# Traveling Waves with Multiple and Nonconvex Fronts for a Bistable Semilinear Parabolic Equation 

MANUEL DEL PINO<br>Departamento de Ingeniería Matemática and Centro de Modelamiento Matemático, Universidad de Chile<br>MICHAŁ KOWALCZYK<br>Departamento de Ingeniería Matemática and Centro de Modelamiento Matemático, Universidad de Chile<br>AND<br>JUNCHENG WEI<br>Chinese University of Hong Kong


#### Abstract

We construct new examples of traveling wave solutions to the bistable and balanced semilinear parabolic equation in $\mathbb{R}^{N+1}, N \geq 2$. Our first example is that of a traveling wave solution with two non planar fronts that move with the same speed. Our second example is a traveling wave solution with a nonconvex moving front. To our knowledge no existence results of traveling fronts with these type of geometric characteristics have been previously known. Our approach explores a connection between solutions of the semilinear parabolic PDE and eternal solutions to the mean curvature flow in $\mathbb{R}^{N+1}$. © 2012 Wiley Periodicals, Inc.


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Communications on Pure and Applied Mathematics, 0001-0067 (PREPRINT)
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## 1 Introduction

The problem of finding traveling wave solutions to autonomous semilinear parabolic PDEs,

$$
\Delta v+f(v)=v_{t} \quad \text { in } \mathbb{R}^{n} \times(-\infty, \infty)
$$

has been studied extensively since the pioneering work of Kolmogorov, Petrovsky, and Piskunow [23] and Fisher [15]. A traveling wave solution propagating in a fixed direction $\mathrm{e} \in \mathbb{R}^{n}$ with speed $c$ is, by definition, a solution of the form $v(x, t)=u(x-c t e)$. When written in the Galilean frame, the traveling wave problem reduces to the following semilinear elliptic PDE:

$$
\begin{equation*}
\Delta u+c \mathrm{e} \cdot \nabla u+f(u)=0 \quad \text { in } \mathbb{R}^{n} . \tag{1.1}
\end{equation*}
$$

The most typical scenario is that of a planar front. Taking the ansatz $u(x)=$ $U(x \cdot \mathrm{e}$ ) reduces (1.1) to the ODE

$$
U^{\prime \prime}+c U^{\prime}+f(U)=0 \quad \text { in } \mathbb{R} .
$$

In this case several examples of existence are well known; most common is the monostable nonlinearity $f(u)=u(1-u)(\mathrm{KPP})$ and the bistable nonlinearity

$$
f(u)=(u+a)\left(1-u^{2}\right), \quad a \in(-1,1) .
$$

In the former case a planar traveling front exists for any $c>2 \sqrt{f^{\prime}(0)}>0$ while in the latter case the nonlinearity determines the speed uniquely,

$$
c=\frac{\int_{-1}^{1} f(t)}{\int_{-1}^{1}\left(U^{\prime}\right)^{2}} .
$$

Note that in the case of balanced bistable nonlinearity we have $c=0$. This means that the traveling wave is a standing wave. These are classical results and we refer the reader to [14] for more information. Other related results in the monostable and bistable cases can be found in [20, 21] (see also [2, 3, 17, 19] and the references therein).

The case of nonplanar fronts is much less understood. Since the subject of this paper is to study the traveling waves with a bistable nonlinearity, we will mention some results in this direction. First let us consider $f$ unbalanced, i.e., $a \neq 0$. When $n=2$ a $V$-shaped traveling wave was found by Ninomiya and Taniguchi and in higher dimension by Hamel, Monneau, and Roquejoffre [18]. Let us comment on the $n=2$ case. Given a traveling wave solution $u(x)$ its traveling front is the nodal set $\{u(x)=0\}$. It can be proven that the front is asymptotic to two straight lines $y=m|x|$, and that it is convex at $\infty$ [25]. Moreover, it is shown that the traveling wave solution is stable. These results are generalized to higher dimensions and fronts of more complex geometric structure, which however share the general characteristics of the V -shaped front; i.e., the front profiles are asymptotically linear and convex, and as solutions of the parabolic problem the traveling wave solutions are stable (see [18, 26, 27]).

Let us now discuss the bistable balanced case. From now on we agree that the direction of propagation will be fixed to coincide with one of the axes. It is convenient to consider the traveling wave problem in $\mathbb{R}^{N+1}$, with the $x_{N+1^{-}}$ axis as the fixed direction of motion and $N$ corresponding to the dimension of the associated traveling front. We will assume that $N \geq 2$. Thus, if we look for solutions to the parabolic Allen-Cahn equation

$$
\begin{equation*}
u_{t}=\Delta u+u-u^{3} \quad \text { in } \mathbb{R}^{N+1} \times(-\infty, \infty), N \geq 2 \tag{1.2}
\end{equation*}
$$

in the following form:

$$
\begin{equation*}
u(x, t)=U\left(x^{\prime}, x_{N+1}-c t\right), \quad x=\left(x^{\prime}, x_{N+1}\right) \tag{1.3}
\end{equation*}
$$

then $U$ will satisfy the traveling wave Allen-Cahn equation

$$
\begin{equation*}
\Delta U+c \partial_{x_{N+1}} U+U-U^{3}=0 \tag{1.4}
\end{equation*}
$$

In [4], the existence of a traveling wave in the form $U\left(r, x_{N+1}\right),\left|x^{\prime}\right|=r$, is obtained for any speed $c>0$. Furthermore, it is shown that asymptotically the 0 -level set of $U$-denoted here by $\Gamma$-is paraboloid-like:

$$
\lim _{\substack{x_{N+1} \rightarrow+\infty \\\left(x^{\prime}, x_{N+1}\right) \in \Gamma}} \frac{r^{2}}{2 x_{N+1}}=\frac{N-1}{c} \quad \text { if } N \geq 2
$$

In the same paper the case $N=1$ is treated as well and the traveling front is shown to be asymptotic to a hyperbolic cosine curve. In all cases traveling fronts are connected, convex surfaces.

The objective of this paper is to show that in the bistable balanced case there exist traveling wave solutions whose traveling fronts are nonconnected, multicomponent surfaces (Theorem 1.1), and also that there are solutions whose fronts are nonconvex (Theorem 1.2). These results are, to our knowledge, the first examples of this type for an autonomous traveling wave problem.

To introduce our results we review some well-known facts about the relation between 1.4 and the so-called translating solutions to the mean curvature flow. These solutions are also called eternal, since they exist for all $t \in(-\infty, \infty)$. In general, we say that an evolving-in-time family of surfaces moves by mean curvature if the following is satisfied:

$$
\begin{equation*}
V=\mathbf{H} \tag{1.5}
\end{equation*}
$$

where $\mathbf{H}$ denotes the mean curvature vector and $V$ the normal velocity of the surface. Translating solutions of this problem are surfaces that do not change shape and are translated by the mean curvature (MC) flow in a fixed direction and with constant velocity. After a rigid motion and rescaling we may assume that a translating solution of the MC flow is represented by a family of surfaces $\left\{\Sigma+c t \mathrm{e}_{N+1}\right\}_{t \in \mathbb{R}}$, where $\Sigma$ is a fixed $N$-dimensional surface in $\mathbb{R}^{N+1}$, and $c \in \mathbb{R}$ is a fixed number. From this we obtain the following equation to determine $\Sigma$ :

$$
\begin{equation*}
H=c v_{N+1}, \tag{1.6}
\end{equation*}
$$

where $H$ is the mean curvature and $v$ is the unit normal vector of the (oriented) surface $\Sigma$ (recall that $\mathbf{H}=H \nu$ ). Observe that the family $\left\{\Sigma+c t e_{N+1}\right\}_{t \in \mathbb{R}}$ is a translating solution of the mean curvature flow, which is translated in the direction parallel to the $x_{N+1}$-axis with the constant speed $c$.

Let us fix a surface $\Sigma$ for which (1.6) holds and such that $c=1$. Let us also define its scaling $\Sigma_{\varepsilon}$ by

$$
\begin{equation*}
y \in \Sigma_{\varepsilon} \Longleftrightarrow \varepsilon y \in \Sigma \tag{1.7}
\end{equation*}
$$

Then, denoting the mean curvatures of these surfaces by $H_{\Sigma}$ and $H_{\Sigma_{\varepsilon}}$, respectively, we see that if $\Sigma$ is a translating solution to the mean curvature flow with speed 1 , then we have

$$
\begin{equation*}
H_{\Sigma_{\varepsilon}}=\varepsilon v_{N+1} \tag{1.8}
\end{equation*}
$$

which means that the scaled surface moves with the constant speed $c=\varepsilon$. In this paper we will consider $\varepsilon$ to be a small parameter, or in other words, we will be interested in translating solutions of the MC flow moving with a small speed.

Several examples of translating solution to the MC equation are known; see, for example, [1, 5, 24, 28]. Here we will discuss a special eternal solution of the mean curvature flow for which $\Sigma$ is a graph of a smooth function $F: \mathbb{R}^{N} \rightarrow \mathbb{R}$, that is, $\Sigma=\left\{\left(x^{\prime}, F\left(x^{\prime}\right)\right), x^{\prime} \in \mathbb{R}^{N}\right\}$. In this case (1.6) reduces to

$$
\begin{equation*}
\nabla\left(\frac{\nabla F}{\sqrt{1+|\nabla F|^{2}}}\right)=\frac{1}{\sqrt{1+|\nabla F|^{2}}} \tag{1.9}
\end{equation*}
$$

It is known from [1, 5] that there exists a unique rotationally symmetric solution $F$ of 1.9 , with the following asymptotic behavior:

$$
\begin{equation*}
F(r)=\frac{r^{2}}{2(N-1)}-\log r+1+O\left(r^{-1}\right), \quad r \gg 1 \tag{1.10}
\end{equation*}
$$

Notice that this asymptotic behavior corresponds (at leading order) to the asymptotic behavior of the nodal set of solutions to (1.4) found in [4]. In what follows we will denote the rotationally symmetric translating solution of the MC flow by $\Gamma$ and the corresponding scaled surface by $\Gamma_{\varepsilon}$. The latter surface is rotationally symmetric, is translating with speed $c=\varepsilon$, and is given as a graph as well:

$$
\Gamma_{\varepsilon}=\left\{x_{N+1}=\varepsilon^{-1} F(\varepsilon r)\right\}
$$

The first result in this paper is about existence of a traveling wave solution to 1.4 whose zero level set consists of two disjoint components, each of which is asymptotically a paraboloid-like surface in a neighborhood of the rotationally symmetric eternal solution to the mean curvature flow $\Gamma_{\varepsilon}$. More precisely, we have the following:

THEOREM 1.1. For each sufficiently small $\varepsilon$, the traveling wave problem (1.4) has a solution $u_{\varepsilon}$ moving with speed $c=\varepsilon$ and with the following properties:
(1) The 0-level set of $u_{\varepsilon}$ consists of two disjoint, rotationally symmetric, and smooth hypersurfaces $\Gamma_{\varepsilon}^{ \pm}$.


Figure 1.1. Schematic view of the surface $\Gamma$ represented as a graph $x_{N+1}=F(r)$ and moving with speed $c=1$, and the surface $\Gamma_{\varepsilon}$, represented as a graph $x_{N+1}=\frac{1}{\varepsilon} F(\varepsilon r)$ and moving with speed $c=\varepsilon$.
(2) The nodal surfaces $\Gamma_{\varepsilon}^{ \pm}$divide the space into three disjoint and unbounded components $\Omega_{\varepsilon}^{ \pm}, \Omega_{\varepsilon}^{0}$. Each of the sets $\Omega_{\varepsilon}^{ \pm}$is a neighborhood, respectively, of $\left(0, x_{N+1}= \pm \infty\right) \in \mathbb{R}^{N+1}$, and it holds $u_{\varepsilon}<0$ in $\Omega_{\varepsilon}^{ \pm}$. The set $\Omega_{\varepsilon}^{0}$ contains $\Gamma_{\varepsilon}$ and $u_{\varepsilon}>0$ in $\Omega_{\varepsilon}^{0}$. Moreover,

$$
\lim _{x_{N+1} \rightarrow \pm \infty} u_{\varepsilon}\left(x^{\prime}, x_{N+1}\right)=-1 \quad \forall x^{\prime} \in \mathbb{R}^{N}
$$

while at the same time

$$
\lim _{\substack{\left(x^{\prime}, x_{N+1)} \rightarrow \infty \\\left(x^{\prime}, x_{N+1}\right) \in \Gamma_{\varepsilon}\right.}} u_{\varepsilon}\left(x^{\prime}, x_{N+1}\right)=1
$$

(3) For any $r>0$ let $C_{r}$ be the cylinder $C_{r}=\left\{\left(x^{\prime}, x_{N+1}\right)| | x^{\prime} \mid=r\right\}$. Let $\Gamma_{\varepsilon}^{ \pm}(r)=\Gamma_{\varepsilon}^{ \pm} \backslash C_{r}$, and similarly $\Gamma_{\varepsilon}(r)=\Gamma_{\varepsilon} \backslash C_{r}$. Then it holds that

$$
\begin{equation*}
\mathrm{d}\left(\Gamma_{\varepsilon}^{ \pm}(r), \Gamma_{\varepsilon}(r)\right)=\mathcal{O}\left(\log \left(\frac{1+\varepsilon^{2} r^{2}}{\varepsilon^{2}}\right)\right) \quad \text { as } r \rightarrow+\infty \tag{1.11}
\end{equation*}
$$

where d is the Hausdorff distance between sets.
Of course, when $u_{\varepsilon}$ is a solution so is $-u_{\varepsilon}$, so our result automatically provides existence of at least two traveling waves with multiple fronts.

Our construction of a traveling wave solution of (1.4 with a two-component traveling front gives more precise information about the moving fronts $\Gamma_{\varepsilon}^{ \pm}$and their relation to $\Gamma_{\varepsilon}$. In particular, it is shown that $\Gamma_{\varepsilon}^{ \pm}$are normal graphs over $\Gamma_{\varepsilon}$ of certain functions $f_{\varepsilon}^{ \pm}: \Gamma_{\varepsilon} \rightarrow \mathbb{R}$, whose asymptotic behavior coincides with the one described in 1.11 above. In Section 2.2 we will discuss this in more detail and we will introduce, based on formal calculations, a system of nonlinear PDEs on $\Gamma_{\varepsilon}$ that determines these functions. A schematic view of the situation is included in Figures 1.1 and 1.2 .

Our second result for the traveling wave problem (1.4) has to do with existence of traveling waves whose fronts are nonconvex surfaces. In fact, in [5] it is proven


Figure 1.2. Illustration of the results of Theorem 1.1 The surfaces $\Gamma_{\varepsilon}^{ \pm}$are presented as well as the asymptotic values of the traveling wave solution $u_{\varepsilon}$.
that in the case of translating solutions of the mean curvature flow in $\mathbb{R}^{N+1}, N \geq$ 2 , there exists a family of rotationally symmetric surfaces $\Sigma_{R}, R>0$, of genus 0 that satisfies

$$
H_{\Sigma_{R}}=v_{R, N+1} .
$$

In other words, $\Sigma_{R}$ is translated by the mean curvature flow in the direction of the $x_{N+1}$-axis with speed $c=1$. Each of these surfaces is formed by taking the union of two graphs of radial functions $W_{R}^{ \pm}:[R, \infty) \rightarrow \mathbb{R}$ in $\mathbb{R}^{N+1}$. These functions satisfy the following asymptotic formulas:

$$
\begin{equation*}
W_{R}^{ \pm}(r)=\frac{r^{2}}{2(N-1)}-\log r+C^{ \pm}+O\left(r^{-1}\right), \quad r \gg 1, \tag{1.12}
\end{equation*}
$$

with some constants $C^{ \pm}$. The graphs of the functions $W_{R}^{ \pm}$are called the ends of $\Sigma_{R}$, and we will refer to them as the upper end $\Sigma_{R}^{+}$and the lower end $\Sigma_{R}^{-}$, respectively. Comparing (1.12) with (1.10) we see that the ends of each of the surfaces $\Sigma_{R}$ are asymptotically "parallel" to the traveling graph $\Gamma$ described above. It is easy to see that $\Sigma_{R}$ divides the space into two disjoint components; we call them $\Omega_{R}^{ \pm}$, respectively, and agree that $\Omega_{R}^{+}$is the component containing the vertical axis $x_{N+1}$ and $\Omega_{R}^{-}$is the other one. Sometimes we refer to the surfaces $\Sigma_{R}$ as traveling catenoids.

We consider a scaling of $\Sigma_{R}$ by a small parameter $\Sigma_{R, \varepsilon}=\frac{1}{\varepsilon} \Sigma_{R}$. The scaled surfaces move now with speed $c=\varepsilon$. We will denote the ends of the scaled traveling catenoid by $\Sigma_{R, \varepsilon}^{ \pm}$. Note that $\Sigma_{R / \varepsilon} \neq \Sigma_{R, \varepsilon}$. Indeed, while both of these surfaces are defined for $r>\frac{R}{\varepsilon}, \Sigma_{R / \varepsilon}$ is a traveling catenoid whose speed is $c=1$, while $\Sigma_{R, \varepsilon}$ is a traveling catenoid whose speed is $c=\varepsilon$. In other words, the surfaces $\Sigma_{R}$ considered for different $R$ are not simple scalings of one another. See Figure 1.3, which illustrates the situation.

We show the following result:


Figure 1.3. Schematic view of the traveling catenoid $\Sigma_{R}$ moving with speed $c=1$ and its rescaled version $\Sigma_{R, \varepsilon}$ moving with speed $c=\varepsilon$. The surface $\Gamma$ is also represented for comparison.

THEOREM 1.2. For each $R>0$ and each $\varepsilon$ sufficiently small there exists a traveling wave solution $u_{\varepsilon}$ of the problem (1.4) moving with speed $c=\varepsilon$, and such that the following hold:
(1) The level set $\tilde{\Sigma}_{R, \varepsilon}=\left\{u_{\varepsilon}=0\right\}$ is a rotationally symmetric, smooth surface of genus 0 .
(2) The surface $\tilde{\Sigma}_{R, \varepsilon}$ divides the space into two disjoint components $D_{R, \varepsilon}^{ \pm}$ such that $u_{\varepsilon}>0$ in $D_{R, \varepsilon}^{+}$and $u_{\varepsilon}<0$ in $D_{R, \varepsilon}^{-}$. Moreover, outside of a sufficiently large ball the set $\Omega_{R, \varepsilon}^{-}$, which is one of the two components into which the traveling catenoid $\Sigma_{R, \varepsilon}$ divides $\mathbb{R}^{N+1}$, is contained in $D_{R, \varepsilon}^{-}$. We also have

$$
\lim _{x_{N+1} \rightarrow \pm \infty} u_{\varepsilon}\left(x^{\prime}, x_{N+1}\right)=1 \quad \forall x^{\prime} \in \mathbb{R}^{N} .
$$

At the same time

$$
\lim _{\substack{\left|\left(x^{\prime}, x_{N+1}\right)\right| \rightarrow \infty \\\left(x^{\prime}, x_{N+1}\right) \in D_{R, \varepsilon}^{\mid} \cap \Omega_{R, \varepsilon}^{-}}} u_{\varepsilon}\left(x^{\prime}, x_{N+1}\right)=-1 .
$$

(3) Let $\widetilde{\Sigma}_{R, \varepsilon}^{ \pm}$denote the ends of the surface $\widetilde{\Sigma}_{R, \varepsilon}$. For each $r>R$ we denote $\widetilde{\Sigma}_{R, \varepsilon}^{ \pm}(r)=\widetilde{\Sigma}_{R, \varepsilon}^{ \pm} \backslash C_{r}$. Correspondingly we introduce the surfaces $\Sigma_{R, \varepsilon}^{ \pm}(r)=\Sigma_{R, \varepsilon}^{ \pm} \backslash C_{r}$. With this notation it holds that

$$
\mathrm{d}\left(\tilde{\Sigma}_{R, \varepsilon}^{ \pm}(r), \Sigma_{R, \varepsilon}^{ \pm}(r)\right)=\mathcal{O}\left(\log \left(\frac{1+\varepsilon^{2} r^{2}}{\varepsilon^{2}}\right)\right)
$$

The existence results in Theorem 1.1 and Theorem 1.2 are rather counterintuitive in view of what happens with the planar fronts. To explain this, let us note that because of statement (2) in Theorem 1.1 the phase labeled -1 has a tendency to invade the other phase. This is because when we take the limit $u_{\varepsilon}\left(x^{\prime}, x_{N+1}\right)$, with $x^{\prime}$


Figure 1.4. Illustration of the results of Theorem 1.2. The surface $\widetilde{\Sigma}_{R, \varepsilon}$ is presented as well as the asymptotic values of the traveling wave solution $u_{\varepsilon}$. For comparison we also include the surface $\Sigma_{\varepsilon, R}$.
fixed and $x_{N+1} \rightarrow \pm \infty$, then $u_{\varepsilon}\left(x^{\prime}, x_{N+1}\right) \rightarrow-1$. In the one-dimensional situation a solution to the parabolic Allen-Cahn equation with initial data satisfying this condition at $\pm \infty$ will eventually converge to -1 . Thus, if this one-dimensional interaction of fronts were the only mechanism present, the nodal hypersurfaces should attract each other and eventually annihilate, and only one phase would remain in the asymptotic limit of infinite time. Based on this a natural statement in higher dimension would then be: if a traveling wave solution of the bistable and balanced problem satisfies $\lim _{x_{N+1} \rightarrow \pm \infty} u\left(x^{\prime}, x_{N+1}\right)=-1$, then $u \equiv-1$.

This turns out to be false because of the mediating effect of the geometry of the front. Indeed, we see that in the situation described by the theorems one stable phase, say -1 , is "surrounded" by the other phase, say +1 , which is also stable thanks to the fact that the nonlinearity is bistable. The nonlinearity being balanced as well, the two phases move with equal speed, and their initial configuration is translated with constant speed and is preserved for all times. As a result we have an eternal solution to the parabolic Allen-Cahn equation. The main effort in this paper is to give a quantitative form of this by deriving and solving a system of PDEs, called the Jacobi-Toda system, for the moving fronts.

Before we close this section, we make several important remarks as well as open questions.

Remark 1.3. The results of Theorems 1.1 and 1.2 hold for general balanced nonlinearity

$$
\begin{equation*}
\Delta U+c \partial_{x_{N+1}} U+f(U)=0 \tag{1.13}
\end{equation*}
$$

where $f(U)=F^{\prime}(U)$ and $F \in \mathcal{C}^{4}(\mathbb{R})$ has two equal wells $\pm 1$ with $F(-1)=$ $F(1)=0$ and $f^{\prime}( \pm 1)<0$. The proofs are similar but the notation and details of some computations become quite cumbersome. For this reason we chose here to work with the cubic, balanced nonlinearity $f(u)=u\left(1-u^{2}\right)$.

On the other hand, it is also possible to construct solutions with multiple traveling fronts when the nonlinearity is unbalanced (see [13]).

Remark 1.4. In the statement of Theorem 1.1] we have assumed that $N \geq 2$. When $N=1$ the traveling wave solution to the mean curvature front is the well-known grim reaper and its properties are quite different. In particular, the ends of the grim reaper become parallel at $\infty$, and as a result to find multiple front traveling waves for the traveling wave problem (1.1) one would have to take into account strong interactions of the ends of the grim reaper. This situation resembles somewhat that of Theorem [1.2, but the problem seems quite technical and is beyond the scope of this paper. In this context it is worth mentioning that according to a result of Gui [16] traveling wave solutions (1.1) with one front must be even symmetric. An interesting question would then be whether multiple front traveling wave solutions with no even symmetry exist.
Remark 1.5. In the proof of Theorem 1.1, for brevity we have only dealt with the case of $k=2$ front traveling wave. The techniques can be extended to a multiple front traveling wave ( $k>2$ ), but the technical details render the proofs a bit longer. The main issue, which is the solution of the Jacobi-Toda system, can be handled similarly as in [12].

This paper is organized as follows. First, we formally explain the result in Theorem 1.1 introducing in the process the Jacobi-Toda system for a traveling solution to the mean curvature flow. Next, we solve this system for $\Gamma_{\varepsilon}$. This is in fact the core of our paper. Then we use the infinite-dimensional Lyapunov-Schmidt reduction procedure to show the existence for (1.1). Finally, we prove Theorem 1.2

## 2 Jacobi-Toda System and Multicomponent Traveling Fronts

The discussion in this section is mostly formal; however, we think that it is useful in order to understand the role played by the Jacobi-Toda system in the existence of a traveling wave with multiple components. We choose to work in the setting that is more general than the one of Theorem 1.1 to emphasize the universality of this system. The notation and many calculations presented here will be used throughout the paper.

### 2.1 Geometric Background

Let us consider a parametrized, regular, $N$-dimensional surface $\Sigma(t)$ for which (1.6) is satisfied. We will consider its parametrization over a family of open sets $\mathcal{U}_{\alpha} \subset \mathbb{R}^{N}, \alpha \in \mathcal{A}$, and associated smooth maps $q_{\alpha}: \mathcal{U}_{\alpha} \rightarrow \mathbb{R}^{N+1}$ such that their images cover $\Sigma$, namely $\bigcup_{\alpha \in \mathcal{A}} q_{\alpha}\left(\mathcal{U}_{\alpha}\right)=\Sigma$. Furthermore, we fix an orientation on $\Sigma$, and by $v$ we will denote the vector field of the unit normal vectors. Let us consider a tubular neighborhood $D_{\delta}$ of $\Sigma$ given by

$$
D_{\delta}=\{|\operatorname{dist}(\Sigma, x)|<\delta\} \subset \mathbb{R}^{N+1}
$$

where dist denotes the signed distance. All our calculations below have local character, and for this reason we will fix a pair $\left(q_{\alpha}, \mathcal{U}_{\alpha}\right)$ and, for simplicity of notation,
drop the subscript $\alpha$. For each sufficiently small $\delta$ the map

$$
(s, z) \longmapsto X \in D_{\delta} \quad \text { where } X(s, z)=q(s)+z v(s), s=\left(s_{1}, \ldots, s_{N}\right) \in \mathcal{U}
$$

is a diffeomorphism onto $D_{\delta}$ (we will consistently abuse the notation, writing $v(s)$ instead of $v(q(s)))$.

In what follows we will work with the scaled version of $\Sigma$, namely $\Sigma_{\varepsilon}$, and we will denote its parametrization and the unit normal by $q_{\varepsilon}$ and $\nu_{\varepsilon}$, respectively. It is easy to see that the following relations hold:

$$
q_{\varepsilon}(s)=\varepsilon^{-1} q(\varepsilon s), \quad v_{\varepsilon}(s)=v(\varepsilon s), \quad s \in \varepsilon^{-1} \mathcal{U}
$$

and that similar scaling formulas can be derived for other functions defined on $\Sigma_{\varepsilon}$. We also have local coordinates in $D_{\delta / \varepsilon}$, which we will still denote by $(s, z)$ and the map $X_{\varepsilon}$ defined by

$$
X_{\varepsilon}(s, z)=q_{\varepsilon}(s)+z v_{\varepsilon}(s)
$$

It is convenient to introduce the following notation for functions $f: D_{\delta / \varepsilon} \rightarrow \mathbb{R}$ :

$$
\left(X_{\varepsilon}^{*} f\right)(s, z)=\left(f \circ X_{\varepsilon}\right)(s, z)
$$

The function $X_{\varepsilon}^{*} f: X_{\varepsilon}^{-1}\left(D_{\delta / \varepsilon}\right) \rightarrow \mathbb{R}$ can be interpreted as the pullback of $f$ via parametrization $X_{\varepsilon}$. In a similar way, we define the pullback of a map $f: D_{\delta / \varepsilon} \rightarrow$ $\mathbb{R}^{d}, d \geq 1$, via $X_{\varepsilon}$. By $\left(X^{*} f\right)$ we denote the pullback of $f: D_{\delta} \rightarrow \mathbb{R}^{d}$ via $X$.

We will now derive formulas expressing $\Delta$ and $\partial_{x_{N+1}}$ in $D_{\delta / \varepsilon}$ in terms of $(s, z) \in \varepsilon \mathcal{U} \times(-\delta / \varepsilon, \delta / \varepsilon)$. We define for each $z \in(-\delta / \varepsilon, \delta / \varepsilon)$

$$
\Sigma_{\varepsilon, z}=\left\{x \in D_{\delta / \varepsilon} \mid \operatorname{dist}\left(\Sigma_{\varepsilon}, x\right)=z\right\}
$$

In other words $\Sigma_{\varepsilon, z}$ is the surface obtained from $\Sigma_{\varepsilon}$ by translation in the direction of the normal by $z$. Then the well-known formula gives

$$
\begin{equation*}
\Delta=\Delta_{\Sigma_{\varepsilon, z}}+\partial_{z}^{2}-H_{\Sigma_{\varepsilon, z}} \partial_{z} \tag{2.1}
\end{equation*}
$$

where $H_{\Sigma_{\varepsilon, z}}$ denotes the mean curvature of $\Sigma_{\varepsilon, z}$. We need to expand these operators in terms of the variable $z$. By $g_{\Sigma_{\varepsilon}}$ and $g_{\Sigma_{\varepsilon, z}}$, respectively, we will denote the metric on $\Sigma_{\varepsilon}$ and $\Sigma_{\varepsilon, z}$ (induced from $\mathbb{R}^{N+1}$ ). In terms of $s \in \varepsilon^{-1} \mathcal{U}$ we get the following expressions:

$$
\begin{equation*}
g_{\Sigma_{\varepsilon, z,}, i j}=g_{\Sigma_{\varepsilon}, i j}+\varepsilon z a_{\varepsilon, i j}+\varepsilon^{2} z^{2} b_{\varepsilon, i j} \tag{2.2}
\end{equation*}
$$

where

$$
\begin{gather*}
g_{\Sigma_{\varepsilon}, i j}=\left(\partial_{j} q \cdot \partial_{i} q\right)(\varepsilon s), \quad a_{\varepsilon, i j}(s)=\left(\partial_{j} q \cdot \partial_{i} v\right)(\varepsilon s)+\left(\partial_{i} q \cdot \partial_{j} v\right)(\varepsilon s)  \tag{2.3}\\
b_{\varepsilon, i j}(s)=\left(\partial_{i} v \cdot \partial_{j} v\right)(\varepsilon s)
\end{gather*}
$$

Then, for the matrix $g_{\Sigma_{\varepsilon, z}}^{-1}=\left(g_{\Sigma_{\varepsilon, z}}^{i j}\right)_{i, j=1, \ldots, N}$, we get, provided that $|\varepsilon z|$ is sufficiently small,

$$
\begin{equation*}
g_{\Sigma_{\varepsilon, z}}^{-1}=g_{\Sigma_{\varepsilon}}^{-1}+\varepsilon z A_{\varepsilon}+\varepsilon^{2} z^{2} B_{\varepsilon} \tag{2.4}
\end{equation*}
$$

where

$$
A_{\varepsilon}=A(\varepsilon s), \quad B_{\varepsilon}=B(\varepsilon s, \varepsilon z)
$$

and $A: \mathcal{U} \rightarrow \mathbb{R}^{N} \times \mathbb{R}^{N}, B: \mathcal{U} \times(-\delta, \delta) \rightarrow \mathbb{R}^{N} \times \mathbb{R}^{N}$ are smooth matrix functions. The expression for the Laplace-Beltrami operator on $\Sigma_{\varepsilon}$ in local coordinates is

$$
\begin{aligned}
\Delta_{\Sigma_{\varepsilon}} & =\frac{1}{\sqrt{\operatorname{det}\left(g_{\left.\Sigma_{\varepsilon}\right)}\right.}} \partial_{j}\left(\sqrt{\operatorname{det}\left(g_{\Sigma_{\varepsilon}}\right)} g_{\Sigma_{\varepsilon}}^{i j} \partial_{i}\right) \\
& =g_{\Sigma_{\varepsilon}}^{i j} \partial_{i j}+\frac{1}{\sqrt{\operatorname{det}\left(g_{\Sigma_{\varepsilon}}\right)}} \partial_{j}\left(\sqrt{\operatorname{det}\left(g \Sigma_{\varepsilon}\right)} g_{\Sigma_{\varepsilon}}^{i j}\right) \partial_{i} \\
& =g_{\Sigma_{\varepsilon}}^{i j} \partial_{i j}-g^{k \ell} \Gamma_{\Sigma_{\varepsilon}, k \ell}^{i} \partial_{i},
\end{aligned}
$$

where $\Gamma_{\Sigma_{\varepsilon}, k \ell}^{i}$ are the Christoffel symbols. A similar formula holds for $\Delta_{\Sigma_{\varepsilon, z}}$. Using this we can write

$$
\Delta_{\Sigma_{\varepsilon, z}}=\Delta_{\Sigma_{\varepsilon}}+\mathbb{A}_{\varepsilon, i j} \partial_{i j}+\mathbb{B}_{\varepsilon, i} \partial_{i},
$$

where

$$
\begin{aligned}
\mathbb{A}_{\varepsilon, i j} & =g_{\Sigma_{\varepsilon, z}}^{i j}-g_{\Sigma_{\varepsilon}}^{i j}, \\
\mathbb{B}_{\varepsilon, i} & =g_{\Sigma_{\varepsilon, z}}^{k \ell}\left[\Gamma_{\Sigma_{\varepsilon, z}, k \ell}^{i}-\Gamma_{\Sigma_{\varepsilon}, k \ell}^{i}\right]+\Gamma_{\Sigma_{\varepsilon}, k \ell}^{i}\left[g_{\Sigma_{\varepsilon, z}}^{k \ell}-g_{\Sigma_{\varepsilon}}^{k \ell}\right] .
\end{aligned}
$$

Expressions in local coordinates for $\mathbb{A}_{\varepsilon, i j}$ and $\mathbb{B}_{\varepsilon, i}$ can be further derived using the above expansions; however, their exact form is not crucial here. The point is that, formally, these functions are small in terms of $|\varepsilon z|$. Finally, for future reference, we notice that if $f_{\varepsilon} \in \mathcal{C}^{2}\left(\Sigma_{\varepsilon}\right)$ is identified with $f \in \mathcal{C}^{2}(\Sigma)$ through the formula $\left(X_{\varepsilon}^{*} f_{\varepsilon}\right)(s)=\left(X^{*} f\right)(\varepsilon s)$, then

$$
\begin{equation*}
\left(X_{\varepsilon}^{*} \Delta_{\Sigma_{\varepsilon}} f_{\varepsilon}\right)(s)=\varepsilon^{2}\left(X^{*} \Delta_{\Sigma} f\right)(\varepsilon s) . \tag{2.5}
\end{equation*}
$$

Next, we will expand the mean curvature $H_{\Sigma_{\varepsilon, z}}$. To this end we will denote by $\mathbb{k}_{\varepsilon, j}, j=1, \ldots, N$, the principal curvatures of $\Sigma_{\varepsilon}$. Then we have

$$
\begin{align*}
H_{\Sigma_{\varepsilon, z}}=\sum_{j=1}^{N} \frac{\mathbb{k}_{\varepsilon, j}}{1-z \mathbb{k}_{\varepsilon, j}} & =\sum_{j=1}^{N} \mathbb{k}_{\varepsilon, j}+z \sum_{j=1}^{N} \mathbb{k}_{\varepsilon, j}^{2}+z^{2} R_{\Sigma_{\varepsilon}}  \tag{2.6}\\
& =H_{\Sigma_{\varepsilon}}+z\left|A_{\Sigma_{\varepsilon}}\right|^{2}+z^{2} R_{\Sigma_{\varepsilon}},
\end{align*}
$$

where

$$
R_{\Sigma_{\varepsilon}}=\sum_{j=1}^{N} \mathbb{k}_{\varepsilon, j}^{3}+z \sum_{j=1}^{N} \mathbb{k}_{\varepsilon, j}^{4}+\cdots,
$$

and $\left|A_{\Sigma_{\varepsilon}}\right|$ is the norm of the second fundamental form on $\Sigma_{\varepsilon}$. Denoting by $\mathbb{k}_{j}$, $j=1, \ldots, N$, the principal curvatures of $\Sigma$, it is straightforward to see that $\left(X_{\varepsilon}^{*} \mathbb{k}_{\varepsilon, j}\right)(s)=\varepsilon\left(X^{*} \mathbb{k}_{j}\right)(\varepsilon s)$, hence

$$
\begin{equation*}
\left(X_{\varepsilon}^{*}\left|A_{\Sigma_{\varepsilon}}\right|^{2}\right)(s)=\varepsilon^{2}\left(X^{*}\left|A_{\Sigma}\right|^{2}\right)(\varepsilon s) . \tag{2.7}
\end{equation*}
$$

To compute the expression for $\partial_{x_{N+1}} \equiv \partial_{N+1}$ in local coordinates of $D_{\delta / \varepsilon}$, we observe that for any function $f_{\varepsilon}$ in $D_{\delta / \varepsilon}$ we have

$$
\partial_{N+1} f_{\varepsilon}=\nabla f_{\varepsilon} \cdot \nabla\left(\pi_{\varepsilon, N+1}\right)
$$

where $\pi_{\varepsilon, j}: D_{\delta / \varepsilon} \rightarrow \mathbb{R}$ denotes the projection on the $j^{\text {th }}$ coordinate. Furthermore, we have the following formula for the gradient (interpreted as a vector field defined on $D_{\delta / \varepsilon}$ ):

$$
\begin{equation*}
\nabla f_{\varepsilon}=\nabla_{\Sigma_{\varepsilon, z}} f_{\varepsilon}+\partial_{z} f_{\varepsilon} \partial_{z} \tag{2.8}
\end{equation*}
$$

where $\nabla_{\Sigma_{\varepsilon, z}}$ denotes the gradient vector field on the hypersurface $\Sigma_{\varepsilon, z}$.
The formula for the gradient in the local coordinates $(s, z) \in \varepsilon^{-1} \mathcal{U} \times(-\delta / \varepsilon, \delta / \varepsilon)$ is given by

$$
\left(X_{\varepsilon}^{*} \nabla f_{\varepsilon}\right)=\partial_{j}\left(X_{\varepsilon}^{*} f_{\varepsilon}\right) g_{\Sigma_{\varepsilon, z}}^{i j} \partial_{i}+\partial_{z}\left(X_{\varepsilon}^{*} f_{\varepsilon}\right) \partial_{z}
$$

hence

$$
\begin{aligned}
X_{\varepsilon}^{*}\left(\partial_{N+1} f_{\varepsilon}\right)= & \left(X_{\varepsilon}^{*} \nabla_{\Sigma_{\varepsilon, z}} f_{\varepsilon}\right) \cdot\left(X_{\varepsilon}^{*} \nabla_{\Sigma_{\varepsilon, z}} \pi_{\varepsilon, N+1}\right) \\
& +X_{\varepsilon}^{*}\left(\partial_{\nu_{\varepsilon}} f_{\varepsilon}\right) X_{\varepsilon}^{*}\left(\partial_{\nu_{\varepsilon}} \pi_{\varepsilon, N+1}\right) \\
= & g_{\Sigma_{\varepsilon, z}}^{j_{j}} \partial_{j}\left(X_{\varepsilon}^{*} f\right) \partial_{i}\left(X_{\varepsilon}^{*} \pi_{\varepsilon, N+1}\right)+\partial_{z}\left(X_{\varepsilon}^{*} f_{\varepsilon}\right) \partial_{z}\left(X_{\varepsilon}^{*} \pi_{\varepsilon, N+1}\right) .
\end{aligned}
$$

Observe that $X_{\varepsilon}^{*} \pi_{\varepsilon, N+1}=q_{\varepsilon, N+1}+z v_{\varepsilon, N+1}$; hence, using (2.4) and neglecting those terms that carry a factor of $\varepsilon z$ in front, we get the following asymptotic formula, valid whenever $|\varepsilon z|$ is small:

$$
\begin{equation*}
X_{\varepsilon}^{*}\left(\partial_{N+1} f_{\varepsilon}\right) \approx g_{\Sigma_{\varepsilon}}^{i j} \partial_{j}\left(X_{\varepsilon}^{*} f\right) \partial_{i}\left(q_{\varepsilon, N+1}\right)+\partial_{z}\left(X_{\varepsilon}^{*} f_{\varepsilon}\right) \nu_{\varepsilon, N+1} \tag{2.9}
\end{equation*}
$$

Here and below we denote $f \approx g$ when $f-g$ is a higher-order term.
To find the scaling formula for this expression, we observe that if $f_{\varepsilon} \in \mathcal{C}^{2}\left(D_{\delta / \varepsilon}\right)$ and $f \in \mathcal{C}^{2}\left(D_{\delta}\right)$ are related through the formula $\left(X_{\varepsilon}^{*} f_{\varepsilon}\right)(s, z)=\left(X^{*} f\right)(\varepsilon s, \varepsilon z)$, then

$$
X_{\varepsilon}^{*}\left(\nabla_{\Sigma_{\varepsilon}} f_{\varepsilon}\right)=\varepsilon X^{*}\left(\nabla_{\Sigma} f\right)
$$

and in particular, since we have

$$
\left(X_{\varepsilon}^{*} \pi_{\varepsilon, N+1}\right)(s, z)=\varepsilon^{-1}\left(X^{*} \pi_{N+1}\right)(\varepsilon s, \varepsilon z), \quad v_{\varepsilon, N+1}(s)=v_{N+1}(\varepsilon s),
$$

therefore

$$
\begin{align*}
X_{\varepsilon}^{*} & \left(\partial_{N+1} f_{\varepsilon}\right)(s, z) \\
& \approx \varepsilon X^{*}\left(\partial_{N+1} f\right)(\varepsilon s, \varepsilon z) \\
& =\varepsilon\left[\left(X^{*}\left(\nabla_{\Sigma} f \cdot \nabla_{\Sigma} \pi_{N+1}\right)\right)+\left(X^{*}\left(\partial_{\nu_{N+1}} f\right)\left(X^{*} \partial_{\nu_{N+1}} \pi_{N+1}\right)\right)\right](\varepsilon s, \varepsilon z)  \tag{2.10}\\
& =\varepsilon\left[g_{\Sigma}^{i j}\left(\partial_{j} X^{*} f\right)\left(\partial_{i} q_{N+1}\right)+\partial_{z}\left(X^{*} f\right) \nu_{N+1}\right](\varepsilon s, \varepsilon z) .
\end{align*}
$$

### 2.2 A Model for Multicomponent Traveling Waves

In this section we will describe an approximate form of the multiple traveling wave solution to equation (1.4), where $c=\varepsilon$ is considered to be a small parameter. This approximate solution models the multiple traveling waves in the sense that the true solution to (1.4) with $c=\varepsilon$ is its small perturbation as $\varepsilon \rightarrow 0$. In general, it is reasonable to assume that each component of the multiple traveling wave is a normal graph over an eternal, translating solution of the MC flow, represented by the hypersurface $\Sigma_{\varepsilon}$. Moreover, the profile of each component of the traveling
front should locally resemble a one-dimensional solution of (1.4) with $\varepsilon=0$. Given these observations we will proceed now with more precise definitions.

Let $H$ be the unique odd and monotonically increasing heteroclinic solution of (1.4) in one dimension:

$$
H^{\prime \prime}+H\left(1-H^{2}\right)=0 \quad \text { in } \mathbb{R} .
$$

For future reference let us recall that $H(t)=\tanh (t / \sqrt{2})$.
Furthermore, let $f_{j}: \Sigma \rightarrow \mathbb{R}, j=1, \ldots, k, k>1$, be smooth functions such that $f_{j}<f_{j+1}$. We also set for convenience $f_{0}=-\infty$ and $f_{k+1}=\infty$. In our formal considerations we do not restrict $k$, however, to keep the paper at a reasonable length, the rigorous construction is carried on for $k=2$ only (see Remark 1.5).

We now define the approximate solution $u_{\varepsilon}$ through its expression in local coordinates $(q, \mathcal{U})$ by

$$
\begin{align*}
\left(X_{\varepsilon}^{*} u_{\varepsilon}\right)(s, z)= & \sum_{j=1}^{k}(-1)^{j+1} H\left(z-\left(X^{*} f_{j}\right)(\varepsilon s)\right)+\frac{1}{2}\left(1-(-1)^{k+1}\right),  \tag{2.11}\\
& \text { where } s \in \varepsilon^{-1} \mathcal{U}, z \in(-\delta / \varepsilon, \delta / \varepsilon) .
\end{align*}
$$

Later on we will have to be more specific about the way the approximate solution is defined outside of $D_{\delta / \varepsilon}$ (which is in fact a nontrivial matter), but for our formal considerations it suffices to know $u_{\varepsilon}$ in $D_{\delta / \varepsilon}$. In what follows we will denote $f_{j}(\varepsilon s)=f_{\varepsilon, j}(s)$, so that $f_{\varepsilon, j}: \Sigma_{\varepsilon} \rightarrow \mathbb{R}$ and that the following relation holds: $\left(X_{\varepsilon}^{*} f_{\varepsilon, j}\right)(s)=\left(X^{*} f_{j}\right)(\varepsilon s), s \in \varepsilon^{-1} \mathcal{U}$.

In order to solve (1.4) we will further introduce a new unknown function $\phi$ and look for a solution in the form $u=u_{\varepsilon}+\phi$. Substituting into (1.4) with $c=\varepsilon$ we get

$$
\Delta u+\varepsilon \partial_{x_{N+1}} u+f(u)=S\left(u_{\varepsilon}\right)+\mathbb{L}(\phi)+N(\phi), \quad f(u)=u\left(1-u^{2}\right),
$$

where

$$
\begin{aligned}
S\left(u_{\varepsilon}\right) & =\Delta u_{\varepsilon}+\varepsilon \partial_{x_{N+1}} u_{\varepsilon}+f\left(u_{\varepsilon}\right), \\
\mathbb{L}(\phi) & =\Delta \phi+\varepsilon \partial_{x_{N+1}} \phi+f^{\prime}\left(u_{\varepsilon}\right) \phi, \\
N(\phi) & =f\left(u_{\varepsilon}+\phi\right)-f\left(u_{\varepsilon}\right)-f^{\prime}\left(u_{\varepsilon}\right) \phi .
\end{aligned}
$$

Then, roughly speaking, (1.4) is reduced to finding $\phi$ and $f_{\varepsilon, j}, j=1, \ldots, k$, such that

$$
\begin{equation*}
\mathbb{L}(\phi)+S\left(u_{\varepsilon}\right)+N(\phi)=0 . \tag{2.12}
\end{equation*}
$$

As we will see later on, this problem requires further modification and in particular to solve it we will analyze in detail invertibility properties of the linear operator $\mathbb{L}$. Let us notice one important fact in this context. If by $H_{\varepsilon, j}^{\prime}$ we denote

$$
\left(X_{\varepsilon}^{*} H_{\varepsilon, j}^{\prime}\right)(s, z)=H^{\prime}\left(z-\left(X^{*} f_{\varepsilon, j}\right)(s)\right),
$$

then

$$
\mathbb{L}\left(H_{\varepsilon, j}^{\prime}\right)=o(1), \quad \varepsilon \rightarrow 0
$$

Thus the inverse of the linear operator $\mathbb{L}$ is not expected to be uniformly bounded as $\varepsilon \rightarrow 0$, since the function $H_{\varepsilon, j}^{\prime}$ is in the approximate kernel of $\mathbb{L}$. On the other hand, to solve 2.12 for $\phi$ we would like to use a fixed point argument for the operator

$$
\phi \longmapsto-\mathbb{L}^{-1}\left(S\left(u_{\varepsilon}\right)+N(\phi)\right),
$$

and this clearly requires that $\left\|\mathbb{L}^{-1}\right\|$ be bounded independently on $\varepsilon$. A standard way to deal with this difficulty is to employ the method of infinite-dimensional Lyapunov-Schmidt reduction. The idea is simple: for any function $\psi: D_{\delta / \varepsilon} \rightarrow \mathbb{R}$ we define a projection operator $\Pi_{\varepsilon}$ by

$$
\left(X_{\varepsilon}^{*} \Pi_{\varepsilon} \psi\right)=\left(X_{\varepsilon}^{*} H_{\varepsilon, j}^{\prime}\right)(s, z) \frac{\int_{-\delta / \varepsilon}^{\delta / \varepsilon}\left[\left(X_{\varepsilon}^{*} \psi\right)\left(X_{\varepsilon}^{*} H_{\varepsilon, j}^{\prime}\right)\right](s, z) d z}{\int_{-\delta / \varepsilon}^{\delta / \varepsilon}\left(X_{\varepsilon}^{*} H_{\varepsilon, j}^{\prime}\right)^{2}(s, z) d z}
$$

Next we decompose $\phi=\phi^{\|}+\phi^{\perp}$ where

$$
\left(X_{\varepsilon}^{*} \phi^{\|}\right)=\left(X_{\varepsilon}^{*} \Pi_{\varepsilon} \phi\right)
$$

Then problem (2.12) reduces to

$$
\begin{align*}
\Pi_{\varepsilon}\left[\mathbb{L}(\phi)+S\left(u_{\varepsilon}\right)+N(\phi)\right] & =0  \tag{2.13}\\
\left(\mathrm{Id}-\Pi_{\varepsilon}\right)\left[\mathbb{L}(\phi)+S\left(u_{\varepsilon}\right)+N(\phi)\right] & =0 \tag{2.14}
\end{align*}
$$

Neglecting formally terms involving $N(\phi)$ and $\mathbb{L}(\phi)$ in (2.13), which should be of higher order, this condition reads

$$
\begin{equation*}
\int_{-\delta / \varepsilon}^{\delta / \varepsilon}\left(X_{\varepsilon}^{*}\left[S\left(u_{\varepsilon}\right) H_{\varepsilon, j}^{\prime}\right]\right)(s, z) d z=0, \quad j=1, \ldots, k, \forall s \in \varepsilon^{-1} \mathcal{U} \tag{2.15}
\end{equation*}
$$

Recall here that we work with a fixed pair $(q, \mathcal{U})$ belonging to the parametrization $\left(q_{\alpha}, \mathcal{U}_{\alpha}\right)_{\alpha \in \mathcal{A}}$ of $\Sigma$, but of course this condition needs to be satisfied for all $\mathcal{U}_{\alpha}$.

We will now use equations 2.15 to derive the Jacobi-Toda system on $\Sigma_{\varepsilon}$. We will write for each fixed $j$ :

$$
\begin{align*}
& \int_{-\delta / \varepsilon}^{\delta / \varepsilon}\left(X_{\varepsilon}^{*}\left(S\left(u_{\varepsilon}\right) H_{\varepsilon, j}^{\prime}\right)\right)(s, z) d z \\
& \left.\quad \approx \int_{-\delta / \varepsilon}^{\delta / \varepsilon} X_{\varepsilon}^{*}\left(\left(\Delta H_{\varepsilon, j}+\varepsilon \partial_{x_{N+1}} H_{\varepsilon, j}+f\left(H_{\varepsilon, j}\right)\right) H_{\varepsilon, j}^{\prime}\right)\right)(s, z) d z  \tag{2.16}\\
& \quad+\int_{-\delta / \varepsilon}^{\delta / \varepsilon} X_{\varepsilon}^{*}\left(\left[f\left(\sum_{i=0}^{2} H_{\varepsilon, j+i-1}\right)-\sum_{i=0}^{2} f\left(H_{\varepsilon, j+i-1}\right)\right] H_{\varepsilon, j}^{\prime}\right)(s, z) d z
\end{align*}
$$

Observe that above we took into account only terms representing the interactions of the $j^{\text {th }}$ wave with its "immediate" neighbors. The remaining terms represent interactions of the $j^{\text {th }}$ wave with those waves whose distances to the $j^{\text {th }}$ wave are large enough to render their interactions negligible.

Now we will consider the integrand of the first of the integrals on the right-hand side of (2.16). Using the expressions for $\Delta$ and $\partial_{x_{N+1}}$, and neglecting small terms (as in the previous section), we get

$$
\begin{aligned}
X_{\varepsilon}^{*} S\left(H_{\varepsilon, j}\right) \approx & \partial_{z z} X_{\varepsilon}^{*} H_{\varepsilon, j}+X_{\varepsilon}^{*} f\left(H_{\varepsilon, j}\right)+X_{\varepsilon}^{*}\left(\varepsilon \nu_{N+1}-H_{\Sigma_{\varepsilon}}\right) \partial_{z} X_{\varepsilon}^{*} H_{\varepsilon, j} \\
& +X_{\varepsilon}^{*}\left[\left(\Delta_{\Sigma_{\varepsilon}}-z\left|A_{\Sigma_{\varepsilon}}\right|^{2} \partial_{z}\right) H_{\varepsilon, j}+\varepsilon \nabla_{\Sigma_{\varepsilon}} H_{\varepsilon, j} \cdot \nabla_{\Sigma_{\varepsilon}}\left(\pi_{\varepsilon, N+1}\right)\right] .
\end{aligned}
$$

Consecutive terms above are organized in such a way that the first term is simply 0 by definition of $H_{\varepsilon, j}$, the second term is also 0 since $\Sigma_{\varepsilon}$ is an eternal solution of the mean curvature flow translating with speed $c=\varepsilon$, and the third is of order $\mathcal{O}\left(\varepsilon^{2}\right)$. In this term we will separate those parts whose projection $\Pi_{\varepsilon}$ onto $H_{\varepsilon, j}^{\prime}$ is nonzero from the rest:

$$
\begin{align*}
& X_{\varepsilon}^{*} S\left(H_{\varepsilon, j}\right) \\
& \quad \approx X_{\varepsilon}^{*}\left[\left(-\Delta_{\Sigma_{\varepsilon}} f_{\varepsilon, j}-\left|A_{\Sigma_{\varepsilon}}\right|^{2} f_{\varepsilon, j}-\varepsilon \nabla_{\Sigma_{\varepsilon}} f_{\varepsilon, j} \cdot \nabla_{\Sigma_{\varepsilon}}\left(\pi_{\varepsilon, N+1}\right)\right) H_{\varepsilon, j}^{\prime}\right]  \tag{2.17}\\
& \quad+X_{\varepsilon}^{*}\left(\left|\nabla_{\Sigma_{\varepsilon}} f_{\varepsilon, j}\right|^{2} H_{\varepsilon, j}^{\prime \prime}\right)-\left(z-X_{\varepsilon}^{*} f_{\varepsilon, j}\right) X_{\varepsilon}^{*}\left(\left|A_{\Sigma_{\varepsilon}}\right|^{2} H_{\varepsilon, j}^{\prime}\right) .
\end{align*}
$$

If we take this formula into account, it is not hard to show that

$$
\begin{align*}
& \int_{-\delta / \varepsilon}^{\delta / \varepsilon} X_{\varepsilon}^{*} S\left(H_{\varepsilon, j}\right) H_{\varepsilon, j}^{\prime} \\
& \left.\quad \approx-c_{0} X_{\varepsilon}^{*}\left(\Delta_{\Sigma_{\varepsilon}} f_{\varepsilon, j}+\left|A_{\Sigma_{\varepsilon}}\right|^{2} f_{\varepsilon, j}+\varepsilon \nabla_{\Sigma_{\varepsilon}} f_{\varepsilon, j} \cdot \nabla_{\Sigma_{\varepsilon}}\left(\pi_{\varepsilon, N+1}\right)\right)(s)\right)  \tag{2.18}\\
& \quad=-\varepsilon^{2} c_{0} X^{*}\left(\Delta_{\Sigma} f_{j}+\left|A_{\Sigma}\right|^{2} f_{j}+\nabla_{\Sigma} f_{j} \cdot \nabla_{\Sigma}\left(\pi_{N+1}\right)\right)(\varepsilon s),
\end{align*}
$$

where $c_{0}=\int_{\mathbb{R}}\left(H^{\prime}\right)^{2}$.
Similarly, we will identify in the integrand in the second integral in (2.16) those parts whose projection onto $H_{\varepsilon, j}^{\prime}$ is nontrivial. Here we use the fact that from $H(t)=\tanh (t / \sqrt{2})$ we get $1-H^{2}=\sqrt{2} H^{\prime}$. After some elementary manipulations we find

$$
\begin{align*}
f\left(\sum_{i=0}^{2} H_{\varepsilon, j+i-1}\right)- & \sum_{i=0}^{2} f\left(H_{\varepsilon, j+i-1}\right) \approx  \tag{2.19}\\
& 3 \sqrt{2} H_{\varepsilon, j}^{\prime}\left(H_{\varepsilon, j-1}-1\right)+3 \sqrt{2} H_{\varepsilon, j}^{\prime}\left(H_{\varepsilon, j+1}+1\right),
\end{align*}
$$

where the terms that we have neglected turn out to have small contributions when projected onto $H_{\varepsilon, j}^{\prime}$. To compute the projection $\Pi_{\varepsilon}$, let us recall that

$$
H(t)-1 \approx-2 e^{-\sqrt{2} t}, \quad t \rightarrow \infty, \quad H(t)+1 \approx 2 e^{\sqrt{2} t}, \quad t \rightarrow-\infty
$$

Then we obtain the following as the leading-order term in the second integral in (2.16):

$$
\begin{align*}
& 3 \sqrt{2} \int_{-\delta / \varepsilon}^{\delta / \varepsilon} X_{\varepsilon}^{*}\left[\left(H_{\varepsilon, j}^{\prime}\right)^{2}\left(H_{\varepsilon, j-1}-1\right)\right](s, z) d z \\
& +3 \sqrt{2} \int_{-\delta / \varepsilon}^{\delta / \varepsilon} X_{\varepsilon}^{*}\left[H_{\varepsilon, j}^{\prime}\left(H_{\varepsilon, j+1}+1\right)\right](s, z) d z  \tag{2.20}\\
& \quad \approx 6 \sqrt{2} c_{1} X_{\varepsilon}^{*}\left(e^{\sqrt{2}\left(f_{\varepsilon, j-1}-f_{\varepsilon, j}\right)}-e^{\sqrt{2}\left(f_{\varepsilon, j}-f_{\varepsilon, j+1}\right)}\right)(s) \\
& \quad=6 \sqrt{2} c_{1} X^{*}\left(e^{\sqrt{2}\left(f_{j-1}-f_{\varepsilon, j}\right)}-e^{\sqrt{2}\left(f_{j}-f_{j+1}\right)}\right)(\varepsilon s)
\end{align*}
$$

where

$$
c_{1}=\int_{-\infty}^{\infty}\left(H^{\prime}(t)\right)^{2} e^{\sqrt{2} t} d t
$$

Denoting

$$
\alpha_{0}=\frac{c_{0}}{6 \sqrt{2} c_{1}}=\frac{1}{6 \sqrt{2}} \frac{\int_{\mathbb{R}}\left(H^{\prime}\right)^{2}}{\int_{\mathbb{R}}\left(H^{\prime}(t)\right)^{2} e^{\sqrt{2} t}}=\frac{\sqrt{2}}{24},
$$

we find that to leading order (2.15) is equivalent to

$$
\begin{align*}
\alpha_{0}\left(\Delta_{\Sigma_{\varepsilon}} f_{\varepsilon, j}+\left|A_{\Sigma_{\varepsilon}}\right|^{2} f_{\varepsilon, j}\right. & \left.+\nabla_{\Sigma_{\varepsilon}} f_{\varepsilon, j} \cdot \nabla_{\Sigma}\left(\pi_{\varepsilon, N+1}\right)\right)  \tag{2.21}\\
& -e^{\sqrt{2}\left(f_{\varepsilon, j-1}-f_{\varepsilon, j}\right)}+e^{\sqrt{2}\left(f_{\varepsilon, j}-f_{\varepsilon, j+1}\right)}=0 .
\end{align*}
$$

This system of $k$ equations will be called the Jacobi-Toda system on $\Sigma_{\varepsilon}$. Let us recall that we have set $f_{\varepsilon, 0}=-\infty$ and $f_{\varepsilon, k+1}=\infty$ to close the system. Let us also observe that by scaling back to $\Sigma$ we get the following singular perturbation problem:

$$
\begin{align*}
& \alpha_{0} \varepsilon^{2}\left(\Delta_{\Sigma} f_{j}+\left|A_{\Sigma_{\varepsilon}}\right|^{2} f_{j}+\nabla_{\Sigma} f_{j} \cdot \nabla_{\Sigma}\left(\pi_{N+1}\right)\right)  \tag{2.22}\\
& -e^{\sqrt{2}\left(f_{j-1}-f_{j}\right)}+e^{\sqrt{2}\left(f_{j}-f_{j+1}\right)}=0 .
\end{align*}
$$

Solutions of (2.21) and (2.22) are related through the formula $f_{\varepsilon, j}(\cdot)=f_{j}(\varepsilon \cdot)$. We should mention here that a similar system appears in the context of foliation by interfaces [9, 12].

## 3 An Existence Result for the Jacobi-Toda System

### 3.1 Rotationally Symmetric Eternal Solutions

The formal calculations of the previous section show that to prove Theorem 1.1 we need to find a suitable approximation of the components of the traveling front, and this in turn requires solving the Jacobi-Toda system (2.22). This will be done in several steps in this section. We begin by writing the Jacobi-Toda system for a
special solution of (1.6). Assuming that the surface $\Sigma$ is given as a graph $\Sigma=$ $\left\{x_{N+1}=F\left(x^{\prime}\right), x^{\prime} \in \mathbb{R}^{N}\right\}$, and that $c=1$, we obtain that 1.6 is equivalent to

$$
\begin{equation*}
\nabla\left(\frac{\nabla F}{\sqrt{1+|\nabla F|^{2}}}\right)=\frac{1}{\sqrt{1+|\nabla F|^{2}}} \tag{3.1}
\end{equation*}
$$

We will further assume that $F\left(x^{\prime}\right)=F\left(\left|x^{\prime}\right|\right)$, i.e., $\Sigma$ is rotationally symmetric. Denoting $\left|x^{\prime}\right|=r$ we get

$$
\begin{equation*}
\frac{F_{r r}}{1+F_{r}^{2}}+(N-1) \frac{F_{r}}{r}=1 \tag{3.2}
\end{equation*}
$$

The following result is proven in [1] in the case $N=2$ and in general in [5]:
PROPOSITION 3.1. There exists an entire, rotationally symmetric, and strictly convex graphical eternal solution to the mean curvature flow. This solution satisfies (3.2), and consequently it is translating with speed $c=1$. Additionally, the following asymptotic expansion as $r \rightarrow \infty$ is valid:

$$
\begin{equation*}
F(r)=\frac{r^{2}}{2(N-1)}-\log r+1+\mathcal{O}\left(r^{-1}\right) \tag{3.3}
\end{equation*}
$$

In what follows by $\Gamma$ we will denote the surface corresponding to the rotationally symmetric eternal solution described in Proposition 3.1.

The Jacobi-Toda system 2.22 for $\Gamma$ becomes

$$
\begin{aligned}
\varepsilon^{2} \alpha_{0}\left(\Delta_{\Gamma} f_{j}+\left|A_{\Gamma}\right|^{2} f_{j}+\nabla_{\Gamma} f_{j} \cdot \nabla_{\Gamma}\right. & F) \\
& \quad-e^{\sqrt{2}\left(f_{j-1}-f_{j}\right)}+e^{\sqrt{2}\left(f_{j}-f_{j+1}\right)}=0
\end{aligned}
$$

Our theory of solvability of the Jacobi-Toda system will be valid for functions of the radial variable $r$ only, and so we need to express the Jacobi-Toda system on $\Gamma$ in terms of the radial variable $r$. For what follows it will be convenient to denote

$$
\begin{equation*}
L[v]=\Delta_{\Gamma} v+\left|A_{\Gamma}\right|^{2} v+\nabla_{\Gamma} v \cdot \nabla_{\Gamma} F \tag{3.4}
\end{equation*}
$$

Now, we will find the expression of this operator when restricted to functions $v=v(r)$, i.e., functions depending on the radial variable only. The LaplaceBetrami operator for a surface $x_{N+1}=F(r)$ acting on $v=v(r)$ is

$$
\begin{align*}
\Delta_{\Gamma} v & =\frac{1}{r^{N-1} \sqrt{1+F_{r}^{2}}} \frac{\partial}{\partial r}\left(\frac{r^{N-1}}{\sqrt{1+F_{r}^{2}}} \frac{\partial}{\partial r}\right) v  \tag{3.5}\\
& =\frac{v_{r r}}{1+F_{r}^{2}}+\left(\frac{N-1}{r}-\frac{F_{r}}{1+F_{r}^{2}}\right) v_{r}
\end{align*}
$$

The principal curvatures are given by

$$
\begin{equation*}
\mathbb{k}_{1}=\cdots=\mathbb{k}_{N-1}=\frac{F_{r}}{r \sqrt{1+F_{r}^{2}}}, \quad \mathbb{k}_{N}=\frac{F_{r r}}{\left(1+F_{r}^{2}\right)^{3 / 2}} \tag{3.6}
\end{equation*}
$$

hence

$$
\begin{equation*}
\left|A_{\Gamma}\right|^{2}=\sum_{j=1}^{N} \mathbb{k}_{j}^{2}=\frac{(N-1) F_{r}^{2}}{r^{2}\left(1+F_{r}^{2}\right)}+\frac{F_{r r}^{2}}{\left(1+F_{r}^{2}\right)^{3}} . \tag{3.7}
\end{equation*}
$$

Finally, we have

$$
\nabla_{\Gamma} v \cdot \nabla_{\Gamma} F=\frac{v_{r} F_{r}}{1+F_{r}^{2}}
$$

hence we find the following expression for the operator $L$ acting on radial functions (we denote this operator by $L_{r}$ ):

$$
\begin{equation*}
L_{r}[v]=\frac{v_{r r}}{1+F_{r}^{2}}+\frac{(N-1) v_{r}}{r}+\left(\frac{(N-1) F_{r}^{2}}{r^{2}\left(1+F_{r}^{2}\right)}+\frac{F_{r r}^{2}}{\left(1+F_{r}^{2}\right)^{3}}\right) v . \tag{3.8}
\end{equation*}
$$

### 3.2 Weighted Hölder Norms on $\Gamma$

We will now proceed to define some weighted norms that we will use in what follows. First we recall that, in general, for a function $h$ given on a manifold $\Sigma$ we have, in some local coordinates,

$$
\nabla_{\Sigma} h=g^{i j} \partial_{j} h \partial_{i}, \quad\left(D_{\Sigma}^{2} h\right)_{i j}=g^{i j} \partial_{i j} h-g^{i j} \Gamma_{i j}^{k} \partial_{k} h
$$

We refer to the vector $\nabla_{\Sigma}$ as the gradient and to matrix $D_{\Sigma}^{2}$ as the second derivative matrix of $h$.

Now, in the case at hand we can use the fact that the surface $\Gamma$ is rotationally symmetric to find $\nabla_{\Gamma}$ and $D_{\Gamma}^{2}$. In particular, when $h=h(r)$, i.e., we are dealing with a radial function, then we have the following relations:

$$
\begin{aligned}
\left|\nabla_{\Gamma} h(r)\right| & \leq \frac{C\left|\partial_{r} h(r)\right|}{\sqrt{1+\left|F_{r}(r)\right|^{2}}} \\
\left|\partial_{r} h(r)\right| & \leq C \sqrt{1+\left|F_{r}(r)\right|^{2}}\left|\nabla_{\Gamma} h(r)\right|, \\
\left|D_{\Gamma}^{2} h(r)\right| & \leq \frac{C\left(\left|\partial_{r}^{2} h(r)\right|+r^{-1}\left|\partial_{r} h(r)\right|\right)}{1+\left|F_{r}(r)\right|^{2}}, \\
\left|\partial_{r}^{2} h(r)\right| & \leq C\left(1+\left|F_{r}(r)\right|^{2}\right)\left(\left|D_{\Gamma}^{2} h(r)\right|+\left|\nabla_{\Gamma} h(r)\right|\right)
\end{aligned}
$$

We define the following weighted norms for $\mathcal{C}^{2, \mu}$ functions on $\Gamma$ :

$$
\begin{aligned}
& \|h\|_{\mathcal{C}_{\beta}^{0, \mu}(\Gamma)}=\sup _{y \in \Gamma}\left(1+\left|F_{r}\left(\left|y^{\prime}\right|\right)\right|^{2}\right)^{\beta}\|h\|_{\mathcal{C}^{0, \mu}(B(y, 1) \cap \Gamma)}, \quad y=\left(y^{\prime}, y_{N+1}\right), \\
& \|h\|_{\mathcal{C}_{\beta}^{2, \mu}(\Gamma)}=\|h\|_{\mathcal{C}_{\beta}^{0, \mu}(\Gamma)}+\left\|\nabla_{\Gamma} h\right\|_{\mathcal{C}_{\beta}^{0, \mu}(\Gamma)}+\left\|D_{\Gamma}^{2} h\right\|_{\mathcal{C}_{\beta}^{0, \mu}(\Gamma)} .
\end{aligned}
$$

### 3.3 A Nonhomogeneous Jacobi-Toda System

We observe that so far we were considering the Jacobi-Toda system with the right-hand side equal to 0 . However, as we will see later on, we have to deal with a more general, nonhomogeneous Jacobi-Toda system. This is because in our formal considerations we neglected some terms that are of lower order but need to be
eventually taken into account. Also in this case we assume $k=2$, and thus we get the following problem:

$$
\begin{align*}
& \varepsilon^{2} \alpha_{0} L\left[f_{1}\right]+e^{\sqrt{2}\left(f_{1}-f_{2}\right)}=\varepsilon^{2} h_{1},  \tag{3.9}\\
& \varepsilon^{2} \alpha_{0} L\left[f_{2}\right]-e^{\sqrt{2}\left(f_{1}-f_{2}\right)}=\varepsilon^{2} h_{2},
\end{align*}
$$

where $f_{j}: \Gamma \rightarrow \mathbb{R}, h_{j}: \Gamma \rightarrow \mathbb{R}$. To solve the above problem we will assume that $f_{j}$ and $h_{j}$ are radial functions. In the remaining part of this section we will consider the problem of the existence of solutions to (3.9) under some assumptions about the decay in $r$ and smallness in $\varepsilon$ for the right-hand side. In general, we will assume that

$$
\begin{equation*}
\left\|h_{j}\right\|_{\mathcal{C}_{\beta}^{0, \mu}(\Gamma)} \leq C \varepsilon^{\tau}, \quad \tau>0, \beta>1 \tag{3.10}
\end{equation*}
$$

Let us briefly explain why a nonhomogeneous problem (3.9) with this type of right-hand side appears in our considerations. Going back to the formal calculations in Section 2.2 we notice that in 2.17) we expanded the mean curvature according to (2.6), and we neglected the error term $R_{\Sigma_{\varepsilon}}$. In the case considered here, i.e., $\Sigma=\Gamma$, this term is small in terms of $\varepsilon$, and it decays like $\mathcal{O}\left(\left(1+r^{2}\right)^{-3 / 2}\right)$ when $r \rightarrow \infty$.

We have the following:
Proposition 3.2. Consider the Jacobi-Toda system (3.9) where $h_{j}, j=1,2$, are radial functions satisfying (3.10). There exists a solution of this problem such that the functions $u$ and $v$ defined by

$$
u=\sqrt{2}\left(f_{2}-f_{1}\right), \quad v=\sqrt{2}\left(f_{1}+f_{2}\right)
$$

satisfy

$$
\begin{gather*}
u(r)=\log \frac{2 \sqrt{2}}{\varepsilon^{2} \alpha_{0}\left|A_{\Gamma}(r)\right|^{2}}+\mathcal{O}\left(\log \log \frac{1}{\varepsilon^{2}\left|A_{\Gamma}(r)\right|^{2}}\right), \\
\text { as } \varepsilon r \rightarrow 0^{+} \text {or } \varepsilon r \gg 1,  \tag{3.11}\\
|v(r)| \leq C \varepsilon^{\tau}\left(1+r^{2}\right)^{-\frac{1}{2}} \log \left(2+r^{2}\right),
\end{gather*}
$$

where $\left|A_{\Gamma}(r)\right|$ is the norm of the second fundamental form on $\Gamma$.
To describe the strategy let us denote

$$
h=\frac{\sqrt{2}}{\alpha_{0}}\left(h_{2}-h_{1}\right) \quad \text { and } \quad g=\frac{\sqrt{2}}{\alpha_{0}}\left(h_{2}+h_{1}\right) .
$$

Then we get the following decoupled system:

$$
\begin{align*}
L[u]-\frac{2 \sqrt{2}}{\varepsilon^{2} \alpha_{0}} e^{-u} & =h,  \tag{3.12}\\
L[v] & =g . \tag{3.13}
\end{align*}
$$

Let us discuss briefly the second of the above equations. The key observation is that the operator $L$ has a decaying, positive element in its kernel

$$
\begin{equation*}
\phi_{0}(r)=\frac{1}{\sqrt{1+F_{r}^{2}}} \sim \frac{1}{r}, \quad r \gg 1 \tag{3.14}
\end{equation*}
$$

from which we can solve (3.13) by a standard ODE method.
The solvability theory for the nonlinear equation 3.12 is where the real difficulty lies. Our approach will be to first use an approximation scheme to find a suitable asymptotic approximation of the solution of (3.12), after which we will be in a position to use a fixed point argument to solve the nonhomogeneous problem, with the right-hand side satisfying (3.10).

The following sections are devoted to the proof of Proposition 3.2 ,

### 3.4 Solvability Theory for the Operator $L$

We begin by proving the claim that we have made in (3.14), namely that $\phi_{0}$ is in the kernel of $L$. Note that since $\Gamma$ is an eternal graph solution to the MC flow, then so is $\Gamma+\tau e_{N+1}$, namely, the graph of $x_{N+1}=F\left(x^{\prime}\right)$ translated by $\tau$ in the direction of the $x_{N+1}$-axis. This results in an invariance of the nonlinear operator on the left-hand side of $(3.2)$, which we will take advantage of in the proof of the following:

LEMMA 3.3. The function $\phi_{0}=1 / \sqrt{1+F_{r}^{2}}$ satisfies $L\left[\phi_{0}\right]=0$; i.e., it is $a$ positive, decaying element in ker $L$.

Proof. Let us consider the nonlinear operator

$$
\mathcal{H}(\Phi)=\frac{\Phi_{r r}}{1+\Phi_{r}^{2}}+(N-1) \frac{\Phi_{r}}{r}
$$

Taking variations of this operator of the form $\Phi_{\sigma}=F+\sigma \phi, \phi=\phi(r)$, we get

$$
\left.\frac{d}{d \sigma} \mathcal{H}\left(\Phi_{\sigma}\right)\right|_{\sigma=0} \equiv \mathcal{H}^{\prime}[\phi]=\frac{\phi_{r r}}{1+F_{r}^{2}}-\frac{2 F_{r r} F_{r} \phi_{r}}{\left(1+F_{r}^{2}\right)^{2}}+\frac{(N-1) \phi_{r}}{r}
$$

In particular, we have $\mathcal{H}^{\prime}[1] \equiv 0$. In addition, the following relation is not hard to prove, again assuming that $\phi=\phi(r)$ :

$$
L[\phi]=\mathcal{H}^{\prime}\left[\phi \sqrt{1+F_{r}^{2}}\right]
$$

From this the assertion of the lemma follows immediately.
The existence result for (3.13) follows from the following:
LEMMA 3.4. Let $g$ be a $\mathcal{C}^{0, \mu}(\Gamma)$ radial function such that

$$
\|g\|_{\mathcal{C}_{\beta}^{0, \mu}\left(\mathbb{R}_{+}\right)}<\infty, \quad \beta \geq 1
$$

There exists a unique, bounded solution to

$$
\begin{equation*}
L[v]=g \tag{3.15}
\end{equation*}
$$

such that

$$
\begin{equation*}
\|v\|_{\mathcal{C}_{\beta-1}^{2, \mu}(\Gamma)} \leq C\|g\|_{\mathcal{C}_{\beta}^{0, \mu}(\Gamma)} . \tag{3.16}
\end{equation*}
$$

Proof. Since the function $g$ in (3.15) is radial, we can use ODE methods to solve the equation. Given $\phi_{0}>0$ as in Lemma 3.3, which is a radial solution of $L[\phi]=0$, we find the second linearly independent solution $\phi_{1}$ of $(1+$ $\left.\left|F_{r}(r)\right|^{2}\right) L_{r}[\phi]=0$ (recall that $L_{r}$ is the radial form of $L$ ) by the reduction-oforder formula:

$$
\begin{aligned}
\phi_{1}(r) & =\phi_{0}(r) \int_{r}^{\infty}\left(1+\left|F_{r}(\rho)\right|^{2}\right) \exp [-A(\rho)] d \rho \\
A(\rho) & =\int_{1}^{\rho} \frac{(N-1)\left(1+\left|F_{r}(\eta)\right|^{2}\right)}{\eta} d \eta
\end{aligned}
$$

From this we readily get that

$$
\phi_{1}(r) \sim\left\{\begin{array}{ll}
\log r, & N=2, \\
r^{2-N}, & N>2,
\end{array} \quad r \ll 1, \quad \phi_{1}(r) \sim r e^{-r^{2}}, r \gg 1\right.
$$

Denoting by $W(r)=W(1) \exp [-A(r)]$ the Wronskian, and letting $\widetilde{g}(r)=(1+$ $\left.\left|F_{r}(r)\right|^{2}\right) g(r)$, we write

$$
v(r)=-\phi_{0}(r) \int_{0}^{r} \frac{\phi_{1}(\rho) \widetilde{g}(\rho)}{W(\rho)} d \rho+\phi_{1}(r) \int_{0}^{r} \frac{\phi_{0}(\rho) \widetilde{g}(\rho)}{W(\rho)} d \rho .
$$

The assertion of the lemma follows from a straightforward argument, using the asymptotic formulas for the functions $\phi_{0}(r)$ and $\phi_{1}(r)$.

### 3.5 Solving for $u$ : The Approximate Solution

Our goal in this and the following section is to solve the problem (3.12). Of course, once this is done Proposition 3.2 will be proven. We begin by finding an approximate solution of (3.12) assuming that $h \equiv 0$, which is equivalent to solving

$$
\begin{equation*}
\mathcal{S}_{\delta}[u]=0, \tag{3.17}
\end{equation*}
$$

where

$$
\begin{equation*}
\mathcal{S}_{\delta}[u] \equiv L[u]-\delta^{-2} e^{-u}, \quad \delta=\frac{\varepsilon \sqrt{\alpha_{0}}}{2^{3 / 4}}, \tag{3.18}
\end{equation*}
$$

and $L$ is the linear operator defined in (3.4). For the purpose of finding a suitable approximate solution we will consider a sequence of approximations $\mathrm{v}_{k}=$ $v_{0}+v_{1}+\cdots+v_{k}$. Once an accurate enough approximation is found, the nonlinear problem (3.12) can be reduced to a fixed point problem. This step involves inverting the linear operator obtained by linearization of the nonlinear operator $\mathcal{S}_{\delta}$ around the approximate solution and will be dealt with in the next section.

The nonlinear operator $\mathcal{S}_{\delta}$ can be written explicitly (using the notation of Section 3.1):

$$
\mathcal{S}_{\delta}[v]=\Delta_{\Gamma} v+\nabla_{\Gamma} v \cdot \nabla_{\Gamma} F+\left|A_{\Gamma}\right|^{2} v-\delta^{-2} e^{-v}
$$

We will now describe the construction of an approximate solution of 3.17). The leading-order term of this approximation is found by solving for $v_{0}$ the following equation:

$$
\begin{equation*}
\left|A_{\Gamma}\right|^{2} v_{0}=\frac{1}{\delta^{2}} e^{-v_{0}} \Longrightarrow v_{0} e^{v_{0}}=\frac{1}{\delta^{2}\left|A_{\Gamma}\right|^{2}} \tag{3.19}
\end{equation*}
$$

For brevity we denote $b(r)=\left|A_{\Gamma}(y)\right|^{2}, y=\left(y^{\prime}, y_{N+1}\right), r=\left|y^{\prime}\right|$. Now, equation 3.19) implies that

$$
\begin{equation*}
v_{0}(r)=\log \frac{1}{\delta^{2} b(r)}-\log \log \frac{1}{\delta^{2} b(r)}+\mathcal{O}\left(\log \log \left|\log \delta^{2} b(r)\right|\right) \tag{3.20}
\end{equation*}
$$

This asymptotic formula is valid when $\delta \ll 1$. This follows from the fact that $b(r)=1+\mathcal{O}\left(r^{2}\right), r \rightarrow 0$, and on the other hand $b(r)=\frac{N-1}{r^{2}}+\mathcal{O}\left(r^{-4}\right), r \gg 1$.

Let us also observe the following relations:

$$
\begin{equation*}
v_{0}^{\prime}=-\frac{b^{\prime}}{b} \frac{v_{0}}{1+v_{0}}, \quad v_{0}^{\prime \prime}=-\left(\frac{b^{\prime}}{b}\right)^{\prime} \frac{v_{0}}{1+v_{0}}-\left(\frac{b^{\prime}}{b}\right)^{2} \frac{v_{0}}{\left(1+v_{0}\right)^{3}} \tag{3.21}
\end{equation*}
$$

from which the asymptotic behavior of the derivatives of $v_{0}$ of any order can be readily deduced. In particular, we observe that

$$
\begin{equation*}
\left|v_{0}^{(j)}\right| \leq \frac{C}{(r+1)^{j}}, \quad j=1,2, \ldots \tag{3.22}
\end{equation*}
$$

Accepting $v_{0}$ as the leading-order approximation and assuming that the next approximate solution is of the form $\mathrm{v}_{1}=v_{0}+v_{1}$, we are left with the following problem:

$$
\begin{equation*}
\left|A_{\Gamma}\right|^{2} v_{1}-\frac{1}{\delta^{2}}\left(e^{-v_{0}-v_{1}}-e^{-v_{0}}\right)=-\left[\Delta_{\Gamma} v_{0}+\nabla_{\Gamma} v_{0} \cdot \nabla_{\Gamma} F\right] \equiv \rho_{0} \tag{3.23}
\end{equation*}
$$

This is a nonlinear equation with a right-hand side that satisfies

$$
\begin{equation*}
\left|\rho_{0}(y)\right| \leq \frac{C}{(1+r)^{2}}, \quad r=\left|y^{\prime}\right| \tag{3.24}
\end{equation*}
$$

This follows from the fact that $v_{0}$ is a smooth function on $\Gamma$ and (3.22). Using this fact we can find a smooth solution of the equation 3.23 that satisfies

$$
\begin{equation*}
\left|v_{1}^{(j)}(y)\right| \leq \frac{C}{\log \left(\frac{2+r^{2}}{\delta^{2}}\right)} \frac{1}{(1+r)^{j}}, \quad j=0,1, \ldots \tag{3.25}
\end{equation*}
$$

The next terms in the approximate solutions will be determined inductively. It is important to keep in mind that the approximations we want to construct must be decaying functions of both $1 / \log \delta^{2}$ and $r$. Given $v_{k-1}=v_{0}+v_{1}+\cdots+v_{k-1}$, for which we already know (suitably adapted) relations (3.24)-(3.25), we determine $v_{k}$
by solving

$$
\begin{align*}
& \left|A_{\Gamma}\right|^{2} v_{k}-\frac{1}{\delta^{2}}\left(e^{-v_{0}-v_{1}-\cdots-v_{k}}-e^{-v_{0}-v_{1}-\cdots-v_{k-1}}\right) \\
& \quad=-\left[\Delta_{\Gamma} v_{k-1}+\nabla_{\Gamma} v_{k-1} \cdot \nabla_{\Gamma} F\right]  \tag{3.26}\\
& \quad \equiv \rho_{k-1}
\end{align*}
$$

Solving this equation gives $\mathrm{v}_{k}=v_{0}+v_{1}+\cdots+v_{k}$ where

$$
\begin{equation*}
\left|v_{k}^{(j)}(y)\right| \leq \frac{C}{\left(\log \frac{2+r^{2}}{\delta^{2}}\right)^{k}} \frac{1}{(1+r)^{k+j-1}}, \quad j=0,1, \ldots \tag{3.27}
\end{equation*}
$$

and

$$
\begin{equation*}
\left|\rho_{k}(y)\right|=\left|\Delta_{\Gamma} v_{k}+\nabla_{\Gamma} v_{k} \cdot \nabla_{\Gamma} F\right| \leq \frac{C}{\left(\log \frac{2+r^{2}}{\delta^{2}}\right)^{k}} \frac{1}{(1+r)^{k+1}} \tag{3.28}
\end{equation*}
$$

Thus we have proven the following:
Lemma 3.5. For each $k>1$ there exists a function $v_{k}$ such that

$$
\mathcal{S}_{\delta}\left[\mathrm{v}_{k}\right] \leq \frac{C}{\left(\log \frac{2+r^{2}}{\delta^{2}}\right)^{k}} \frac{1}{(1+r)^{k+1}}
$$

## Another Parametrization of $\boldsymbol{\Gamma}$

The next step in the proof of Proposition 3.2 is to linearize the operator $\mathcal{S}_{\delta}$ around $\mathrm{v}_{k}$ and find a solution of $\mathcal{S}_{\delta}[u]=g$ in the form $u=\mathrm{v}_{k}+h$ using ODE methods.

To have a convenient form of the linear operator $\mathcal{S}_{\delta}^{\prime}\left[{ }^{k} k\right]$ we define another parametrization of $\Gamma$, which is obtained by taking the arc length along the curve $(r, F(r))$. Thus we define

$$
\begin{equation*}
s=\int_{0}^{r} \sqrt{1+F_{r}^{2}} d \rho \tag{3.29}
\end{equation*}
$$

Of course, the function $r \mapsto s(r)$ is invertible and its inverse is $s \mapsto r(s)$. We also note the following relations:

$$
\begin{align*}
c\left|\partial_{s} h\right| & \leq\left|\nabla_{\Gamma} h\right| \leq C\left|\partial_{s} h\right| \\
c\left(\left|\partial_{s}^{2} h\right|+s^{-1}\left|\partial_{s} h\right|\right) & \leq\left|D_{\Gamma}^{2} h\right| \leq C\left(\left|\partial_{s}^{2} h\right|+s^{-1}\left|\partial_{s} h\right|\right) \tag{3.30}
\end{align*}
$$

Using the asymptotic formula 3.3 for $F$, we get that

$$
\begin{equation*}
s \sim r, \quad r \ll 1, \quad \text { and } \quad s=\frac{r^{2}}{2(N-1)}+\mathcal{O}(\log r), \quad r \gg 1 \tag{3.31}
\end{equation*}
$$

By a straightforward computation we obtain the following expression for the operator $L$ but now in terms of the arc length variable $s$ :

$$
\begin{equation*}
L_{s}[v]=v_{s s}+a(s) v_{s}+b(s) v \tag{3.32}
\end{equation*}
$$

where

$$
\begin{equation*}
a(s)=\frac{F_{r}(r(s))+\frac{N-1}{r(s)}}{\sqrt{1+F_{r}^{2}(r(s))}}, \quad b(s)=\left|A_{\Gamma}(r(s))\right|^{2} \tag{3.33}
\end{equation*}
$$

Note that

$$
\begin{align*}
& a(s)= \begin{cases}\frac{N-1}{s}\left(1+\mathcal{O}\left(s^{2}\right)\right), & s \ll 1, \\
1+\mathcal{O}\left(s^{-1}\right), & s \gg 1,\end{cases}  \tag{3.34}\\
& b(s)=\frac{N-1}{r^{2}(s)}+\mathcal{O}\left(r^{-4}\right)=\frac{1}{2 s}+\mathcal{O}\left(s^{-2} \log s\right), \quad s \gg 1,
\end{align*}
$$

and that in general $a(s), b(s)>0$ since $\Gamma$ is convex and $F_{r}(0)=0$. We also have $b(0)=1$ and $b^{\prime}(0)=0$. Another important fact is that

$$
\begin{equation*}
b^{\prime \prime}(0)=-\frac{N^{2}+4 N+2}{N^{4}(N+2)}<0, \quad N=2,3, \ldots \tag{3.35}
\end{equation*}
$$

This last identity follows by a direct computation. Setting $\mathfrak{b}_{N}=\frac{N^{2}+4 N+2}{2 N^{4}(N+2)}$, we have

$$
\begin{equation*}
b(s)=1-\mathfrak{b}_{N} s^{2}+\mathcal{O}\left(s^{4}\right), \quad s \rightarrow 0 \tag{3.36}
\end{equation*}
$$

## Definition of the Linearized Operator $\mathfrak{L}_{\boldsymbol{\delta}}$

From the above considerations we see that linearization of $\mathcal{S}_{\delta}$ around the approximate solution $\mathrm{v}_{k}$ expressed in terms of $r$ is the following operator:

$$
\begin{equation*}
\mathfrak{L}_{\delta}[h]=\frac{h_{r r}}{1+F_{r}^{2}}+\frac{N-1}{r} h_{r}+p_{\delta}(r) h, \quad p_{\delta}(r)=b(r)\left(1+v_{0} e^{-v_{k}+v_{0}}\right) . \tag{3.37}
\end{equation*}
$$

We will often use the approximate solution $\mathrm{v}_{k}$ expressed in terms of the arc length variable $s$, which we will denote by $\mathrm{u}_{k}(s)=\mathrm{v}_{k}(r(s))$. We will also set $u_{j}(s)=$ $v_{j}(r(s)), j=0,1, \ldots$ We let $b(s)=b(r(s))$.

Later on we will consider the linearized operator in the space of functions that decay in both $s$ and $\log \left(\frac{s}{\delta^{2}}\right)$ as $s$ increases. We will see that for our purposes we need to determine $\mathrm{v}_{k}$ (or $\mathrm{u}_{k}$ ) for $k$ sufficiently large.

With some abuse of notation we will denote by the same symbol $\mathfrak{L}_{\delta}$ the linearized operator expressed in terms of the arc length variable $s$ :

$$
\begin{align*}
& \mathfrak{L}_{\delta}[h]=h_{s s}+a(s) h_{s}+p_{\delta}(s) h, \\
& p_{\delta}(s)=b(s)\left(1+u_{0}(s) e^{-\left(u_{1}(s)+\cdots+u_{k}(s)\right)}\right) \tag{3.38}
\end{align*}
$$

Our goal is to find a right inverse of $\mathfrak{L}_{\delta}$. The idea is very simple. Since (3.38) is an ODE, an inverse can always be written using the variation-of-parameters formula. To control the norm of $\mathfrak{L}_{\delta}^{-1}$, we need to understand the behavior of a fundamental set. This is complicated by the fact that on the one hand the operator depends on $\delta$ and on the other hand its properties change as $s$ varies from 0 to $\infty$.

In fact, we observe that from $3.19-(3.20$ and $3.26-3.27$ ) it follows that

$$
\begin{equation*}
p_{\delta}(s) \sim \log \frac{1}{\delta^{2}} \tag{3.39}
\end{equation*}
$$

when $s \leq \bar{s}$ with some $\bar{s}>0$ fixed, independent of $\delta$, while when $s \gg 1$ we have

$$
\begin{equation*}
p_{\delta}(s) \sim \frac{\log \frac{s}{\delta^{2}}}{s} \tag{3.40}
\end{equation*}
$$

This can be summarized:

$$
p_{\delta}(s) \sim \frac{1}{2+s} \log \left(\frac{2+s}{\delta^{2}}\right)
$$

for all $s$ and $\delta \ll 1$. At the same time $a(s) \sim \frac{1}{s}, s \ll 1$, and $a(s) \sim 1, s \gg 1$. In particular, we will need to study carefully $\mathfrak{L}_{\delta}$ in these ranges of $s$.

### 3.6 An Inverse of $\mathfrak{L}_{\boldsymbol{\delta}}$

In this section, we solve the problem

$$
\begin{equation*}
\mathfrak{L}_{\delta}[h]=g(s) \tag{3.41}
\end{equation*}
$$

Clearly, solving this problem is the key to implementing a fixed point argument needed to solve (3.12). The point is to construct a right inverse of $\mathfrak{L}_{\delta}$ that is bounded in suitable Hölder weighted norms. Let us define these norms first:

$$
\begin{align*}
\|g\|_{\mathcal{C}_{\beta, v}^{0, \mu}\left(\mathbb{R}_{+}\right)} & =\sup _{s>1}\left\{(2+s)^{\beta}\left(\log \frac{2+s}{\delta^{2}}\right)^{v}\|g\|_{\mathcal{C}^{0, \mu}((s-1, s+1))}\right\} \\
\|g\|_{\mathcal{C}_{\beta, v}^{\ell, \mu}\left(\mathbb{R}_{+}\right)} & :=\sum_{j=0}^{\ell}\left\|g^{(j)}\right\|_{\mathcal{C}_{\beta, \nu}^{0, \mu}\left(\mathbb{R}_{+}\right)} \tag{3.42}
\end{align*}
$$

Because of the relations 3.30 these norms are easily translated into the norms of $g$ as a function (of the radial variable) on $\Gamma$.

More precisely, we will show the following:
Lemma 3.6. Suppose that $\beta>0, v>0$. Then there exists a constant $C>0$ and a solution $h$ to 3.41 such that

$$
\begin{align*}
&\|h\|_{\mathcal{C}_{\beta, v}^{0, \mu}\left(\mathbb{R}_{+}\right)}+\left\|h^{\prime}\right\|_{\mathcal{C}_{\beta+1, v}^{0, \mu}\left(\mathbb{R}_{+}\right)}+\left\|h^{\prime \prime}\right\|_{\mathcal{C}_{\beta+1, v}^{0, \mu}\left(\mathbb{R}_{+}\right)} \leq  \tag{3.43}\\
& C\left(\log \frac{1}{\delta^{2}}\right)^{4+2 \beta}\|g\|_{\mathcal{C}_{\beta+1, v+1}^{0, \mu}}\left(\mathbb{R}_{+}\right)
\end{align*}
$$

In the rest of this section we prove this important lemma.
To begin with, we make the following transformation:

$$
\begin{equation*}
\hat{h}=\exp \left(\frac{1}{2} \int_{1}^{s} a(\tau) d \tau\right) h \tag{3.44}
\end{equation*}
$$

Then, when $s \rightarrow 0, \hat{h} \sim s^{(N-1) / 2} h$, and when $s \rightarrow+\infty, \hat{h} \sim e^{s / 2} h$, by 3.34). Equation (3.41) is transformed to

$$
\begin{equation*}
\hat{h}^{\prime \prime}+\left(p_{\delta}(s)-\hat{a}(s)\right) \hat{h}(s)=\widehat{g}, \tag{3.45}
\end{equation*}
$$

where

$$
\hat{a}=\frac{1}{2} a^{\prime}+\frac{1}{4} a^{2}, \quad \hat{g}=\exp \left(\frac{1}{2} \int_{1}^{s} a(\tau) d \tau\right) g .
$$

In what follows we will mainly work with the transformed equation (3.45). The idea of the proof of the lemma follows the same lines as the construction of the approximate solutions. The situation now is more complicated since we have to consider a second-order ODE.

Let us denote

$$
\widehat{\mathfrak{L}}_{\delta}[h]=h^{\prime \prime}+\hat{p}_{\delta} h, \quad \hat{p}_{\delta}=p_{\delta}-\hat{a} .
$$

When we consider the operator $\hat{\mathfrak{L}}_{\delta}$ for functions defined in the interval $I_{1}=\left(0, s_{1}\right)$ for some $s_{1}>0$ independent of $\delta$, then we refer to this problem as the inner problem. We speak of the outer problem when we take $I_{s_{\delta}}=\left(s_{\delta}, \infty\right), s_{\delta} \gg s_{1}>$ 0 , as the domain of the functions involved.

First, we will describe the way we choose $s_{1}$ and $s_{\delta}$. For $s \rightarrow 0$, we have, by (3.34)-(3.36)

$$
\begin{align*}
p_{\delta}(s) & =\left(1-\mathfrak{b}_{N} s^{2}+\mathcal{O}\left(s^{4}\right)\right)\left(\log \frac{1}{\delta^{2}}+1+\mathcal{O}\left(s^{2}\right)\right) \\
\hat{a}(s) & =s^{-2}\left[\frac{(N-2)^{2}}{4}-\frac{1}{4}\right]+\mathcal{O}(1) \tag{3.46}
\end{align*}
$$

As a consequence, there exist

$$
M>0 \quad \text { and } \quad s_{1}>\frac{M}{\sqrt{\log \frac{1}{\delta^{2}}}}>0
$$

which are independent of $\delta$, such that

$$
\begin{equation*}
\hat{p}_{\delta}(s)=p_{\delta}(s)-\hat{a}(s)>0, \quad \frac{M}{\sqrt{\log \frac{1}{\delta^{2}}}} \leq s \leq s_{1} . \tag{3.47}
\end{equation*}
$$

When $s \rightarrow \infty$ we have by (3.34) that $p_{\delta}$ satisfies (3.40) and

$$
\begin{equation*}
\widehat{a}(s)=\frac{1}{4}+\mathcal{O}\left(s^{-1}\right) \tag{3.48}
\end{equation*}
$$

with similar formulas for the derivatives. From this we can find the asymptotic behavior of $\hat{p}_{\delta}(s)$ for $s$ large and infer the existence of $s_{2} \geq s_{1}$, again independent of $\delta$, such that for $s>s_{2}$ it holds that

$$
\begin{equation*}
\hat{p}_{\delta}^{\prime}(s) \leq 0 . \tag{3.49}
\end{equation*}
$$

Observe that $s_{1}$ and $s_{2}$ in general do not coincide and we need to solve an intermediate problem to glue the inner solution and the solution for $s$ between $s_{1}$ and $s_{2}$.

Finally, we will assume that $\delta$ is chosen sufficiently small so that

$$
\begin{equation*}
\hat{p}_{\delta}(s)>0, \quad s_{1}<s<s_{2} . \tag{3.50}
\end{equation*}
$$

This can be achieved since, when $s$ is bounded away from 0 and $\infty$ independently of $\delta$, we have $\hat{p}_{\delta}(s) \sim p_{\delta}(s) \sim b(s) \log \frac{1}{\delta^{2}}$. For future reference we observe that from (3.49) and (3.50), it follows that there exists a unique $s_{\delta}$ such that $\hat{p}_{\delta}\left(s_{\delta}\right)=0$ and

$$
\begin{equation*}
\hat{p}_{\delta}(s)>0, \quad s_{1} \leq s<s_{\delta}, \quad \hat{p}_{\delta}(s)<0, \quad s>s_{\delta} \tag{3.51}
\end{equation*}
$$

Actually, from (3.34) it follows that there exist constants $M_{1}<M_{2}$ such that

$$
\begin{equation*}
s_{\delta} \in\left(M_{1} \log \frac{1}{\delta^{2}}, M_{2} \log \frac{1}{\delta^{2}}\right) \tag{3.52}
\end{equation*}
$$

One more observation we make is that on any interval $I=\left(0, s^{*}\right)$, with $s^{*}<$ $C \log \frac{1}{\delta^{2}}$, the norms $\|\cdot\|_{\mathcal{C}_{\beta, \nu}^{\ell, \mu}(I)}$ and $\|\cdot\|_{\mathcal{C}_{\beta, 0}^{\ell, \mu}(I)}$ are equivalent in the following sense:

$$
\|g\|_{\mathcal{C}_{\beta, v}^{\ell, \mu}(I)} \leq C\left(\log \frac{1}{\delta^{2}}\right)^{v}\|g\|_{\mathcal{C}_{\beta, 0}^{\ell, \mu}(I)} \leq C\|g\|_{\mathcal{C}_{\beta, v}^{\ell, \mu}(I)}
$$

We agree that $\|\cdot\|_{\mathcal{C}_{\beta, 0}^{\ell, \mu}(I)}=\|\cdot\|_{\mathcal{C}_{\beta}^{\ell, \mu}(I)}$. We will use this equivalence of norms when we consider the operator $\mathfrak{L}_{\delta}$ on the interval $\left(0, s_{\delta}\right)$.

## Inner Problem for the Operator $\mathfrak{L}_{\boldsymbol{\delta}}$

In this section we will consider the following problem:

$$
\begin{array}{cl}
\mathfrak{L}_{\delta}\left[h_{i}\right]=g & \text { in } I_{1}=\left(0, s_{1}\right) \\
h_{i}(0)=0, & h_{i}^{\prime}(0)=0 \tag{3.53}
\end{array}
$$

Our goal is to show that there exists a unique solution $h_{i}$ to 3.53 such that

$$
\begin{equation*}
\left\|h_{i}\right\|_{\mathcal{C}_{\beta}^{2, \mu}\left(I_{1}\right)} \leq C \log \frac{1}{\delta^{2}}\|g\|_{\mathcal{C}_{\beta+1}^{0, \mu}\left(I_{1}\right)} \tag{3.54}
\end{equation*}
$$

We will work with the transformed operator $\mathfrak{L}_{\delta}$ so that 3.53 becomes

$$
\begin{gather*}
\widehat{\mathfrak{L}}_{\delta}\left[\hat{h}_{i}\right]=\widehat{g} \quad \text { in } I_{1}=\left(0, s_{1}\right) \\
\hat{h}_{i}(0)=0, \quad \widehat{h}_{i}^{\prime}(0)=0 \tag{3.55}
\end{gather*}
$$

For convenience we will denote $\lambda=\sqrt{1+\log \left(1 / \delta^{2}\right)}$. Taking into account the asymptotic behavior of $b(s)$ and $\hat{a}(s)$ when $s \rightarrow 0$, we see that the operator $\widehat{\mathfrak{L}}_{\delta}$ can be written in the form

$$
\widehat{\mathfrak{L}}_{\delta}[\widehat{h}]=\widehat{h}^{\prime \prime}+\left[\lambda^{2}-s^{-2}\left(\frac{(N-2)^{2}}{4}-\frac{1}{4}\right)\right]\left(1+\mathcal{O}\left(s^{2}\right)\right) \hat{h}
$$

It is convenient to make a further change of variables:

$$
\widehat{h}_{i}(s)=\tilde{h}_{i}(\lambda s), \quad \widehat{g}(s)=\widetilde{g}(\lambda s), \quad \hat{p}_{\delta}(s)=\lambda^{-2} \widetilde{p}(\lambda s), \quad \text { etc. }
$$

Then, denoting by $\tilde{\mathfrak{L}}_{\delta}$ the rescaled operator, we have

$$
\tilde{\mathfrak{L}}_{\delta}[\widetilde{h}]=\widetilde{h}^{\prime \prime}+\left[1-s^{-2}\left(\frac{(N-2)^{2}}{4}-\frac{1}{4}\right)\right]\left(1+\mathcal{O}\left(\lambda^{-2} s^{2}\right)\right) \tilde{h}
$$

and 3.55 becomes

$$
\tilde{\mathfrak{L}}_{\delta}\left[\tilde{h}_{i}\right]=\lambda^{-2} \widetilde{g} \quad \text { in } I_{\lambda}=\left(0, \lambda s_{1}\right)
$$

Formally $\tilde{\mathfrak{L}}_{\delta}[\tilde{h}]=0$ resembles the modified Bessel equation, and the operator $\tilde{\mathfrak{L}}_{\delta}$ should have an element of the kernel $\tilde{h}_{i, 1}$ such that

$$
\begin{equation*}
\tilde{h}_{i, 1}(s) \sim s^{\frac{1}{2}} J_{\frac{N-2}{2}}(s) \tag{3.56}
\end{equation*}
$$

where $J_{(N-2) / 2}(s)$ is the Bessel function. The second linearly independent element in the kernel is such that

$$
\begin{equation*}
\tilde{h}_{i, 2}(s) \sim s^{1 / 2} J_{(-N+2) / 2}(s) \tag{3.57}
\end{equation*}
$$

when $\frac{N-2}{2}$ is not an integer and

$$
\tilde{h}_{i, 2}(s) \sim s^{1 / 2} Y_{(N-2) / 2}(s)
$$

when $\frac{N-2}{2}$ is an integer, where $Y_{(N-2) / 2}$ is the modified Bessel function of the second kind [6].

We choose the solution to 3.55 given by

$$
\begin{align*}
\tilde{h}_{i}(s)= & -\lambda^{-2} \tilde{h}_{i, 1}(s) \int_{0}^{s} \tilde{h}_{i, 2}(\tau) \widetilde{g}(\tau) d \tau  \tag{3.58}\\
& +\lambda^{-2} \tilde{h}_{i, 2}(s) \int_{0}^{s} \tilde{h}_{i, 1}(\tau) \widetilde{g}(\tau) d \tau
\end{align*}
$$

Note that $\tilde{h}_{i}(0)=0$ and $\tilde{h}_{i}^{\prime}(0)=0$ since, after the change of variables, we have $\tilde{g}(s)=\mathcal{O}\left(s^{(N-1) / 2}\right)$.

To make use of the above formula and to estimate $\tilde{h}_{i}$ we need some information about the functions $\widetilde{h}_{i, j}, j=1,2$. We recall that the Bessel functions oscillate, and the same is expected for $\widetilde{h}_{i, j}$. We observe first that passing to the limit over compacts we can justify the asymptotic statements (3.56-(3.57) and show the uniform convergence of $\widetilde{h}_{i, j}$ to the corresponding solutions of the Bessel equation as $\lambda \rightarrow \infty$. In particular, it follows that for each $K>0$ and each sufficiently large $\lambda$ the function $\tilde{h}_{i, 1}$ is uniformly bounded on the interval $(0, K)$, and for each small $\tau>0$ the function $\tilde{h}_{i, 2}$ is uniformly bounded over the interval ( $\left.\tau, K\right)$. Furthermore, taking $K$ sufficiently large, we may assume that

$$
\tilde{p}(s)=\left[1-s^{-2}\left(\frac{(N-2)^{2}}{4}-\frac{1}{4}\right)\right]\left(1+\mathcal{O}\left(\lambda^{-2} s^{2}\right)\right)>0, \quad s \in\left(K, \lambda s_{1}\right)
$$

In fact, we even have

$$
c_{1} \leq \tilde{p}(s) \leq c_{2}, \quad s \in\left(K, \lambda s_{1}\right)
$$

with some constants $c_{1}, c_{2}>0$. Now we will make an important observation: let $\widetilde{h}$ be a solution of $\widetilde{\mathfrak{L}}_{\delta}[\widetilde{h}]=0$ in ( $K, \lambda s_{1}$ ) and consider the following expressions:

$$
Q_{1}(\tilde{h}) \equiv\left[\tilde{h}^{\prime}(s)\right]^{2}+\widetilde{p}(s)[\widetilde{h}(s)]^{2}, \quad Q_{2}(\tilde{h})=\frac{\left[\tilde{h}^{\prime}(s)\right]^{2}}{\widetilde{p}(s)}+[\widetilde{h}(s)]^{2} .
$$

It is easy to see that

$$
\begin{equation*}
\frac{d}{d s} Q_{1}(\tilde{h})=\widetilde{p}^{\prime}(\tilde{h})^{2}, \quad \frac{d}{d s} Q_{2}(\tilde{h})=-\frac{\tilde{p}^{\prime}}{\widetilde{p}}\left(\tilde{h}^{\prime}\right)^{2} . \tag{3.59}
\end{equation*}
$$

Let now $K \leq \xi_{1}<\xi_{2}<\lambda s_{1}$ be two points such that $\widetilde{h}^{\prime}\left(\xi_{j}\right)=0$. Then from (3.59) and the bound on $\widetilde{p}$, it follows that there exist constants $C_{1}$ and $C_{2}$ such that

$$
\begin{equation*}
C_{2}\left[\tilde{h}\left(\xi_{1}\right)\right]^{2} \geq\left[\tilde{h}\left(\xi_{2}\right)\right]^{2} \geq C_{1}\left[\tilde{h}\left(\xi_{1}\right)\right]^{2} \tag{3.60}
\end{equation*}
$$

as long as $\widetilde{p}^{\prime}$ does not change sign in the interval $\left(\xi_{1}, \xi_{2}\right)$ (recall that $\widetilde{p}>0$ in $\left(K, \lambda s_{1}\right)$ ).

We claim that from this and the uniform bound for the functions $\tilde{h}_{i, j}$ for $s<K$, which we have already proven, it follows that these functions are actually bounded uniformly for $s \geq K$ as well. To prove this we observe that from (3.46) it follows that
$\widetilde{p}(s)=\left\{\left[1-\mathfrak{b}_{N} \lambda^{-2} s^{2}+\mathcal{O}\left(\lambda^{-4} s^{4}\right)\right]-s^{-2}\left[\frac{(N-2)^{2}}{4}-\frac{1}{4}\right]\right\}\left(1+\mathcal{O}\left(\lambda^{-2} s^{2}\right)\right) ;$
hence when $N=2,3$ we have $\tilde{p}^{\prime}(s)<0$ for all $s \in\left(0, \lambda s_{1}\right)$, while when $N>3$ there exists a unique $s_{\lambda}<C \sqrt{\lambda}$ such that

$$
\tilde{p}^{\prime}(s)>0, \quad s \in\left(0, s_{\lambda}\right), \quad \tilde{p}^{\prime}(s)<0, \quad s \in\left(s_{\lambda}, \lambda s_{1}\right) .
$$

Therefore when $N=2,3$ the uniform bound on $\tilde{h}_{i, j}$ follows immediately from (3.60). When $N>3$ we need to consider the growth of $\widetilde{h}_{i, j}$ between $\zeta_{1}<s_{\lambda}<\zeta_{2}$ where $\zeta_{\ell}$ are zeros of $\widetilde{h}_{i, j}$. Observe that since $\widetilde{p}(s)$ is bounded uniformly for $0<$ $s<\lambda s_{1}$, therefore using the relations (3.59) but now considering those points $\zeta$ at which $\widetilde{h}_{i, j}(\zeta)=0$, we get, as long as $\zeta<s_{\lambda}$, that $\left[\widetilde{h}_{i, j}^{\prime}(\zeta)\right]^{2}$ is bounded uniformly in $\lambda$. Then, for each $s \in\left(\zeta_{1}, s_{\lambda}\right)$, we get

$$
\frac{d}{d s} Q_{2}\left(\widetilde{h}_{i, j}\right)(s) \leq 0 \Longrightarrow C\left[\widetilde{h}_{i, j}^{\prime}\left(\zeta_{1}\right)\right]^{2} \geq\left[\widetilde{h}_{i, j}^{\prime}(s)\right]^{2}+\widetilde{p}(s)\left[\tilde{h}_{i, j}(s)\right]^{2},
$$

and in particular $\left[\widetilde{h}_{i, j}^{\prime}\left(s_{\lambda}\right)\right]^{2}+\left[\widetilde{h}_{i, j}\left(s_{\lambda}\right)\right]^{2}$ is bounded. A similar argument, but using $Q_{1}\left(\widetilde{h}_{i, j}\right)(s)$ for $s \in\left(s_{\lambda}, \zeta_{2}\right)$, gives that $\left[\widetilde{h}_{i, j}^{\prime}(s)\right]^{2}+\left[\widetilde{h}_{i, j}(s)\right]^{2}$ is bounded as well. Now (3.60) applies in $\left(\zeta_{2}, \lambda s_{1}\right)$ and the claim follows.

The asymptotic formulas (3.56-3.57) for $s$ small, and the uniform bound on $\tilde{h}_{i, j}$, together with the variation-of-parameters formula (3.58, give the following bound:

$$
\begin{equation*}
\left\|s^{\frac{1-N}{2}} \widetilde{h}_{i}\right\|_{\mathcal{C}^{0}(0, K)} \leq \frac{C}{\lambda^{2}}\left\|s^{2+\frac{1-N}{2}} \widetilde{g}\right\|_{\mathcal{C}^{0}(0, K)} . \tag{3.61}
\end{equation*}
$$

On the other hand, the uniform bounds on $\widetilde{h}_{i, j}$ yield

$$
\begin{equation*}
\left\|s^{\frac{1-N}{2}} \widetilde{h}_{i}\right\|_{\mathcal{C}^{0}\left(K, \lambda s_{1}\right)} \leq \frac{C}{\lambda^{2}}\left\|s^{1+\frac{1-N}{2}} \tilde{g}\right\|_{\mathcal{C}^{0}\left(K, \lambda s_{1}\right)} . \tag{3.62}
\end{equation*}
$$

Scaling back these estimates we get for the solution of the inner problem the estimate

$$
\left\|h_{i}\right\|_{\mathcal{C}^{0, \mu}\left(I_{1}\right)} \leq C\|g\|_{\mathcal{C}^{0, \mu}\left(I_{1}\right)} .
$$

Using then equation (3.53), we can write

$$
h_{s s}+a(s) h_{s}=g-p_{\delta}(s) h,
$$

and since $p_{\delta}(s) \sim \log \frac{1}{\delta^{2}}$ on $I_{1}$,

$$
\left\|h_{i}\right\|_{\mathcal{C}^{2, \mu}\left(I_{1}\right)} \leq C \log \frac{1}{\delta^{2}}\|g\|_{\mathcal{C}^{0, \mu}\left(I_{1}\right)}
$$

from where we get (3.54), using the fact that on the interval $I_{1}=\left(0, s_{1}\right)$, with $s_{1}$ bounded independently of $\delta$, the weight in the definition of the $\mathcal{C}_{\beta}^{0, \mu}$ norm is bounded by a constant.

## Continuation of the Solution from $s=s_{1}$ to $s=s_{2}$

Let $s_{1}<s_{2}$ be as defined above (see (3.47)-( 3.50 ). We will solve now

$$
\begin{array}{ll}
\hat{\mathfrak{L}}_{\delta}\left[\hat{h}_{n}\right]=\hat{g} & \text { in } I_{2}=\left(s_{1}, s_{2}\right), \\
\hat{h}_{n}\left(s_{1}\right)=\hat{h}_{i}\left(s_{1}\right), & \hat{h}_{n}^{\prime}\left(s_{1}\right)=\widehat{h}_{i}^{\prime}\left(s_{1}\right) \tag{3.63}
\end{array}
$$

Let us recall that in the interval considered here we have $\hat{p}_{\delta}(s)>0, \hat{p}_{\delta}(s) \sim$ $b(s) \log \frac{1}{\delta^{2}}$, and $s_{2}$ is a point such that $p_{\delta}^{\prime}(s) \leq 0$ for $s>s_{2}$.

The solution of (3.63) can be written using the variation-of-parameters formula

$$
\begin{align*}
\hat{h}_{n}(s)= & \hat{h}_{n, 1}(s) \hat{h}_{i}\left(s_{1}\right)+\hat{h}_{n, 2}(s) \hat{h}_{i}^{\prime}\left(s_{1}\right)-\hat{h}_{n, 1}(s) \int_{s_{1}}^{s} \hat{h}_{2, n}(\tau) \hat{g}(\tau) d \tau  \tag{3.64}\\
& +\hat{h}_{n, 2}(s) \int_{s_{1}}^{s} \hat{h}_{1, n}(\tau) \hat{g}(\tau) d \tau,
\end{align*}
$$

where the $\hat{h}_{n, j}$ form a fundamental set of the ODE (3.63) with

$$
\hat{h}_{n, 1}\left(s_{1}\right)=1=\widehat{h}_{n, 2}^{\prime}\left(s_{1}\right), \quad \widehat{h}_{n, 1}^{\prime}\left(s_{1}\right)=0=\widehat{h}_{n, 2}\left(s_{1}\right) .
$$

Using the fact that, by the choices of $s_{1}, s_{2}$, and $\delta$ in (3.47)-(3.50), $\hat{p}_{\delta}(s)>c>0$ in $I_{2}$, we can employ the identities (3.59] to obtain a uniform bound on $\left[\hat{h}_{n, j}(s)\right]^{2}$ and $\left[\hat{h}_{n, j}^{\prime}(s)\right]^{2}$ in $I_{2}$.

Then from the estimate on $\hat{h}_{j}\left(s_{1}\right)$ and $\hat{h}_{i}^{\prime}\left(s_{1}\right)$ and (3.64) we get, after changing back to the original functions $h_{n}$ and $g$,

$$
\begin{equation*}
\left\|h_{n}\right\|_{\mathcal{C}^{0}, \mu}\left(I_{2}\right) \leq C \log \frac{1}{\delta^{2}}\|g\|_{\mathcal{C}^{0, \mu}\left(I_{1} \cup I_{2}\right)} \tag{3.65}
\end{equation*}
$$

hence we get, again using the equation

$$
\begin{equation*}
\left\|h_{n}\right\|_{\mathcal{C}^{2, \mu}\left(I_{2}\right)} \leq C\left(\log \frac{1}{\delta^{2}}\right)^{2}\|g\|_{\mathcal{C}^{0, \mu}\left(I_{1} \cup I_{2}\right)} \tag{3.66}
\end{equation*}
$$

and since $s_{2}$ is bounded,

$$
\begin{equation*}
\left\|h_{n}\right\|_{\mathcal{C}_{\beta}^{2}, \mu\left(I_{2}\right)} \leq C\left(\log \frac{1}{\delta^{2}}\right)^{2}\|g\|_{\mathcal{C}_{\beta+1}^{0, \mu}\left(I_{1} \cup I_{2}\right)} . \tag{3.67}
\end{equation*}
$$

Continuation of the Solution from $s=s_{2}$ to $s=s_{\delta}$
Next we will solve

$$
\begin{array}{ll}
\hat{\mathfrak{L}}_{\delta}\left[\hat{h}_{m}\right]=\hat{g} & \text { in } I_{3}=\left(s_{2}, s_{\delta}\right), \\
\hat{h}_{m}\left(s_{2}\right)=\hat{h}_{n}\left(s_{2}\right) & \hat{h}_{m}^{\prime}\left(s_{2}\right)=\widehat{h}_{n}^{\prime}\left(s_{2}\right), \tag{3.68}
\end{array}
$$

where $s_{\delta}$ is defined in (3.51). Notice that in $I_{3}$ we have $\hat{p}_{\delta}^{\prime}(s)<0$; however, $\hat{p}_{\delta}(s)$ is not bounded away from 0 since by definition of $s_{\delta}, \hat{p}_{\delta}\left(s_{\delta}\right)=0$. But we can still use the quadratic form $Q_{1}(h)$ in $\sqrt{3.59}$ to find a uniform bound on $\left[\hat{h}_{m, j}^{\prime}(s)\right]^{2}$, where the $\hat{h}_{m, j}$ are elements of a fundamental set. From this we find

$$
\begin{equation*}
\left|\widehat{h}_{m, j}(s)\right| \leq C\left(1+\left(s-s_{2}\right)\right), \quad s \in I_{3} . \tag{3.69}
\end{equation*}
$$

Then, the variation-of-parameters formula gives

$$
\begin{align*}
\hat{h}_{m}(s)= & \hat{h}_{m, 1}(s) \hat{h}_{n}\left(s_{2}\right)+\hat{h}_{m, 2}(s) \hat{h}_{n}^{\prime}\left(s_{2}\right)-\hat{h}_{m, 1}(s) \int_{s_{2}}^{s} \hat{h}_{2, m}(\tau) \widehat{g}(\tau) d \tau  \tag{3.70}\\
& +\hat{h}_{m, 2}(s) \int_{s_{2}}^{s} \hat{h}_{1, m}(\tau) \widehat{g}(\tau) d \tau .
\end{align*}
$$

Multiplying this identity by $\exp \left\{-\frac{1}{2} \int_{1}^{s} a(\tau) d \tau\right\}$ and using (3.69) we infer that the function

$$
h_{m}(s)=\hat{h}_{m}(s) \exp \left\{-\frac{1}{2} \int_{1}^{s} a(\tau) d \tau\right\}
$$

satisfies

$$
\left\|h_{m}\right\|_{\mathcal{C}^{0, \mu}\left(I_{3}\right)} \leq C\left(\log \frac{1}{\delta^{2}}\right)\left(\left|h_{n}\left(s_{2}\right)\right|+\left|h_{n}{ }^{\prime}\left(s_{2}\right)\right|\right)+C\left(\log \frac{1}{\delta^{2}}\right)^{2}\|g\|_{\mathcal{C}_{\beta+1}^{0, \mu}\left(I_{3}\right)} .
$$

Taking into account (3.66) we find

$$
\begin{equation*}
\left\|h_{m}\right\|_{\mathcal{C}^{0, \mu}\left(I_{3}\right)} \leq C\left(\log \frac{1}{\delta^{2}}\right)^{3}\|g\|_{\mathcal{C}_{\beta+1}^{0, \mu}\left(I_{1} \cup I_{2} \cup I_{3}\right)}, \tag{3.71}
\end{equation*}
$$

and then using the equation $\mathfrak{L}_{\delta}\left[h_{m}\right]=g$ in $I_{3}$ :

$$
\begin{equation*}
\left\|h_{m}\right\|_{\mathcal{C}^{2}, \mu\left(I_{3}\right)} \leq C\left(\log \frac{1}{\delta^{2}}\right)^{4}\|g\|_{\mathcal{C}_{\beta+1}^{0, \mu}\left(I_{1} \cup I_{2} \cup I_{3}\right)} . \tag{3.72}
\end{equation*}
$$

Finally, noting that for $s_{2}<s<s_{\delta}$ we have $(2+s)^{\beta} \leq C\left(\log 1 / \delta^{2}\right)^{\beta}$, we obtain the following estimate:

$$
\begin{equation*}
\left\|h_{m}\right\|_{\mathcal{C}_{\beta}^{2, \mu}\left(I_{3}\right)} \leq C\left(\log \frac{1}{\delta^{2}}\right)^{4+\beta}\|g\|_{\mathcal{C}_{\beta+1}^{0, \mu}\left(I_{1} \cup I_{2} \cup I_{3}\right)} . \tag{3.73}
\end{equation*}
$$

## Outer Problem for the Operator $\mathfrak{L}_{\boldsymbol{\delta}}$

Now we will find a solution $\hat{h}_{o}$ of (3.45) such that

$$
\begin{gather*}
2 \widehat{h}_{o}^{\prime \prime}+\widehat{p}_{\delta} \hat{h}_{o}=\hat{g}, \quad s>s_{\delta}, \\
\hat{h}_{o}\left(s_{\delta}\right)=\hat{h}_{m}\left(s_{\delta}\right), \quad \hat{h}_{o}^{\prime}\left(s_{\delta}\right)=\widehat{h}_{m}^{\prime}\left(s_{\delta}\right) . \tag{3.74}
\end{gather*}
$$

It is convenient to change variables $s=s_{\delta}+t$ and regard at first this problem for $t \in \mathbb{R}_{+}$. We will use the same symbols for the functions involved. Again we will use the variation-of-parameters formula. To this end, we need to choose two linearly independent solutions of the homogeneous problem such that

$$
\hat{h}_{o, 1}(t) \rightarrow \infty \quad \text { and } \quad \hat{h}_{o, 2}(t) \rightarrow 0, \quad t \rightarrow \infty
$$

A fundamental set with these properties can be found (for instance, see [22]) given that $\hat{p}_{\delta}\left(s_{\delta}+t\right)=-\frac{1}{4}+o(1)$ as $t \rightarrow \infty$. Moreover, we can choose $\hat{h}_{o, j}$ in such a way that

$$
\begin{array}{ll}
\hat{h}_{o, 1}(0)=0, & \hat{h}_{o, 2}(0)=1 \\
\hat{h}_{o, 1}^{\prime}(0)=1, & \widehat{h}_{o, 2}^{\prime}(0)=-\eta \tag{3.75}
\end{array}
$$

where $\eta>0$ is bounded independently of $\delta$. Observe that the Wronskian of these functions is $W\left(\hat{h}_{o, 1}, \hat{h}_{o, 2}\right)(t)=-1$. Then we get

$$
\begin{align*}
\hat{h}_{o}\left(s_{\delta}+t\right)= & {\left[\eta \hat{h}_{o}\left(s_{\delta}\right)+\hat{h}_{o}^{\prime}\left(s_{\delta}\right)\right] \hat{h}_{o, 1}(t)+h_{o}\left(s_{\delta}\right) \hat{h}_{o, 2}(t) } \\
& +\hat{h}_{o, 1}(t) \int_{0}^{t} h_{o, 2}(\tau) \hat{g}\left(s_{\delta}+\tau\right) d \tau  \tag{3.76}\\
& -\hat{h}_{o, 2}(t) \int_{0}^{t} h_{o, 1}(\tau) \widehat{g}\left(s_{\delta}+\tau\right) d \tau .
\end{align*}
$$

Since $\hat{p}_{\delta}^{\prime}\left(s_{\delta}+t\right)<0$ and $\hat{p}_{\delta}^{\prime \prime}\left(s_{\delta}+t\right)>0$ for $t>0$, by the general theory for second-order linear ODEs (see, for instance, [22, chap. 9.2]), we get that for some
$c_{j}, C_{j}>0, j=1,2:$

$$
\begin{align*}
C_{1} \exp \left\{\int_{0}^{t}\left[-\hat{p}_{\delta}\left(s_{\delta}+\tau\right)\right]^{1 / 2} d \tau\right\} & \leq \hat{h}_{o, 1}(t) \\
& \leq C_{2} \exp \left\{\int_{0}^{t}\left[-\hat{p}_{\delta}\left(s_{\delta}+\tau\right)\right]^{1 / 2} d \tau\right\},  \tag{3.77}\\
c_{1} \exp \left\{-\int_{0}^{t}\left[-\hat{p}_{\delta}\left(s_{\delta}+\tau\right)\right]^{1 / 2} d \tau\right\} & \leq \hat{h}_{o, 2}(t) \\
& \leq c_{2} \exp \left\{-\int_{0}^{t}\left[-\hat{p}_{\delta}\left(s_{\delta}+\tau\right)\right]^{1 / 2} d \tau\right\} .
\end{align*}
$$

We note that for any $\alpha>0, \nu>0$, and $\delta$ sufficiently small, the functions

$$
\begin{align*}
& \left(s_{\delta}+t\right)^{\alpha}\left(\log \frac{s_{\delta}+t}{\delta^{2}}\right)^{v+1} \\
& \quad \times \exp \left\{\int_{0}^{t}\left(\left[-\hat{p}_{\delta}\left(s_{\delta}+\tau\right)\right]^{1 / 2}-\frac{1}{2} a\left(s_{\delta}+\tau\right)\right) d \tau\right\}  \tag{3.78}\\
& \left(s_{\delta}+t\right)^{\alpha}\left(\log \frac{s_{\delta}+t}{\delta^{2}}\right)^{v+1} \\
& \quad \times \exp \left\{\int_{0}^{t}\left(-\left[-\hat{p}_{\delta}\left(s_{\delta}+\tau\right)\right]^{1 / 2}-\frac{1}{2} a\left(s_{\delta}+\tau\right)\right) d \tau\right\}
\end{align*}
$$

are monotone decreasing for $t>0$, hence using that $s_{\delta}=\mathcal{O}\left(\log \frac{1}{\delta}\right)$ and denoting

$$
\omega_{\beta, v+1}\left(s_{\delta}+t\right)=\left(s_{\delta}+t\right)^{\beta}\left(\log \frac{s_{\delta}+t}{\delta^{2}}\right)^{v+1} \exp \left\{-\int_{1}^{s_{\delta}+t} a(\tau) d \tau\right\}
$$

we get by (3.73):

$$
\begin{align*}
& \omega_{\beta, v+1}\left(s_{\delta}+t\right)\left|\left[\eta \hat{h}_{o}\left(s_{\delta}\right)+\hat{h}_{o}^{\prime}\left(s_{\delta}\right)\right] \hat{h}_{o, 1}(t)+\hat{h}_{o}\left(s_{\delta}\right) \hat{h}_{o, 2}(t)\right| \\
& \quad \leq C\left(\log \frac{1}{\delta}\right)^{5+2 \beta+v}\|g\|_{\mathcal{C}_{\beta+1,0}^{0, \mu}\left(0, s_{\delta}\right)}  \tag{3.79}\\
& \quad \leq C\left(\log \frac{1}{\delta}\right)^{4+2 \beta}\|g\|_{\mathcal{C}_{\beta+1, v+1}^{0, \mu}\left(0, s_{\delta}\right)} .
\end{align*}
$$

On the other hand, for any $\beta>0, v>0$, and $\delta$ sufficiently small, the functions

$$
\begin{align*}
& \left(s_{\delta}+t\right)^{-\beta-1}\left(\log \frac{s_{\delta}+t}{\delta^{2}}\right)^{-v-1} \\
& \quad \times \exp \left\{\int_{0}^{t}\left(-\left[-\hat{p}_{\delta}\left(s_{\delta}+\tau\right)\right]^{1 / 2}+\frac{1}{2} a\left(s_{\delta}+\tau\right)\right) d \tau\right\},  \tag{3.80}\\
& \left(s_{\delta}+t\right)^{-\beta-1}\left(\log \frac{s_{\delta}+t}{\delta^{2}}\right)^{-\nu-1} \\
& \quad \times \exp \left\{\int_{0}^{t}\left(\left[-\hat{p}_{\delta}\left(s_{\delta}+\tau\right)\right]^{1 / 2}+\frac{1}{2} a\left(s_{\delta}+\tau\right)\right) d \tau\right\},
\end{align*}
$$

are monotone increasing for $t>0$. Then, assuming $\|g\|_{\mathcal{C}_{\beta+1, v+1}^{0, \mu}\left(\mathbb{R}_{+}\right)}<\infty$, we get that the functions

$$
\begin{aligned}
& y_{1}(t)=\hat{h}_{o, 1}(t) \int_{0}^{t} h_{o, 2}(\tau) \hat{g}\left(s_{\delta}+\tau\right) d \tau \\
& y_{2}(t)=\hat{h}_{o, 2}(t) \int_{0}^{t} h_{o, 1}(\tau) \hat{g}\left(s_{\delta}+\tau\right) d \tau
\end{aligned}
$$

satisfy

$$
\begin{equation*}
\omega_{\beta, v+1}\left(s_{\delta}+t\right)\left(\left|y_{1}(t)\right|+\left|y_{2}(t)\right|\right) \leq C\|g\|_{\mathcal{C}_{\beta+1, v+1}^{0, \mu}}\left(\mathbb{R}_{+}\right) . \tag{3.81}
\end{equation*}
$$

We recall that

$$
h_{o}\left(s_{\delta}+t\right)=\hat{h}_{o}\left(s_{\delta}+t\right) \exp \left\{-\frac{1}{2} \int_{1}^{s_{\delta}+t} a(\tau) d \tau\right\} .
$$

Thus, by the variation-of-parameters formula (3.76) and (3.79)-(3.81), it follows that

$$
\begin{equation*}
\left\|h_{o}\right\|_{\mathcal{C}_{\beta, v+1}^{0, \mu}\left(s_{\delta}, \infty\right)} \leq C\left(\log \frac{1}{\delta}\right)^{4+2 \beta}\|g\|_{\mathcal{C}_{\beta+1, v}^{0}}\left(\mathbb{R}_{+}\right) \tag{3.82}
\end{equation*}
$$

To estimate the Hölder norms of the derivatives, we write the equation for $h_{o}$ in the form

$$
\left(h_{o}^{\prime} \exp \left\{\int_{s^{*}}^{s} a(\tau) d \tau\right\}\right)^{\prime}=\exp \left\{\int_{s^{*}}^{s} a(\tau) d \tau\right\}\left(g-p_{\delta} h\right),
$$

where $s^{*}<s_{\delta}$ is large and fixed independently of $\delta$. Integrating this equation from $s^{*}$ to $s>s_{\delta}$ we get

$$
\begin{align*}
& \left|h_{o}^{\prime}(s) \exp \left\{\int_{s^{*}}^{s} a(\tau) d \tau\right\}\right| \\
& \quad \leq\left|h_{o}^{\prime}\left(s^{*}\right)\right|+\left|\int_{s^{*}}^{s} \exp \left\{\int_{s^{*}}^{\sigma} a(\tau) d \tau\right\}\left(g-p_{\delta} h\right) d \sigma\right|^{\leq} \quad  \tag{3.83}\\
& \quad \leq\left|h_{o}^{\prime}\left(s^{*}\right)\right|+C\left(\|g\|_{\mathcal{C}_{\beta+1, v}^{0}\left(\mathbb{R}_{+}\right)}+\left\|h_{o}\right\|_{\mathcal{C}_{\beta, \nu}^{0, \mu}\left(\mathbb{R}_{+}\right)}\right) \int_{s^{*}}^{s} \widetilde{\omega}_{\beta, \nu}(\sigma) d \sigma,
\end{align*}
$$

where

$$
\widetilde{\omega}_{\beta, v}(\sigma)=(2+\sigma)^{-\beta-1}\left(\log \frac{2+\sigma}{\delta^{2}}\right)^{-v} \exp \left\{\int_{s^{*}}^{\sigma} a(\tau) d \tau\right\} .
$$

When $s^{*}$ is taken sufficiently large, we have for $\sigma>s^{*}$

$$
\widetilde{\omega}_{\beta, v}(\sigma) \leq C(2+\sigma)^{-\beta-1}\left(\log \frac{2+\sigma}{\delta^{2}}\right)^{-v} \exp \left\{\int_{s^{*}}^{s} a(\tau) d \tau\right\} .
$$

Using this for $s \in\left(s^{*}, s^{*}+1\right)$ we find by (3.83) that

$$
(2+s)^{\beta+1}\left(\log \frac{2+s}{\delta^{2}}\right)^{v}\left|h_{o}^{\prime}(s)\right| \leq C\left(\log \frac{1}{\delta}\right)^{4+2 \beta}\|g\|_{\mathcal{C}_{\beta+1, v}^{0}}\left(\mathbb{R}_{+}\right)
$$

by the previous argument. Then we argue inductively considering intervals of the form $\left(s^{*}+k, s^{*}+k+1\right)$ to show that for $s \in\left(s^{*}+k, s^{*}+k+1\right)$ we have an analogous estimate. This gives at the end

$$
\begin{equation*}
\left\|h_{o}^{\prime}\right\|_{\mathcal{C}_{\beta+1, v}^{0, \mu}\left(\left(s_{\delta}, \infty\right)\right)} \leq C\left(\log \frac{1}{\delta}\right)^{4+2 \beta}\|g\|_{\mathcal{C}_{\beta+1, v+1}^{0, \mu}}\left(\mathbb{R}_{+}\right) \tag{3.84}
\end{equation*}
$$

Then we estimate $h_{o}^{\prime \prime}$ using the equation directly.
Now the solution of (3.41) can be written in the form

$$
h=h_{i} \chi_{I_{1}}+h_{n} \chi_{I_{2}}+h_{m} \chi_{I_{3}}+h_{o} \chi_{\left(s_{\delta}, \infty\right)},
$$

where $\chi_{I}$ is the characteristic function of the interval $I$. We conclude the proof of Lemma 3.6 by combining estimates (3.54), (3.67), (3.73), and (3.84). For future purposes we will denote the right inverse of $\mathfrak{L}_{\delta}$ by $\mathfrak{L}_{\delta}^{-1}$. According to the statement of Lemma 3.6 we have in particular

$$
\begin{align*}
& \left\|\mathfrak{L}_{\delta}^{-1}(g)\right\|_{\mathcal{C}_{\beta, \nu}^{0, \mu}\left(\mathbb{R}_{+}\right)}+\left\|\left(\mathfrak{L}_{\delta}^{-1}(g)\right)^{\prime}\right\|_{\mathcal{C}_{\beta+1, v}^{1, \mu}\left(\mathbb{R}_{+}\right)} \leq  \tag{3.85}\\
& C\left(\log \frac{1}{\delta^{2}}\right)^{4+2 \beta}\|g\|_{\mathcal{C}_{\beta+1, v+1}^{0, \mu}\left(\mathbb{R}_{+}\right)} .
\end{align*}
$$

## Conclusion of the Proof of Proposition 3.2

We will now use the theory of the previous two sections to solve (3.12)-(3.13) and thereby complete the proof of Proposition 3.2

Notice that the existence of the function $v_{\varepsilon}$ solving (3.13) has been established already. Thus we only need to consider 3.12 . We will use a fixed point argument for the nonlinear operator $\mathcal{S}_{\delta}$. Let $k>0$ be fixed and take the approximate solution $\mathrm{v}_{k}$; see Lemma 3.5. We define $\mathrm{u}_{k}(s)=\mathrm{v}_{k}(r(s))$. Then the result of Lemma 3.5 reads

$$
\left|\mathcal{S}_{\delta}\left[u_{k}\right](s)\right| \leq \frac{C}{\left(\log \frac{2+s}{\delta^{2}}\right)^{k}} \frac{1}{(1+s)^{(k+1) / 2}}
$$

We will look for a solution in the form $u=u_{k}+\phi$. We will write

$$
\mathcal{S}_{\delta}\left[u_{k}+\phi\right]=\mathfrak{L}_{\delta}[\phi]+\mathcal{S}_{\delta}\left[u_{k}\right]+\mathfrak{N}_{\delta}(\phi),
$$

where

$$
\begin{aligned}
\mathfrak{N}_{\delta}(\phi) & =-\frac{1}{\delta^{2}} e^{-u_{k}}\left(e^{-\phi}-1+\phi\right) \\
& =-b(s) u_{0}\left[1+\mathcal{O}\left(\frac{1}{\log \frac{2+s}{\delta^{2}}}\right)\right]\left(e^{-\phi}-1+\phi\right) \\
& \sim \frac{1}{2+s} \log \left(\frac{2+s}{\delta^{2}}\right)\left(e^{-\phi}-1+\phi\right),
\end{aligned}
$$

is a nonlinear function with quadratic growth in its argument. Thus, we need to solve

$$
\mathfrak{L}_{\delta}[\phi]+\mathcal{S}_{\delta}\left[u_{k}\right]+\mathfrak{N}_{\delta}(\phi)=h_{\delta} .
$$

Now given the right inverse of $\mathfrak{L}_{\delta}$, we can put the above equation in the form of a fixed point problem for

$$
\mathcal{T}_{\delta}[\phi]:=-\mathfrak{L}_{\delta}^{-1}\left[\mathcal{S}_{\delta}\left[\mathrm{u}_{k}\right]+\mathfrak{N}_{\delta}(\phi)-h_{\delta}\right] .
$$

Given the result of Lemma 3.6 and (3.85) the existence of $\phi$ can be established. To see this let us fix real numbers $\beta, \nu, \gamma>0$ and a positive integer $k$, which satisfy in addition

$$
\frac{1}{2}>\beta, \quad v>6+2 \beta+\gamma, \quad k>4+2 \beta+v+\gamma
$$

With this choice one can verify that

$$
\begin{align*}
\left(\log \frac{1}{\delta^{2}}\right)^{4+2 \beta}\left\|\mathcal{S}_{\delta}\left[u_{k}\right]\right\|_{\mathcal{C}_{\beta+1, v+1}^{0, \mu}}\left(\mathbb{R}_{+}\right) & \leq C\left(\log \frac{1}{\delta^{2}}\right)^{-\gamma}, \\
\left(\log \frac{1}{\delta^{2}}\right)^{4+2 \beta}\left\|\mathfrak{N}_{\delta}(\phi)\right\|_{\mathcal{C}_{\beta+1, v+1}^{0, \mu}}\left(\mathbb{R}_{+}\right) & \leq C\left(\log \frac{1}{\delta^{2}}\right)^{-\gamma}\|\phi\|_{\mathcal{C}_{\beta, v}^{2, \mu}\left(\mathbb{R}_{+}\right)}^{2}  \tag{3.86}\\
\left(\log \frac{1}{\delta^{2}}\right)^{4+2 \beta}\left\|h_{\delta}\right\|_{\mathcal{C}_{\beta+1, v+1}^{0, \mu}\left(\mathbb{R}_{+}\right)} & \leq C\left(\log \frac{1}{\delta^{2}}\right)^{-\gamma} .
\end{align*}
$$

Then we see that, for each sufficiently small $\delta$, the map $\mathcal{T}_{\delta}$ takes the set

$$
\left\{\phi \left\lvert\,\|\phi\|_{\mathcal{C}_{\beta, \nu}^{0, \mu}\left(\mathbb{R}_{+}\right)}+\left\|\phi^{\prime}\right\|_{\mathcal{C}_{\beta+1, \nu}^{1, \mu}\left(\mathbb{R}_{+}\right)}<\left(\log \frac{1}{\delta^{2}}\right)^{-\frac{1}{2} \gamma}\right.\right\}
$$

into itself. Also, one can verify in a similar manner that this map is a Lipschitz contraction on this set and thus the proof of Proposition 3.2 follows.

## 4 Setting Up the Infinite-Dimensional Reduction

### 4.1 Construction of the Approximation

Let $\Gamma$ be the eternal solution of the mean curvature flow with $c=1$, and let $\Gamma_{\varepsilon}$ be the corresponding surface translating with speed $c=\varepsilon \ll 1$. We will use the natural representation of $\Gamma$ as a graph of the radial function $x_{N+1}=F(r)$. The scaled surface is given by $\Gamma_{\varepsilon}=\left\{x_{N+1}=F_{\varepsilon}(r) \mid F_{\varepsilon}(r)=\varepsilon^{-1} F(\varepsilon r)\right\}$. In general, we will take advantage of the radial symmetry of the eternal solution and employ the infinite-dimensional Lyapunov-Schmidt reduction method to reduce the original PDE,

$$
\begin{equation*}
\Delta u+\varepsilon \partial_{x_{N+1}} u+u-u^{3}=0 \quad \text { in } \mathbb{R}^{N+1} \tag{4.1}
\end{equation*}
$$

to a one-dimensional system of two equations whose independent variable is the radial variable $r$. This will be in fact the Jacobi-Toda system treated above.

We will now proceed to define an approximation of a solution of (4.1) that depends on the radial variable $r$ and the signed distance $z$ to $\Gamma_{\varepsilon}$. We will use the notation introduced in Sections $2.1-2.2$, with obvious modifications taking into account the fact that $\Gamma_{\varepsilon}$ is radially symmetric and thus has a globally defined parametrization.

## A Model for the Multicomponent Traveling Wave near $\Gamma_{\boldsymbol{\varepsilon}}$

In what follows, it will be useful to keep in mind that a global system of coordinates on $\Gamma$ and $\Gamma_{\varepsilon}$ can be defined by

$$
\begin{aligned}
\Gamma & =\left\{(r \Theta, F(r)) \mid r>0, \Theta \in S^{N-1}\right\} \\
\Gamma_{\varepsilon} & =\left\{\left.\left(r \Theta, \frac{1}{\varepsilon} F(\varepsilon r)\right) \right\rvert\, r>0, \Theta \in S^{N-1}\right\}
\end{aligned}
$$

There are other ways to introduce local coordinates on $\Gamma$. For instance, around each point $y \in \Gamma$ we have the normal geodesic coordinates. It is not hard to show that there exists $\delta_{0}>0$ such that these coordinates are well defined for each $y \in \Gamma$, at least in a neighborhood of $y$ of the form $U_{y, \delta_{0}}=B\left(y, \delta_{0}\right) \cap \Gamma$. A similar statement can be made when $y \in \Gamma_{\varepsilon}$ are considered, now with $U_{y, \delta_{0} / \varepsilon}=$ $B\left(y, \delta_{0} / \varepsilon\right) \cap \Gamma$.

We choose an orientation

$$
v(y)=\frac{(-\nabla F(r(y)), 1)}{\sqrt{1+|\nabla F(r(y))|^{2}}}
$$

on $\Gamma$ and take $z=z(x)=\operatorname{dist}(x, \Gamma)$ compatible with this orientation. Let us introduce the following weight functions:

$$
\omega(x)=2+\left|F_{r}(r)\right|^{2}, \quad \omega_{\varepsilon}(x)=2+\left|F_{r}(\varepsilon r)\right|^{2}, \quad x=\left(x^{\prime}, x_{N+1}\right), r=\left|x^{\prime}\right|
$$

We recall here that $F_{r}(r) \sim r, r \gg 1$. Also in what follows we will write $\omega(r)$, $\omega_{\varepsilon}(r)$, understanding that $r=r(x)=\left|x^{\prime}\right|$.

It is not hard to show that there exists an $\eta_{0}>0$ such that for all points $x$ such that $|z(x)| \leq \eta_{0} \log \omega(r)$ the map

$$
x \mapsto y+z v(y), \quad y \in \Gamma
$$

is a diffeomorphism. We denote this diffeomorphism by $X(x)=(y, z)$ and for a function $u$ given in a neighborhood of $\Gamma$ we set $\left(X^{*} u\right)(y, z)=\left(u \circ X^{-1}\right)(y, z)$. The coordinates $(y, z)$ above are called Fermi coordinates of $\Gamma$. Similar claims are true when we consider $\Gamma_{\varepsilon}$ and points $x$ such that $|z(x)| \leq \frac{\eta_{0}}{\varepsilon} \log \left(\omega_{\varepsilon}(r)\right)$. Taking this into account we introduce the following neighborhood of $\Gamma_{\varepsilon}$ :

$$
U_{\Gamma_{\varepsilon}}(M)=\left\{x \in \mathbb{R}^{N+1}| | z(x)\left|=\left|\operatorname{dist}\left(x, \Gamma_{\varepsilon}\right)\right| \leq M \log \left(\frac{\omega_{\varepsilon}(r)}{\varepsilon^{2}}\right)\right\}\right.
$$

Clearly Fermi coordinates are well defined in $U_{\Gamma_{\varepsilon}}(M)$ for all $M>0$ large and $\varepsilon>0$ small. If by $X_{\varepsilon}$ we denote the diffeomorphism in $U_{\Gamma_{\varepsilon}}(M)$ defined by $X_{\varepsilon}(x)=(y, z)$, then for a function $u$ defined in this neighborhood we set

$$
\left(X_{\varepsilon}^{*} u\right)(y, z)=\left(u \circ X_{\varepsilon}^{-1}\right)(y, z)
$$

We will describe functions $f_{j}$ representing the leading order for the location of the nodal set of our traveling wave. To this end we appeal to the results of Proposition 3.2 and let the functions $f_{j}, j=1,2$, be solutions of the Jacobi-Toda system (3.9) with $h_{j} \equiv 0$. We get that functions $f_{j}$ satisfy

$$
\begin{equation*}
f_{j}(r)=\frac{(-1)^{j}}{2 \sqrt{2}} \log \frac{2 \sqrt{2}}{\varepsilon^{2} \alpha_{0}\left|A_{\Gamma}(r)\right|^{2}}+\mathcal{O}\left(\log \log \frac{1}{\varepsilon^{2}\left|A_{\Gamma}(r)\right|^{2}}\right) \tag{4.2}
\end{equation*}
$$

In addition, we have $f_{1}=-f_{2}$.
In what follows we will use scaled versions of these functions, namely $f_{\varepsilon, j}$ : $\Gamma_{\varepsilon} \rightarrow \mathbb{R}$, defined by

$$
f_{\varepsilon, j}(r)=f_{j}(\varepsilon r), \quad r=r(y)=\left|y^{\prime}\right|, \quad y=\left(y^{\prime}, y_{N+1}\right) \in \Gamma_{\varepsilon}
$$

We recall here that $\varepsilon^{2}\left|A_{\Gamma}(\varepsilon r)\right|^{2}=\left|A_{\Gamma_{\varepsilon}}(r)\right|^{2}$.
In reality functions $f_{\varepsilon, j}$ give only the leading-order behavior of the traveling fronts and thus we further need two functions, which will be for a moment unknown parameters to be determined in the course of the Lyapunov-Schmidt scheme we use.

Thus we let $h_{j}, j=1,2$, be functions of the radial variable $r$ on $\Gamma$ such that for some $\beta, \tau \in(0,1)$ we have

$$
\begin{equation*}
\left\|h_{j}\right\|_{\mathcal{C}_{\beta}^{2, \mu}(\Gamma)} \leq \varepsilon^{\tau} \tag{4.3}
\end{equation*}
$$

As before we introduce scaled versions of these functions $h_{\varepsilon, j}: \Gamma_{\varepsilon} \rightarrow \mathbb{R}$ defined by $h_{\varepsilon, j}(r)=h_{j}(\varepsilon r)$. Let us make an elementary observation about the relation between the weighted norms on $\Gamma$ and $\Gamma_{\varepsilon}$. Defining the $\mathcal{C}_{\beta}^{2, \mu}\left(\Gamma_{\varepsilon}\right)$ norm in a natural way, namely using the weight function $\omega_{\varepsilon}^{\beta}(r)=\omega^{\beta}(\varepsilon r)$ and letting $h_{\varepsilon}(y)=$ $h(\varepsilon y)$, for $y \in \Gamma_{\varepsilon}$ we get

$$
\left\|h_{\varepsilon}\right\|_{\mathcal{C}_{\beta}^{2, \mu}\left(\Gamma_{\varepsilon}\right)} \leq\|h\|_{\mathcal{C}_{\beta}^{2, \mu}(\Gamma)} \leq \varepsilon^{-2-\mu}\left\|h_{\varepsilon}\right\|_{\mathcal{C}_{\beta}^{2, \mu}\left(\Gamma_{\varepsilon}\right)} .
$$

In particular, we get from this and (4.3)

$$
\begin{equation*}
\left\|h_{\varepsilon, j}\right\|_{\mathcal{C}_{\beta}^{2, \mu}\left(\Gamma_{\varepsilon}\right)} \leq \varepsilon^{\tau}, \quad j=1,2 \tag{4.4}
\end{equation*}
$$

Given the functions $f_{\varepsilon, j}$ and $h_{\varepsilon, j}$ as described above we will denote

$$
\mathbf{f}_{\varepsilon}=\left(f_{\varepsilon, 1}, f_{\varepsilon, 2}\right), \quad \mathbf{h}_{\varepsilon}=\left(h_{\varepsilon, 1}, h_{\varepsilon, 2}\right)
$$

etc.
To define a model for the traveling profile we first recall that by $H$ we have denoted the unique, odd, and monotonically increasing solution of $H^{\prime \prime}+H(1-$ $\left.H^{2}\right)=0$. Next we consider a cutoff function

$$
\chi(t)= \begin{cases}0, & |t|<1 \\ 1, & |t|>2\end{cases}
$$

Now, let $M>0$ be a fixed large number and let

$$
\begin{equation*}
\chi_{\varepsilon}(x)=\chi\left(\frac{z(x)}{M \log \left(\frac{\omega_{\varepsilon}(r)}{\varepsilon^{2}}\right)}\right), \quad z(x)=\operatorname{dist}\left(x, \Gamma_{\varepsilon}\right) \tag{4.5}
\end{equation*}
$$

Taking $M$ large and $\varepsilon$ small we define the initial approximation of the solution in the support of $\chi_{\varepsilon}$ by
(4.6) $\left(X_{\varepsilon}^{*} u_{\varepsilon}\right)(r, z)=H\left(z-f_{\varepsilon, 1}(r)-h_{\varepsilon, 1}(r)\right)-H\left(z-f_{\varepsilon, 2}(r)-h_{\varepsilon, 2}(r)\right)-1$.

Next we define the initial approximation globally in $\mathbb{R}^{N+1}$ by

$$
\begin{equation*}
w_{\varepsilon}(x)=\left(1-\chi_{\varepsilon}(x)\right) u_{\varepsilon}(x)-\chi_{\varepsilon}(x) \tag{4.7}
\end{equation*}
$$

### 4.2 Reduction to the Nonlinear Projected Problem

We look for a solution of

$$
S(u)=\Delta u+\varepsilon \partial_{x_{N+1}} u+u\left(1-u^{2}\right)=0
$$

in the form $u=w_{\varepsilon}+\varphi_{\varepsilon}$, where $\varphi_{\varepsilon}$ is a small function. We write

$$
S\left(w_{\varepsilon}+\varphi_{\varepsilon}\right)=S\left(w_{\varepsilon}\right)+L \varphi_{\varepsilon}+N\left(\varphi_{\varepsilon}\right)
$$

where

$$
\begin{aligned}
L \varphi_{\varepsilon} & =\Delta \varphi_{\varepsilon}+\varepsilon \partial_{x_{N+1}} \varphi_{\varepsilon}+\left(1-3 w_{\varepsilon}^{2}\right) \varphi_{\varepsilon}, \\
N\left(\varphi_{\varepsilon}\right) & =-3 w_{\varepsilon} \varphi_{\varepsilon}^{2}-\varphi_{\varepsilon}^{3}
\end{aligned}
$$

We will decompose our nonlinear problem into a system suitable to apply an infinite-dimensional Lyapunov-Schmidt reduction scheme. To this end we recall that we have given functions $\mathbf{f}_{\varepsilon}$ and also unknown functions $\mathbf{h}_{\varepsilon}$.

Given a large number $M$ as in the definition of $w_{\varepsilon}$ above we consider smooth cutoff functions $\zeta_{j} \geq 0, j=1,2$, which satisfy the following conditions:

$$
\zeta_{1}(t)+\zeta_{2}(t)=\left\{\begin{array}{ll}
1, & |t| \leq M,  \tag{4.8}\\
0, & |t| \geq 2 M,
\end{array} \quad \zeta_{1}(t)= \begin{cases}1, & -M<t<-1, \\
0, & t>1 .\end{cases}\right.
$$

We define cutoff functions $\zeta_{\varepsilon, j}$ by

$$
\begin{equation*}
\left(X_{\varepsilon}^{*} \zeta_{\varepsilon, j}\right)(r, z)=\zeta_{j}\left(z-\left(\frac{1}{2}+\delta\right)\left|f_{\varepsilon, 1}(r)-f_{\varepsilon, 2}(r)\right|\right) \tag{4.9}
\end{equation*}
$$

where $\delta$ is a small constant. Note that with this definition we have

$$
\begin{aligned}
& \zeta_{\varepsilon, 1}+\zeta_{\varepsilon, 2}=1, \quad|z|<M+\left(\frac{1}{2}+\delta\right)\left|f_{\varepsilon, 1}(r)-f_{\varepsilon, 2}(r)\right| \\
& \zeta_{\varepsilon, 1}+\zeta_{\varepsilon, 2}=0, \quad|z|>2 M+\left(\frac{1}{2}+\delta\right)\left|f_{\varepsilon, 1}(r)-f_{\varepsilon, 2}(r)\right|
\end{aligned}
$$

Also we have

$$
\zeta_{\varepsilon, j}\left(r,\left(f_{\varepsilon, j}+h_{\varepsilon, j}\right)(\varepsilon r)\right)=1
$$

Furthermore, we choose cutoff functions $\tilde{\zeta}_{\varepsilon, j}$ such that

$$
\begin{aligned}
& \operatorname{supp} \tilde{\zeta}_{\varepsilon, 1}=\left\{-3 M-\frac{1}{2}\left|f_{\varepsilon, 1}(r)-f_{\varepsilon, 2}(r)\right|<z<\left(\frac{1}{2}+2 \delta\right)\left|f_{\varepsilon, 1}(r)-f_{\varepsilon, 2}(r)\right|\right\}, \\
& \operatorname{supp} \tilde{\zeta}_{\varepsilon, 2}=\left\{3 M+\frac{1}{2}\left|f_{\varepsilon, 1}(r)-f_{\varepsilon, 2}(r)\right|>z>-\left(\frac{1}{2}+2 \delta\right)\left|f_{\varepsilon, 1}(r)-f_{\varepsilon, 2}(r)\right|\right\},
\end{aligned}
$$

and additionally

$$
\tilde{\zeta}_{\varepsilon, j} \zeta_{\varepsilon, j}=\zeta_{\varepsilon, j}
$$

Now we look for a solution of our problem $\varphi_{\varepsilon}$ in the form

$$
\varphi_{\varepsilon}=\sum_{j=1,2} \zeta_{\varepsilon, j} \phi_{\varepsilon, j}+\psi_{\varepsilon}
$$

The functions $\phi_{\varepsilon, j}$ and $\psi_{\varepsilon}$ must still be determined from a system of equations that we will now describe. First we introduce functions $H_{\varepsilon, j}^{\prime}$ defined by

$$
\left(X_{\varepsilon}^{*} H_{\varepsilon, j}^{\prime}\right)(y, z)=H^{\prime}\left(z-f_{\varepsilon, j}(\varepsilon r)\right), \quad r=\left|y^{\prime}\right| .
$$

We also introduce new unknowns $\mathrm{c}_{\varepsilon, j}, j=1,2$, which are functions on $\Gamma_{\varepsilon}$. Next, we ask that the functions $\phi_{\varepsilon, j}, \psi_{\varepsilon}$, and $\mathrm{c}_{\varepsilon, j}$ be solutions of the following coupled system of equations:

$$
\begin{align*}
\tilde{\zeta}_{\varepsilon, j} L \phi_{\varepsilon, j}=\tilde{\zeta}_{\varepsilon, j} & \left\{-\left(S\left(w_{\varepsilon}\right)+N\right)-\left(L-\Delta-\varepsilon \partial_{x_{N+1}}+2\right) \psi_{\varepsilon}\right.  \tag{4.10}\\
& \left.-\left[L, \zeta_{\varepsilon, j}\right] \phi_{\varepsilon, j}+c_{\varepsilon, j} H_{\varepsilon, j}^{\prime}\right\},
\end{align*}
$$

$$
\begin{align*}
\left(\Delta+\varepsilon \partial_{x_{N+1}}-2\right) \psi_{\varepsilon}= & -\left(1-\sum_{i=1,2} \zeta_{i, \varepsilon}\right)\left\{S\left(w_{\varepsilon}\right)+N+\left[L, \zeta_{\varepsilon, i}\right] \phi_{\varepsilon, i}\right\} \\
& -\left(1-\sum_{i=1,2} \zeta_{\varepsilon, i}\right)\left(L-\Delta-\varepsilon \partial_{X_{N+1}}+2\right) \psi_{\varepsilon} \tag{4.11}
\end{align*}
$$

where $N=N\left(\sum_{j=1,2} \phi_{\varepsilon, j} \zeta_{\varepsilon j}+\psi_{\varepsilon}\right)$. Note that after multiplying (4.10) by $\zeta_{\varepsilon, j}$, $j=1,2$, using the fact that $\zeta_{\varepsilon, j} \tilde{\zeta}_{\varepsilon, j} \equiv 1$, and adding the resulting expression and (4.11), we obtain

$$
\begin{equation*}
L \varphi_{\varepsilon}+S\left(w_{\varepsilon}\right)+N\left(\varphi_{\varepsilon}\right)=\sum_{j=1,2} \mathrm{c}_{\varepsilon, j} H_{\varepsilon, j}^{\prime} \zeta_{\varepsilon, j} . \tag{4.12}
\end{equation*}
$$

As is usual in a Lyapunov-Schmidt reduction approach, the functions $\mathrm{c}_{\varepsilon, j}$ will be initially determined in such a way that (4.10) has a solution for any given parameter function $\mathbf{h}_{\varepsilon}$. Later we will adjust the traveling front, whose location is represented by $\mathbf{f}_{\varepsilon}+\mathbf{h}_{\varepsilon}$, so that $\mathrm{c}_{\varepsilon}, j \equiv 0$. After this is done, we will get the solution to our original problem.

In fact, a slight modification of (4.10), which we will describe now, is needed. We introduce the following functions:

$$
\left(X_{\varepsilon}^{*} w_{\varepsilon, j}\right)(y, z)=H\left(z-f_{\varepsilon, j}(\varepsilon r)\right), \quad j=1,2, r=\left|y^{\prime}\right|,
$$

and check that we have, say in the set $\tilde{\zeta}_{\varepsilon, j} \equiv 1$,

$$
\begin{aligned}
L \phi_{\varepsilon, j}= & \Delta_{\Gamma_{\varepsilon}}+\partial_{z}^{2} \phi_{\varepsilon, j}+f^{\prime}\left(w_{\varepsilon, j}\right) \phi_{\varepsilon, j} \\
& +\left[f^{\prime}\left(w_{\varepsilon}\right)-f^{\prime}\left(w_{\varepsilon, j}\right)\right] \phi_{\varepsilon, j}+\left[\Delta_{\Gamma_{\varepsilon, z}}-\Delta_{\Gamma_{\varepsilon}}\right] \phi_{\varepsilon, j} \\
& -\left(H_{\Gamma_{\varepsilon, z}}-\varepsilon v_{\Gamma_{\varepsilon}, N+1}\right) \partial_{z} \phi_{\varepsilon, j}+\varepsilon \nabla_{\Gamma_{\varepsilon, z}}\left(\pi_{\varepsilon, N+1}\right) \cdot \nabla_{\Gamma_{\varepsilon, z}} \phi_{\varepsilon, j} .
\end{aligned}
$$

Then, we can write (4.10) in the form

$$
\begin{equation*}
\Delta_{\Gamma_{\varepsilon}} \phi_{\varepsilon, j}+\partial_{z}^{2} \phi_{\varepsilon, j}+f^{\prime}\left(w_{\varepsilon, j}\right) \phi_{\varepsilon, j}=\mathfrak{g}_{\varepsilon, j}+c_{\varepsilon, j} H_{\varepsilon, j}^{\prime} \tag{4.13}
\end{equation*}
$$

at least when $\tilde{\zeta}_{\varepsilon, j} \equiv 1$. However, it is convenient to view this problem in the set $\Gamma_{\varepsilon} \times \mathbb{R}$. Indeed, the operator $L_{\varepsilon, j}=\Delta_{\Gamma_{\varepsilon}}+\partial_{z}^{2}+f^{\prime}\left(w_{\varepsilon, j}\right)$ is defined on functions whose domain is $\Gamma_{\varepsilon} \times \mathbb{R}$, while the right-hand side is a function supported on a set $\operatorname{supp} \tilde{\zeta}_{\varepsilon, j}$. More precisely, we have

$$
\begin{align*}
\mathfrak{g}_{\varepsilon, j}= & \widetilde{\zeta}_{\varepsilon, j}\left(S\left(w_{\varepsilon}\right)+N\right)-\widetilde{\zeta}_{\varepsilon, j}\left(L-\Delta-\varepsilon \partial_{x_{N+1}}+2\right) \psi_{\varepsilon} \\
& -\widetilde{\zeta}_{\varepsilon, j}\left[L, \zeta_{\varepsilon, j}\right] \phi_{\varepsilon, j}+\widetilde{\zeta}_{\varepsilon, j}\left[f^{\prime}\left(w_{\varepsilon}\right)-f^{\prime}\left(w_{\varepsilon, j}\right)\right] \phi_{\varepsilon, j}  \tag{4.14}\\
& +\widetilde{\zeta}_{j, \varepsilon}\left[\Delta_{\Gamma_{\varepsilon, z}}-\Delta_{\Gamma_{\varepsilon}}\right] \phi_{\varepsilon, j} \\
& +\widetilde{\zeta}_{\varepsilon, j}\left[\left(H_{\Gamma_{\varepsilon, z}}-\varepsilon v_{\Gamma_{\varepsilon}, N+1}\right) \partial_{z} \phi_{\varepsilon, j}-\varepsilon \nabla_{\Gamma_{\varepsilon, z}}\left(\pi_{\varepsilon, N+1}\right) \cdot \nabla_{\Gamma_{\varepsilon, z}} \phi_{\varepsilon, j}\right] .
\end{align*}
$$

Again, multiplying (4.13) by $\zeta_{\varepsilon, j}$ and adding the resulting equations and (4.11), we get (4.12).

For future purposes we write (4.11) in the form

$$
\begin{equation*}
\left(\Delta+\varepsilon \partial_{x_{N+1}}-2\right) \psi_{\varepsilon}=\mathfrak{h}_{\varepsilon}, \tag{4.15}
\end{equation*}
$$

where by $\mathfrak{h}_{\varepsilon}$ we have denoted the right-hand side of (4.11). Note that if we assume that $\phi_{\varepsilon, j}$ and $\psi_{\varepsilon}$ are functions of ( $r, x_{N+1}$ ) only with $r=\left|x^{\prime}\right|$, then so are the functions $\mathfrak{g}_{\varepsilon, j}$ and $\mathfrak{h}_{\varepsilon}$. Conversely, if we consider more generally problems of the form (4.13) and (4.15) with $\mathfrak{g}_{\varepsilon, j}$ and $\mathfrak{h}_{\varepsilon}$ depending on $\left(r, x_{N+1}\right)$ only, then the solutions of these problems $\phi_{\varepsilon, j}$ and $\psi_{\varepsilon}$ will also depend on ( $r, x_{N+1}$ ) only.

### 4.3 Further Modification of (4.13)

Let us look now at equation (4.13) more closely. We have in general the following system to solve:

$$
\left[\Delta_{\Gamma_{\varepsilon}}+\partial_{z}^{2}+f^{\prime}\left(w_{\varepsilon, j}\right)\right] \phi_{\varepsilon, j}=\mathfrak{g}_{\varepsilon, j} \quad \text { in } \Gamma_{\varepsilon} \times \mathbb{R}, j=1,2
$$

It is convenient to rewrite this system in the following way: first, we introduce shifted Fermi coordinates

$$
\mathrm{t}_{j}=z-f_{\varepsilon, j}(r), \quad j=1,2 .
$$

Second, we write each of the operators above in terms of these new coordinates:

$$
\begin{aligned}
\Delta_{\Gamma_{\varepsilon}}+\partial_{z}^{2}+f^{\prime}\left(w_{\varepsilon, j}\right)= & \Delta_{\Gamma_{\varepsilon}}+\partial_{\mathrm{t}_{j}}^{2}+f^{\prime}\left(H\left(\mathrm{t}_{j}\right)\right)-\Delta_{\Gamma_{\varepsilon}} f_{\varepsilon, j} \partial_{\mathrm{t}_{j}} \\
& -\nabla_{\Gamma_{\varepsilon}} f_{\varepsilon, j} \cdot \nabla_{\Gamma_{\varepsilon}} \partial_{\mathrm{t}_{j}}+\left|\nabla_{\Gamma_{\varepsilon}} f_{\varepsilon, j}\right|^{2} \partial_{\mathrm{t}_{j}}^{2} .
\end{aligned}
$$

Usually the second line above is relatively small in the sense that its norm can be controlled by the norm of the solution times a small factor, and thus we can absorb it on the right-hand side of the corresponding equation. Note also that variables $t_{j}$ are related through the formula

$$
\begin{equation*}
\mathrm{t}_{1}-\mathrm{t}_{2}=f_{\varepsilon, 2}-f_{\varepsilon, 1} . \tag{4.16}
\end{equation*}
$$

Then letting

$$
\begin{aligned}
\tilde{\mathfrak{g}}_{\varepsilon, j}\left(y, \mathrm{t}_{j}\right)=\mathfrak{g}_{\varepsilon, j}+\tilde{\zeta}_{\varepsilon, j}[ & \Delta_{\Gamma_{\varepsilon}} f_{\varepsilon, j} \partial_{\mathrm{t}_{j}}+\nabla_{\Gamma_{\varepsilon}} f_{\varepsilon, j} \cdot \nabla_{\Gamma_{\varepsilon}} \partial_{\mathrm{t}_{j}} \\
& \left.-\left|\nabla_{\Gamma_{\varepsilon}} f_{\varepsilon, j}\right|^{2} \partial_{\mathrm{t}_{j}}^{2}\right] \phi_{\varepsilon, j},
\end{aligned}
$$

we obtain the following system:

$$
\begin{equation*}
\left[\Delta_{\Gamma_{\varepsilon}}+\partial_{\mathrm{t}_{j}}^{2}+f^{\prime}\left(H\left(\mathrm{t}_{j}\right)\right)\right] \phi_{\varepsilon, j}=\tilde{\mathfrak{g}}_{\varepsilon, j}\left(y, \mathrm{t}_{j}\right)+\mathrm{c}_{\varepsilon, j} H^{\prime}\left(\mathrm{t}_{j}\right), \quad j=1,2, \tag{4.17}
\end{equation*}
$$

where now, with some abuse of notation, $\phi_{\varepsilon, j}=\phi_{\varepsilon, j}\left(y, \mathrm{t}_{j}\right)$. This system can be considered as a system for functions defined on two copies $\Gamma_{\varepsilon} \times \mathbb{R}$, and it looks at first sight as being decoupled. However, in reality we have, in the original setting,

$$
\tilde{\mathfrak{g}}_{\varepsilon, j}=\tilde{\mathfrak{g}}_{\varepsilon, j}\left(y, z ; \phi_{\varepsilon, 1}, \phi_{\varepsilon, 2}, \psi_{\varepsilon}\right) .
$$

Therefore when considering for instance the equation for $\phi_{\varepsilon, 1}$ in the shifted variable $t_{1}$, we need to use the above relation between $t_{1}$ and $t_{2}$ to express all functions involved in terms of $y \in \Gamma_{\varepsilon}$ and $t_{1} \in \mathbb{R}$. Of course, the same must be done with the second equation. As a result, we will obtain a nonlinear and nonlocal system for $\phi_{\varepsilon, j}, j=1,2$. The advantage of making this transformation is that we always work with the same, basic linearized operator on the left-hand side. Again,
we point out that all the functions involved depend on $y$ through the radial variable $r=\left|y^{\prime}\right|$.

## 5 Linear Theory

We recall that we have denoted $\omega(r(y))=1+|\nabla F(r(y))|^{2}$ and $\omega_{\varepsilon}(r)=$ $\omega(\varepsilon r)$. Given a $\mathcal{C}^{2, \mu}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)$ function $u$ we define its weighted norms by

$$
\begin{align*}
&\|u\|_{\mathcal{C}_{\beta, \eta}^{0, \mu}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)}= \sup _{(y, z) \in \Gamma_{\varepsilon} \times \mathbb{R}}(\cosh z)^{\eta} \omega_{\varepsilon}^{\beta}(r(y))\|u\|_{\mathcal{C}^{0, \mu}\left(B(y, 1) \cap \Gamma_{\varepsilon} \times(z-1, z+1)\right)} \\
&\|u\|_{\mathcal{C}_{\beta, \eta}^{2, \mu}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)}=\|u\|_{\mathcal{C}_{\beta, \eta}^{0, \mu}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)}+\left\|\nabla_{\Gamma_{\varepsilon} \times \mathbb{R}} u\right\|_{\mathcal{C}_{\beta, \eta}^{0, \mu}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)}  \tag{5.1}\\
&+\left\|D_{\Gamma_{\varepsilon} \times \mathbb{R}}^{2} u\right\|_{\mathcal{C}_{\beta, \eta}^{0, \mu}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)}
\end{align*}
$$

Above $\nabla_{\Gamma_{\varepsilon} \times \mathbb{R}}$ and $D_{\Gamma_{\varepsilon} \times \mathbb{R}}^{2}$ denote the gradient and second derivative on the manifold $\Gamma_{\varepsilon} \times \mathbb{R}$ equipped with a natural product metric and the associated Levi-Civita connection.

In this section we will consider the following basic linearized operator:

$$
\Delta_{\Gamma_{\varepsilon}} \phi+\partial_{z}^{2} \phi+f^{\prime}(H(z)) \phi \equiv \mathrm{L}_{\varepsilon} \phi
$$

We note that

$$
\partial_{z}^{2} H^{\prime}+f^{\prime}(H) H^{\prime}=0
$$

In fact, $H^{\prime}$ is the unique bounded element in the kernel of $\partial_{z}^{2}+f^{\prime}(H)$. In particular, we have, with some $\nu_{0}>0$,

$$
\int_{\mathbb{R}}\left|\phi^{\prime}(z)\right|^{2}-f^{\prime}(H(z))|\phi(z)|^{2} \geq v_{0} \int_{\mathbb{R}}|\phi(z)|^{2}
$$

whenever $\phi$ satisfies

$$
\int_{\mathbb{R}} \phi(z) H^{\prime}(z) d z=0
$$

In general, we will consider the following problem:

$$
\begin{gather*}
\Delta_{\Gamma_{\varepsilon}} \phi+\partial_{z}^{2} \phi+f^{\prime}(H) \phi=\mathfrak{g} \\
\int_{-\infty}^{\infty} \phi(y, z) H^{\prime}(z) d z=0, \quad y \in \Gamma_{\varepsilon} \times \mathbb{R} \tag{5.2}
\end{gather*}
$$

We will assume that

$$
\|\mathfrak{g}\|_{\mathcal{C}_{\beta, \eta}^{0, \mu}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)} \leq \infty
$$

with some $\beta, \eta>0$. In the case at hand we have $\beta \in(0,1)$ and $\eta \in(0, \sqrt{2})$.

### 5.1 A Priori Estimates

Most of what will be said in this section follows the argument of [8], and so we will only outline the main points.

First we need the following:
Lemma 5.1. The only bounded solutions of

$$
\Delta \phi+\partial_{z}^{2} \phi+f^{\prime}(H(z)) \phi=0 \quad \text { in } \mathbb{R}^{N+1}, N \geq 0
$$

are of the form $\phi=c H^{\prime}(z)$ with some constant $c$.
This lemma is proven in [11] (see also [10]) .
Next, we show the following a priori estimate:
Lemma 5.2. Let $\phi$ be a solution of the problem (5.2). There holds

$$
\begin{equation*}
\|\phi\|_{\mathcal{C}_{\beta, \eta}^{2, \mu}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)} \leq C\|\mathfrak{g}\|_{\mathcal{C}_{\beta, \eta}^{0, \mu}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)} . \tag{5.3}
\end{equation*}
$$

Proof. The proof of this lemma follows arguments in [10, 11], with only small changes due to the fact that here we use slightly different norms.

We argue by contradiction. Thus we assume that there exist sequences $\left\{\varepsilon_{n}\right\}$, $\left\{\phi_{\varepsilon_{n}}\right\}$, and $\left\{\mathfrak{g}_{\varepsilon_{n}}\right\}$ such that

$$
\begin{aligned}
& \Delta_{\Gamma_{\varepsilon_{n}}} \phi_{\varepsilon_{n}}+\partial_{z}^{2} \phi_{\varepsilon_{n}}+f^{\prime}(H) \phi_{\varepsilon_{n}}=\mathfrak{g}_{\varepsilon_{n}} \quad \text { in } \Gamma_{\varepsilon_{n}} \times \mathbb{R}, \\
& \int_{-\infty}^{\infty} \phi_{\varepsilon_{n}}(y, z) H^{\prime}(z) d z=0, \quad y \in \Gamma_{\varepsilon_{n}},
\end{aligned}
$$

and such that as $\varepsilon_{n} \rightarrow 0$

$$
\left\|\phi_{\varepsilon_{n}}\right\|_{\mathcal{C}_{\beta, n}^{2, \mu}\left(\Gamma_{\varepsilon_{n}} \times \mathbb{R}\right)}=1, \quad\left\|\mathfrak{g}_{\varepsilon_{n}}\right\|_{\mathcal{C}_{\beta, \eta}^{0, \mu}\left(\Gamma_{\varepsilon_{n}} \times \mathbb{R}\right)} \rightarrow 0
$$

In particular, from the definition of the norm there exists $y_{n} \in \Gamma_{\varepsilon_{n}}$ and $z_{n} \in \mathbb{R}$ such that

$$
\begin{equation*}
\left(\cosh z_{n}\right)^{\eta} \omega_{\varepsilon_{n}}^{\beta}\left(r\left(y_{n}\right)\right)\left\|\phi_{\varepsilon_{n}}\right\|_{\mathcal{C}^{0, \mu}}\left(B\left(y_{n}, 1\right) \cap \Gamma_{\varepsilon_{n}} \times\left(z_{n}-1, z_{n}+1\right)\right)>\frac{1}{2} . \tag{5.4}
\end{equation*}
$$

We consider four cases depending on the behavior of the sequences $\left\{\varepsilon_{n} r\left(y_{n}\right)\right\}$ and $\left\{z_{n}\right\}$. The various possibilities are, for example, (1) $\varepsilon_{n} r\left(y_{n}\right)$ and $z_{n}$ bounded, (2) $\varepsilon_{n} r\left(y_{n}\right) \rightarrow \infty$ while $z_{n}$ bounded, etc. In each of these cases we use essentially the same argument with just slight modifications. This has been done in detail in [10, 11].

To get the idea of the general scheme, we assume for instance that $\left\{\varepsilon_{n} r\left(y_{n}\right)\right\}$ and $\left\{z_{n}\right\}$ are bounded. We take the normal geodesic coordinates on $\Gamma_{\varepsilon_{n}}$, which are defined around each $y_{n}$ at least in the set $U_{n}=B\left(y_{n}, \delta_{0} / \varepsilon_{n}\right) \cap \Gamma_{\varepsilon_{n}}$, where $\delta_{0}>0$ is a small number independent of $y_{n}$. We denote the coordinates of $y \in U_{n}$ by $\xi=\left(\xi_{1}, \ldots, \xi_{N}\right)$ and set

$$
\widetilde{\phi}_{n}(\xi, z)=\phi_{\varepsilon_{n}}(y, z), \quad(y, z) \in U_{n} \times \mathbb{R}
$$

In the local coordinates we have

$$
\begin{aligned}
\Delta_{\Gamma_{\varepsilon_{n}}} \phi_{\varepsilon_{n}}+\partial_{z}^{2} \phi_{\varepsilon_{n}}+f^{\prime}(H) \phi_{\varepsilon_{n}}= & \Delta \widetilde{\phi}_{n}+\partial_{z}^{2} \widetilde{\phi}_{n}+f^{\prime}(H) \widetilde{\phi}_{n} \\
& +a_{\varepsilon_{n}, i j} \partial_{i j} \widetilde{\phi}_{n}+b_{\varepsilon_{n}, j} \partial_{j} \widetilde{\phi}_{n}
\end{aligned}
$$

Passing to the limit over compacts we obtain that $\tilde{\phi}_{n} \rightarrow \tilde{\phi}$ in $\mathcal{C}_{\text {loc }}^{2, \mu^{\prime}}\left(\mathbb{R}^{N} \times \mathbb{R}\right)$, $\mu^{\prime}<\mu$, where $\tilde{\phi}(0)>0$ and $\tilde{\phi}$ is bounded, and at the same time

$$
\Delta \widetilde{\phi}+\partial_{z}^{2} \widetilde{\phi}+f^{\prime}(H) \widetilde{\phi}=0
$$

Lemma 5.1 implies that $\tilde{\phi}=c H^{\prime}$, but this contradicts the fact that we also have

$$
\int_{R} \tilde{\phi}(\cdot, z) H^{\prime}(z) d z=0
$$

passing to the limit in the orthogonality condition.
To get an idea of how the other cases are handled, let us consider the case $\varepsilon_{n} r\left(y_{n}\right) \rightarrow \infty$, while $\left\{z_{n}\right\}$ remains bounded. Then we proceed in a similar manner as above letting

$$
\widetilde{\phi}_{n}(\xi, z)=\omega_{\varepsilon}^{\beta}(r(y)) \phi_{\varepsilon_{n}}(y, z) .
$$

For the remaining cases we refer the reader to [10] (see also [11]).

### 5.2 An Existence Result for the Model Linear Problem

Proposition 5.3. Given $\mathfrak{g} \in \mathcal{C}_{\beta, \eta}^{0, \mu}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)$ such that $\int_{\mathbb{R}} \mathfrak{g}(\cdot, z) H^{\prime}(z) d z=0$, there exists a unique solution of (5.2).

Proof. We will argue by approximations. Let us replace $\mathfrak{g}$ in (5.2) by a function $\mathfrak{g}_{R}(y, z)=\mathfrak{g}(y, z) \chi_{(0, R)}(y)$ where we will take $R \rightarrow \infty$ later on. With this right-hand side we can give to the problem (5.2) a weak formulation in the closed subspace of the Sobolev space $H^{1}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)$ of functions that satisfy the orthogonality conditions in (5.2). Thus we have

$$
\begin{align*}
\Delta_{\Gamma_{\varepsilon}} \phi_{R}+\partial_{z}^{2} \phi_{R}+f^{\prime}(H(z)) \phi_{R} & =\mathfrak{g}_{R} \\
\int_{\mathbb{R}} \phi_{R}(y, z) H^{\prime}(z) d z & =0 \tag{5.5}
\end{align*}
$$

With this operator we associate the bilinear form

$$
\mathfrak{a}_{R}(\phi, \psi)=\int_{\Gamma_{\varepsilon} \times \mathbb{R}}\left[\nabla_{\Gamma_{\varepsilon}} \phi \cdot \nabla_{\Gamma_{\varepsilon}} \psi+\partial_{z} \phi \partial_{z} \psi-f^{\prime \prime}(H(z)) \phi \psi\right] d V\left(\Gamma_{\varepsilon}\right) d z
$$

Then we say that $\phi_{R}$ is a weak solution of this problem if for all tests functions $\psi$ we have

$$
\mathfrak{a}_{R}\left(\phi_{R}, \psi\right)=\int_{\Gamma_{\varepsilon} \cap B(0, R) \times \mathbb{R}} \mathfrak{g}_{R} \psi d V\left(\Gamma_{\varepsilon}\right) d z
$$

Since we have as well, by our assumption,

$$
\int_{\mathbb{R}} \mathfrak{g}_{R}(y, z) H^{\prime}(z) d z=0 \quad \forall y \in \Gamma_{\varepsilon},
$$

and, under the orthogonality conditions, the bilinear form $\mathfrak{a}_{R}(\psi, \eta)$ is actually positive definite, it follows that there exists a unique $\phi_{R} \in H^{1}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)$ that satisfies weakly the equation and the orthogonality condition.

Letting $R \rightarrow+\infty$ and using the uniform a priori estimates valid for the approximations completes the proof of the Proposition.

### 5.3 A Priori Estimates and Existence for (4.11)

In this section we will consider the following problem:

$$
\begin{equation*}
\left(\Delta+\partial_{x_{N+1}}^{2}+\varepsilon \partial_{x_{N+1}}-2\right) \psi=\mathfrak{h} . \tag{5.6}
\end{equation*}
$$

We observe that if $\mathfrak{h}$ depends on $r=\left|x^{\prime}\right|, x^{\prime} \in \mathbb{R}^{N-1}$, and $x_{N+1}$ only, so does $\psi$.
We will use the following weighted norms:

$$
\|\mathfrak{h}\|_{\mathcal{C}_{\beta}^{0, \mu}\left(\mathbb{R}^{N} \times \mathbb{R}\right)}=\sup _{x^{\prime} \in \mathbb{R}^{N}}\left(1+\varepsilon^{2}\left|x^{\prime}\right|^{2}\right)^{\beta}\|\mathfrak{h}\|_{\mathcal{C}^{0, \mu}\left(B\left(x^{\prime}, 1\right) \times \mathbb{R}\right)}, \quad \beta>0 .
$$

The weighted Hölder norms $\mathcal{C}_{\beta}^{2, \mu}\left(\mathbb{R}^{N} \times \mathbb{R}\right)$ are defined similarly. Note that the definition of the norm implies in particular that

$$
\|\mathfrak{h}\|_{\mathcal{C}_{\beta}^{0, \mu}\left(\mathbb{R}^{N} \times \mathbb{R}\right)}<\infty \Longrightarrow\|\mathfrak{h}\|_{\mathcal{C}^{0, \mu}\left(\mathbb{R}^{N} \times \mathbb{R}\right)}<\infty
$$

and thus, by a standard argument, we obtain the existence of a solution to (5.6), $\psi \in \mathcal{C}^{2, \mu}\left(\mathbb{R}^{N} \times \mathbb{R}\right)$. Now, to show that in fact

$$
\|\psi\|_{\mathcal{C}_{\beta}^{2, \mu}}\left(\mathbb{R}^{N} \times \mathbb{R}\right)=C\|\mathfrak{h}\|_{\mathcal{C}_{\beta}^{0, \mu}}^{\left(\mathbb{R}^{N} \times \mathbb{R}\right)},
$$

one can use a comparison argument based on the fact that the reciprocal of the weight function $\left(1+\varepsilon^{2}\left|x^{\prime}\right|^{2}\right)^{\beta}$ is a positive supersolution of (5.6). Details are left to the reader.

## 6 Infinite-Dimensional Reduction

### 6.1 Estimates for the Error of the Initial Approximation

Our first goal is to estimate the functions $\tilde{\mathfrak{g}}_{\varepsilon, j}$ defined in (4.14) and (4.16). Whenever convenient, we will indicate the fact that these functions depend on their functional arguments by writing $\tilde{\mathfrak{g}}_{\varepsilon, j}=\tilde{\mathfrak{g}}_{\varepsilon, j}\left(\phi_{\varepsilon, 1}, \phi_{\varepsilon, 2}, \psi_{\varepsilon}, \mathbf{h}_{\varepsilon}\right)$. In general, besides the assumptions on $\mathbf{h}_{\varepsilon}$, which we have made in (4.3)-(4.4), we will also assume that for some $\sigma \in(0, \sqrt{2})$ and $K>0$ we have, with $\beta_{\sigma}=1-\sigma / \sqrt{2}$,

$$
\begin{equation*}
\left\|\phi_{\varepsilon, j}\right\|_{\mathcal{C}_{\beta \sigma, \sigma}^{2, \mu}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)} \leq K \varepsilon^{2-\sigma \sqrt{2}} . \tag{6.1}
\end{equation*}
$$

About the function $\psi_{\varepsilon}$ we assume that, with some $\kappa>3$, we have

$$
\begin{equation*}
\|\psi\|_{\mathcal{C}_{\kappa}^{2}, \mu}^{\left(\mathbb{R}^{N} \times \mathbb{R}\right)}, ~ \leq K \varepsilon^{3} . \tag{6.2}
\end{equation*}
$$

Lemma 6.1. Under the preceding hypotheses there exists $a \sigma \in(0, \sqrt{2})$ such that the following estimate holds:

$$
\begin{align*}
& \left\|\tilde{\mathfrak{g}}_{\varepsilon, j}\right\|_{\mathcal{C}_{\beta \sigma, \sigma}^{0, \mu}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)} \leq  \tag{6.3}\\
& \quad C\left\{\varepsilon^{2-\sigma \sqrt{2}}+o(1) \sum_{j=1,2}\left\|\phi_{\varepsilon, j}\right\|_{\mathcal{C}_{\beta \sigma, \sigma}^{2, \mu}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)}+\left\|\psi_{\varepsilon}\right\|_{\mathcal{C}_{K}^{2, \mu}\left(\mathbb{R}^{N} \times \mathbb{R}\right)}\right\} .
\end{align*}
$$

The function $\mathfrak{g}_{\varepsilon, j}$ is a Lipschitz function of its arguments and we have

$$
\begin{align*}
& \left\|\tilde{\mathfrak{g}}_{\varepsilon, j}\left(\phi_{\varepsilon, 1}^{(1)}, \phi_{\varepsilon, 2}^{(1)}, \psi_{\varepsilon}^{(1)}, \mathbf{h}_{\varepsilon}^{(1)}\right)-\tilde{\mathfrak{g}}_{\varepsilon, j}\left(\phi_{\varepsilon, 1}^{(2)}, \phi_{\varepsilon, 2}^{(2)}, \psi_{\varepsilon}^{(2)}, \mathbf{h}_{\varepsilon}^{(2)}\right)\right\|_{\mathcal{C}_{\beta \sigma, \sigma}^{0, \mu}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)} \\
& \leq C\left\{\varepsilon^{2-\sigma \sqrt{2}}\left\|\mathbf{h}_{\varepsilon}^{(1)}-\mathbf{h}_{\varepsilon}^{(2)}\right\|_{\mathcal{C}_{\beta \sigma}^{2}, \mu}\left(\Gamma_{\varepsilon}\right)\right.  \tag{6.4}\\
& \\
& \left.\quad+\left\|(1) \sum_{j=1,2}\right\| \psi_{\varepsilon, j}^{(1)}-\psi_{\varepsilon, j}^{(2)}\left\|_{\mathcal{C}_{\beta \sigma, \sigma}^{2}, \mu\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)}\right\|_{\mathcal{C}_{\kappa}^{2}, \mu}\left(\mathbb{R}^{N} \times \mathbb{R}\right)\right\} .
\end{align*}
$$

Proof. The proof of this lemma follows by a somewhat tedious but rather straightforward calculation. Similar calculations can also be found in [7, 8]. We will outline the main part, which is the estimate of the term involving $S\left(w_{\varepsilon}\right)$. Note that $\tilde{\zeta}_{\varepsilon, j} S\left(w_{\varepsilon}\right)=\tilde{\zeta}_{\varepsilon, j} S\left(u_{\varepsilon}\right)$ (see (4.6)-(4.7)).

Let us denote $\tilde{u}_{\varepsilon}(y, z)=\left(X_{\varepsilon}^{*} u_{\varepsilon}\right)(y, z)$. We expand $\Delta$ near $\Gamma_{\varepsilon}$ in terms of the Fermi coordinates to get

$$
\begin{align*}
\left(X_{\varepsilon}^{*} S\left(u_{\varepsilon}\right)\right)= & \Delta_{\Gamma_{\varepsilon}} \tilde{u}_{\varepsilon}+\left[\partial_{z}^{2} \tilde{u}_{\varepsilon}+f\left(\tilde{u}_{\varepsilon}\right)\right] \\
& +\left[\varepsilon \partial_{z}\left(\pi_{\varepsilon, N+1}\right)-H_{\Gamma_{\varepsilon}} \partial_{z} \tilde{u}_{\varepsilon}-z\left|A_{\Gamma_{\varepsilon}}\right|^{2} \partial_{z} \tilde{u}_{\varepsilon}\right.  \tag{6.5}\\
& +\varepsilon \nabla_{\Gamma_{\varepsilon}}\left(\pi_{\varepsilon, N+1}\right) \cdot \nabla_{\Gamma_{\varepsilon}} \tilde{u}_{\varepsilon}+\mathbb{A}_{\varepsilon}\left[\tilde{u}_{\varepsilon}\right]+\mathbb{B}_{\varepsilon}\left[\tilde{u}_{\varepsilon}\right]-z^{2} R_{\Gamma_{\varepsilon}} \partial_{z} \widetilde{u}_{\varepsilon} .
\end{align*}
$$

Above, $\mathbb{A}_{\varepsilon}$ and $\mathbb{B}_{\varepsilon}$ are linear differential operators of second and first order, respectively, whose expressions in terms of local coordinates on $\Gamma_{\varepsilon}$ are given in Section 2.1. Most of the terms in (6.5) are estimated directly. The leading-order term is in fact given by

$$
\partial_{z}^{2} \tilde{u}_{\varepsilon}+f\left(\widetilde{u}_{\varepsilon}\right)=f\left(\widetilde{u}_{\varepsilon}\right)-f\left(H\left(z-f_{\varepsilon, 1}-h_{\varepsilon, 1}\right)\right)-f\left(-H\left(z-f_{\varepsilon, 2}-h_{\varepsilon, 2}\right)\right) .
$$

Using this, the definition of $\widetilde{u}_{\varepsilon}$, and (2.19), we can estimate, taking $\sigma \in(0, \sqrt{2})$,

$$
\begin{aligned}
\left|\partial_{z}^{2} \widetilde{u}_{\varepsilon}+f\left(\widetilde{u}_{\varepsilon}\right)\right| \leq & C\left\{H^{\prime}\left(z-f_{\varepsilon, 1}-h_{\varepsilon, 1}\right)\left[1+H\left(z-f_{\varepsilon, 2}-h_{\varepsilon, 2}\right)\right]\right. \\
& \left.+H^{\prime}\left(z-f_{\varepsilon, 2}-h_{\varepsilon, 1}\right)\left[1-H\left(z-f_{\varepsilon, 1}-h_{\varepsilon, 1}\right)\right]\right\} \\
\leq & C \max _{j}\left\{e^{-\sigma\left|z-f_{\varepsilon, j}\right|}\right\} \exp \left(-\frac{\sqrt{2}-\sigma}{\sqrt{2}} \log \frac{\omega_{\varepsilon}}{\varepsilon^{2}}\right) \\
\leq & C \varepsilon^{2-\sigma \sqrt{2}} \max _{j}\left\{e^{-\sigma\left|z-f_{\varepsilon, j}\right|}\right\} \omega_{\varepsilon}^{-\beta_{\sigma}} .
\end{aligned}
$$

Since we have $\varepsilon \partial_{z}\left(\pi_{\varepsilon, N+1}\right)-H_{\Gamma_{\varepsilon}}=0$, the remaining nonzero term in the first line in (6.5) is

$$
\begin{align*}
\Delta_{\Gamma_{\varepsilon}} \tilde{u}_{\varepsilon}-z\left|A_{\Gamma_{\varepsilon}}\right|^{2} \partial_{z} \tilde{u}_{\varepsilon}= & \sum_{j=1,2}(-1)^{j} H^{\prime}\left(z-f_{\varepsilon, j}-h_{\varepsilon, j}\right) \Delta_{\Gamma_{\varepsilon}}\left(f_{\varepsilon, j}+h_{\varepsilon, j}\right) \\
& +\left|A_{\Gamma_{\varepsilon}}\right|^{2} \sum_{j=1,2}\left(f_{\varepsilon, j}+h_{\varepsilon, j}\right) H^{\prime}\left(z-f_{\varepsilon, j}-h_{\varepsilon, j}\right) \\
& +\sum_{j=1,2} H^{\prime \prime}\left(z-f_{\varepsilon, j}-h_{\varepsilon, j}\right)\left|\nabla_{\Gamma_{\varepsilon}}\left(f_{\varepsilon, j}-h_{\varepsilon, j}\right)\right|^{2}  \tag{6.6}\\
& +\left|A_{\Gamma_{\varepsilon}}\right|^{2} \sum_{j=1,2}\left(z-f_{\varepsilon, j}-h_{\varepsilon, j}\right) H^{\prime}\left(z-f_{\varepsilon, j}-h_{\varepsilon, j}\right) .
\end{align*}
$$

We note that

$$
\begin{equation*}
\left|A_{\Gamma_{\varepsilon}}(r)\right|^{2}=\varepsilon^{2}\left|A_{\Gamma}(\varepsilon r)\right|^{2} \leq C \varepsilon^{2} \omega_{\varepsilon}^{-2}(r) . \tag{6.7}
\end{equation*}
$$

Each term in (6.6) is then estimated directly. The second line in (6.5) is easily seen to be smaller relative to the terms we have just considered. As for the terms involving functions $\phi_{\varepsilon, j}$, we observe that the largest among them is

$$
\left[L, \zeta_{\varepsilon, j}\right] \phi_{\varepsilon, j}=\Delta\left(\zeta_{\varepsilon, j} \phi_{\varepsilon, j}\right)-\zeta_{\varepsilon, j} \Delta \phi_{\varepsilon, j} .
$$

Using the fact that $\Delta \zeta_{\varepsilon, j}=o(1)$ and $\nabla \zeta_{\varepsilon, j}=o(1)$, which follows from the choice of the cutoff functions $\zeta_{\varepsilon, j}$, we can estimate this term by $o(1)\left\|\phi_{\varepsilon, j}\right\|_{\mathcal{C}_{\beta \sigma, \sigma}^{2, \mu}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)}$.

The rest of the proof is straightforward, and we leave the details to the reader.

Going back to the system (4.17) and taking into account the theory of the preceding section, we see that the functions $c_{\varepsilon, j}$ need to be determined from the formula

$$
\begin{equation*}
\mathrm{c}_{\varepsilon, j}=\frac{\int_{\mathbb{R}} \tilde{\mathfrak{a}}_{\varepsilon, j}\left(y, \mathrm{t}_{j}\right) H^{\prime}\left(\mathrm{t}_{j}\right) d \mathrm{t}_{j}}{\int_{\mathbb{R}}\left(H^{\prime}\left(\mathrm{t}_{j}\right)\right)^{2} \zeta_{\varepsilon, j}\left(y, \mathrm{t}_{j}\right) d \mathrm{t}_{j}} . \tag{6.8}
\end{equation*}
$$

Using Lemma 6.1 we see that statements analogous to (6.3) and (6.4) hold when we replace $\tilde{\mathfrak{g}}_{\varepsilon, j}$ by $\tilde{\mathfrak{g}}_{\varepsilon, j}+\mathrm{c}_{\varepsilon, j} H^{\prime}\left(\mathrm{t}_{j}\right) \zeta_{\varepsilon, j}$.

Next we will consider the right-hand side of equation (4.15). We have the following:

Lemma 6.2. Under the same hypotheses as in Lemma 6.1 and assuming that the constant $M>0$ in (4.5) and (4.8) is large enough, there exist $\kappa>3$ and $\gamma>1$ such that we have

$$
\begin{gather*}
\left\|\mathfrak{h}_{\varepsilon}\right\|_{\mathcal{C}_{\mathcal{K}}^{0}, \mu}\left(\mathbb{R}^{N} \times \mathbb{R}\right)  \tag{6.9}\\
\leq C
\end{gather*} \varepsilon^{3}+\varepsilon^{\gamma} \sum_{j=1,2}\left\|\phi_{\varepsilon, j}\right\|_{\mathcal{C}_{\beta \sigma, \sigma}^{2, \mu}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)} .
$$

Considering $\mathfrak{h}_{\varepsilon}$ as a function of ( $\phi_{\varepsilon, 1}, \phi_{\varepsilon, 2}, \psi_{\varepsilon}, \mathbf{h}_{\varepsilon}$ ),

$$
\begin{align*}
& \left\|\mathfrak{h}_{\varepsilon}\left(\phi_{\varepsilon, 1}^{(1)}, \phi_{\varepsilon, 2}^{(1)}, \psi_{\varepsilon}^{(1)}, \mathbf{h}_{\varepsilon}^{(1)}\right)-\mathfrak{h}_{\varepsilon}\left(\phi_{\varepsilon, 1}^{(2)}, \phi_{\varepsilon, 2}^{(2)}, \psi_{\varepsilon}^{(2)}, \mathbf{h}_{\varepsilon}^{(2)}\right)\right\|_{\mathcal{C}_{\beta \sigma+\sigma \delta}^{0, \mu}\left(\mathbb{R}^{N} \times \mathbb{R}\right)} \\
& \leq C  \tag{6.10}\\
& \leq \\
& \quad+\varepsilon^{3}\left\|\mathbf{h}_{\varepsilon}^{(1)}-\mathbf{h}_{\varepsilon}^{(2)}\right\|_{\mathcal{C}_{\beta \sigma}^{2, \mu}\left(\Gamma_{\varepsilon}\right)}+\varepsilon^{\gamma} \sum_{j=1,2}\left\|\phi_{\varepsilon, j}^{(1)}-\phi_{\varepsilon, j}^{(2)}\right\|_{\mathcal{B}_{\beta \sigma, \sigma}^{2, \mu}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)} \\
& \left.\quad \psi_{\varepsilon}^{(1)}-\psi_{\varepsilon}^{(2)} \|_{\mathcal{C}_{\kappa}^{2, \mu}\left(\mathbb{R}^{N} \times \mathbb{R}\right)}\right\} .
\end{align*}
$$

A proof of this estimate is omitted, since similar results are proven in [7, 8] and no essentially new elements are needed to carry out the argument in the present case. We only point out that the support of the function $\mathfrak{h}_{\varepsilon}$ is in the set where $\left|z-f_{\varepsilon, j}\right|>M \log \frac{\omega_{\varepsilon}}{\varepsilon^{2}}$, from which it follows that all exponentially decaying terms are very small, like $\mathcal{O}\left(\varepsilon^{3}\right)$ at least.

### 6.2 Projected Nonlinear Problem

Our objective in this section is to solve (4.13)-(4.15). Given the linear theory available and the results of the preceding section, we will achieve this by a simple fixed point argument.

Let functions $\widetilde{\phi}_{\varepsilon, j}, j=1,2$, and $\widetilde{\psi}_{\varepsilon}$ satisfying assumptions (6.1)-6.2) be fixed. We will also choose $\mathbf{h}_{\varepsilon}$ satisfying (4.4). We first use the linear theory of Section 5 to solve the following system:

$$
\begin{align*}
& \left(\Delta_{\Gamma_{\varepsilon}}+\partial_{\mathrm{t}_{j}}^{2}+f^{\prime}\left(H\left(\mathrm{t}_{j}\right)\right)\right) \phi_{\varepsilon, j}  \tag{6.11}\\
& \quad=\tilde{\mathfrak{g}}_{\varepsilon, j}\left(y, \mathrm{t}_{j} ; \widetilde{\phi}_{\varepsilon, 1}, \widetilde{\phi}_{\varepsilon, 2}, \widetilde{\psi}_{\varepsilon}, \mathbf{h}_{\varepsilon}\right)+\mathrm{c}_{\varepsilon, j} H^{\prime}\left(\mathrm{t}_{j}\right), \quad j=1,2, \\
& \quad \int_{\mathbb{R}} \phi_{\varepsilon, j}\left(y, \mathrm{t}_{j}\right) H^{\prime}\left(\mathrm{t}_{j}\right) d \mathrm{t}_{j}=0, \quad j=1,2,  \tag{6.12}\\
& \quad\left(\Delta+\varepsilon \partial_{x_{N+1}}-2\right) \psi_{\varepsilon}=\mathfrak{h}_{\varepsilon}\left(x ; \tilde{\phi}_{\varepsilon, 1}, \widetilde{\phi}_{\varepsilon, 2}, \widetilde{\psi}_{\varepsilon}, \mathbf{h}_{\varepsilon}\right) . \tag{6.13}
\end{align*}
$$

This is equivalent to (4.13)-4.15) when $\tilde{\phi}_{\varepsilon, j}=\phi_{\varepsilon, j}$ and $\tilde{\psi}_{\varepsilon}=\psi_{\varepsilon}$. In fact, using Lemma 6.1 and Lemma 6.2, we obtain existence of such a fixed point satisfying (6.1)-(6.2) by the Banach fixed point theorem. To do this we first solve (6.13) for $\psi_{\varepsilon}$ as a function of ( $\widetilde{\phi}_{\varepsilon, 1}, \widetilde{\phi}_{\varepsilon, 2}, \mathbf{h}_{\varepsilon}$ ). Existence of $\psi_{\varepsilon}$ follows by a fixed point argument using Lemma 6.2 and the results in Section 5.3 . We have in fact

$$
\left\|\psi_{\varepsilon}\right\|_{\mathcal{C}_{K}^{2, \mu}\left(\mathbb{R}^{N} \times \mathbb{R}\right)} \leq C\left\{\varepsilon^{3}+\varepsilon^{\gamma} \sum_{j=1,2}\left\|\phi_{\varepsilon, j}\right\|_{\mathcal{C}_{\beta \sigma, \sigma}^{2, \mu}\left(\Gamma_{\varepsilon} \times \mathbb{R}\right)}\right\},
$$

with a similar estimate showing the Lipschitz character of $\psi_{\varepsilon}$. Given this we solve (6.11) using again the Banach fixed point theorem. Let us summarize this:

Lemma 6.3. Under the above hypotheses there exists a unique solution ( $\phi_{\varepsilon, 1}, \phi_{\varepsilon, 2}$, $\psi_{\varepsilon}$ ) of (6.11) and (6.13) satisfying (6.1) and (6.2).

### 6.3 Solution of the Reduced Problem

At this point we are left with the task of adjusting $\mathbf{h}_{\varepsilon}$ in such a way that $\mathrm{c}_{\varepsilon, j} \equiv 0$. For this let us observe that the map

$$
\left(\widetilde{\phi}_{\varepsilon, 1}, \widetilde{\phi}_{\varepsilon, 2}, \widetilde{\psi}_{\varepsilon} ; \mathbf{h}_{\varepsilon}\right) \longmapsto\left(\phi_{\varepsilon, 1}, \phi_{\varepsilon, 2}, \psi_{\varepsilon}\right)
$$

is a uniform contraction (with a small Lipschitz constant) with respect to $\mathbf{h}_{\varepsilon}$. It follows that $\left(\phi_{\varepsilon, 1}, \phi_{\varepsilon, 2}, \psi_{\varepsilon}\right)$ are Lipschitz functions of $\mathbf{h}_{\varepsilon}$ with small Lipschitz constants. This last fact can be easily seen from Lemma 6.1 and Lemma 6.2. Another important fact is that since we have assumed initially that $f_{\varepsilon, j}$ and $h_{\varepsilon, j}$ are functions of $r$, where $r=\left|x^{\prime}\right|,\left(x^{\prime}, x_{N+1}\right) \in \mathbb{R}^{N+1}$, we have $\left(\phi_{\varepsilon, 1}, \phi_{\varepsilon, 2}, \psi_{\varepsilon}\right)$ are functions of $(r, z)$ only, at least near $\Gamma_{\varepsilon}$, i.e., where the Fermi coordinates are defined. In fact, instead of working in an abstract setting, which does not refer to the rotational symmetry of $\Gamma_{\varepsilon}$, we could have reduced the whole problem to the one in the half-plane $\mathbb{R}_{+}^{2}=\left(r, x_{N+1}\right)$, and think of $\Gamma_{\varepsilon}$ as a curve, with $(r, z)$ as its Fermi coordinates. Then the end result, from the point of view of existence of ( $\phi_{\varepsilon, 1}, \phi_{\varepsilon, 2}, \psi_{\varepsilon}$ ), would of course be the same. Summarizing, all functions involved depend on $x=\left(x^{\prime}, x_{N+1}\right)$, through $r(x)=\left|x^{\prime}\right|$ and $x_{N+1}$, and when expressed in Fermi coordinates $(y, z)$, they depend on $r(y)=\left|y^{\prime}\right|$ and $z$ only.

Now we will find the exact conditions for $\mathbf{h}_{\varepsilon}$ that guarantee that $\mathrm{c}_{\varepsilon, j} \equiv 0$. We will show that they result in a nonhomogeneous and nonlocal Jacobi-Toda system, quite similar to the one already studied in Section 3 . From the theory developed in this section, the existence of $\mathbf{h}_{\varepsilon}$ will follow immediately, thus completing the proof of Theorem 1.1. Our first task is then to justify rigorously the formal calculations in Section 2.2. In fact, with the notation as in the previous sections, we need to adjust $\mathbf{h}_{\varepsilon}$ so that

$$
\int_{\mathbb{R}} \tilde{\mathfrak{g}}_{\varepsilon, j}\left(r, \mathrm{t}_{j}\right) H^{\prime}\left(\mathrm{t}_{j}\right) d \mathrm{t}_{j}=0, \quad j=1,2 .
$$

Let us recall that $\tilde{\mathfrak{g}}_{\varepsilon, j}$ depends on ( $\phi_{\varepsilon, 1}, \phi_{\varepsilon, 2}, \psi_{\varepsilon}, \mathbf{h}_{\varepsilon}$ ), that ( $\phi_{\varepsilon, 1}, \phi_{\varepsilon, 2}, \psi_{\varepsilon}$ ) depend nonlocally on $\mathbf{h}_{\varepsilon}$, and that this dependence involves second derivatives of $\mathbf{h}_{\varepsilon}$. Thus its projection onto $H^{\prime}\left(\mathrm{t}_{j}\right)$ will be a nonlocal, second-order ODE in terms of the radial variable $r$.

Let us write

$$
\tilde{\mathfrak{g}}_{\varepsilon, j}=\widetilde{\zeta}_{\varepsilon, j} S\left(w_{\varepsilon}\right)+\widehat{\mathfrak{g}}_{\varepsilon, j}, \quad \hat{\mathfrak{g}}_{\varepsilon, j}=\hat{\mathfrak{g}}_{\varepsilon, j}\left(\phi_{\varepsilon, 1}, \phi_{\varepsilon, 2}, \psi_{\varepsilon}, \mathbf{h}_{\varepsilon}\right) .
$$

Examining the expression for $S\left(u_{\varepsilon}\right)$ in (6.5), we see that as a function of $\left(r, \mathrm{t}_{j}\right)$ it has general form (say, where $\left.\tilde{\zeta}_{\varepsilon, j} \equiv 1\right) S\left(u_{\varepsilon}\right)\left(r, \mathrm{t}_{j}\right)=S\left(u_{\varepsilon}\right)\left(r, \mathrm{t}_{j}-h_{\varepsilon, j}\right)$. It is therefore more convenient to integrate $\tilde{\mathfrak{g}}_{\varepsilon, j}$ against $H^{\prime}\left(\mathrm{t}_{j}-h_{\varepsilon, j}\right)$ rather than
$H^{\prime}\left(\mathrm{t}_{j}\right)$. It is easily seen that $\mathrm{c}_{j, \varepsilon}=0$ when

$$
\begin{align*}
& \int_{\mathbb{R}} \widetilde{\mathfrak{g}}_{\varepsilon, j}\left(r(y), \mathrm{t}_{j}\right) H^{\prime}\left(\mathrm{t}_{j}-h_{\varepsilon, j}\right) d \mathrm{t}_{j} \\
& \quad=\int_{\mathbb{R}} \tilde{\zeta}_{\varepsilon, j} S\left(w_{\varepsilon}\right)\left(r(y), \mathrm{t}_{j}\right) H^{\prime}\left(\mathrm{t}_{j}-h_{\varepsilon, j}\right) \mathrm{t}_{j}  \tag{6.14}\\
& \quad \quad+\int_{\mathbb{R}} \hat{\mathfrak{g}}_{\varepsilon, j} H^{\prime}\left(\mathrm{t}_{j}-h_{\varepsilon, j}\right) d \mathrm{t}_{j} \\
& \quad=\Pi_{\varepsilon, j}+\hat{\Pi}_{\varepsilon, j}=0 .
\end{align*}
$$

As we have argued in Section (2.2) the main term in the above integral (remembering that by definition $w_{\varepsilon}=u_{\varepsilon}$ in supp $\zeta_{\varepsilon, j}$ ) comes from

$$
\Pi_{\varepsilon, j}=\int_{\mathbb{R}} \zeta_{\varepsilon, j} S\left(u_{\varepsilon}\right)\left(r(y), \mathrm{t}_{j}\right) H^{\prime}\left(\mathrm{t}_{j}-h_{\varepsilon, j}\right) \mathrm{t}_{j}
$$

while the remaining part of the projection, denoted by $\hat{\Pi}_{\varepsilon, j}$, is a lower-order term.
Repeating calculations in Section 2.2 and taking into account formula (6.5), one can derive the following expression:

$$
\begin{equation*}
\Pi_{\varepsilon, j}=\alpha_{0} J_{\Gamma_{\varepsilon}}\left(f_{\varepsilon, j}+h_{\varepsilon, j}\right)+\mathcal{T}_{j}\left(\mathbf{f}_{\varepsilon}+\mathbf{h}_{\varepsilon}\right)+q_{\varepsilon, j}\left(\mathbf{f}_{\varepsilon}+\mathbf{h}_{\varepsilon}\right), \tag{6.15}
\end{equation*}
$$

where, for a vector function $\mathbf{v}=\left(v_{1}, v_{2}\right)$, on $\Gamma_{\varepsilon}$ we have denoted

$$
\begin{align*}
J_{\Gamma_{\varepsilon}}\left(v_{j}\right) & =\Delta_{\Gamma_{\varepsilon}} v_{j}+\left|A_{\Gamma_{\varepsilon}}\right|^{2} v_{j}+\varepsilon \nabla_{\Gamma_{\varepsilon}}\left(\pi_{\varepsilon, N+1}\right) \cdot \nabla_{\Gamma_{\varepsilon}} v_{j} \\
\mathcal{T}_{j}(\mathbf{v}) & =-e^{\sqrt{2}\left(v_{j-1}-v_{j}\right)}+e^{\sqrt{2}\left(v_{j}-v_{j+1}\right)} . \tag{6.16}
\end{align*}
$$

We observe that the main order term in $q_{\varepsilon, j}$ (see (6.5)) comes from

$$
z^{2} \tilde{\zeta}_{\varepsilon, j} R_{\Gamma_{\varepsilon}} \partial_{z} \tilde{u}_{\varepsilon} \approx\left(\mathrm{t}_{j}-f_{\varepsilon, j}\right)^{2} \sum_{\ell=1}^{N} \mathbb{k}_{\Gamma_{\varepsilon}, \ell}^{3} H^{\prime}\left(\mathrm{t}_{j}-h_{\varepsilon, j}\right),
$$

where $\mathbb{k}_{\Gamma_{\varepsilon}, \ell}$ are the principal curvatures of $\Gamma_{\varepsilon}$. Direct calculations show that

$$
\left|\mathbb{k}_{\Gamma_{\varepsilon}, \ell}^{3}\right| \approx \varepsilon^{3} \omega_{\varepsilon}^{-3 / 2} .
$$

Taking into account the assumptions we have made at the beginning on $\mathbf{f}_{\varepsilon}$, and $\mathbf{h}_{\varepsilon}$ in (4.2)-(4.3), we see that there exist $\beta>0$ and $\rho>0$ such that

$$
\left\|q_{\varepsilon, j}\right\|_{\mathcal{C}_{1+\beta}^{0, \mu}\left(\Gamma_{\varepsilon}\right)} \leq C \varepsilon^{2+\rho}
$$

Identifying functions on $\Gamma_{\varepsilon}$ and $\Gamma$ by $v_{\varepsilon}(r)=v(\varepsilon r)$, so that $q_{\varepsilon, j}(r)=q_{j}(\varepsilon r)$, we get from the above

$$
\left\|q_{j}\right\|_{\mathcal{C}_{1+\beta}^{0, \mu}(\Gamma)} \leq C \varepsilon^{2+\rho-\mu}
$$

Function $q_{j}$ now depends on the functions $\mathbf{f}$ and $\mathbf{h}$ defined on $\Gamma$. Similar statements hold for the remaining term in 6.14; namely, we have

$$
\left\|\hat{\Pi}_{\varepsilon, j}\right\|_{C_{1+\beta}^{0, \mu}\left(\Gamma_{\varepsilon}\right)} \leq C \varepsilon^{2+\rho}
$$

and, scaling back to $\Gamma$, we can write

$$
\left\|\widehat{\Pi}_{j}\right\|_{\mathcal{C}_{1+\beta}^{0, \mu}(\Gamma)} \leq C \varepsilon^{2+\rho-\mu} .
$$

We let $\mu>0$ be a small number and set $\kappa=\rho-\mu>0$, also choosing it in such a way that $\tau<\kappa$ (see (4.3). Denoting by $J_{\Gamma}$ the scaled operator in (6.16) and setting $\widehat{q}_{j}=q_{j}+\widehat{\Pi}_{j}$, we then get

$$
\begin{equation*}
\alpha_{0} \varepsilon^{2} J_{\Gamma}\left(f_{j}+h_{j}\right)+\mathcal{T}_{j}(\mathbf{f}+\mathbf{h})=\widehat{q}_{j} \tag{6.17}
\end{equation*}
$$

This is a Jacobi-Toda system, which can be solved using the theory we developed in the proof of Proposition 3.2 and in particular the result of Lemma 3.6 In fact, $\widehat{q}_{j}$ is a Lipschitz function of $\mathbf{h}$ since it follows from the Lipschitz character of $S\left(w_{\varepsilon}\right)$, $\phi_{\varepsilon, j}$, and $\psi_{\varepsilon}$ as functions of $\boldsymbol{h}$ that

$$
\left\|\widehat{q}_{j}\left(\mathbf{h}^{(1)}\right)-\widehat{q}_{j}\left(\mathbf{h}^{(2)}\right)\right\|_{\mathcal{C}_{1+\beta}^{0, \mu}(\Gamma)} \leq C \varepsilon^{2+\kappa}\left\|\mathbf{h}^{(1)}-\mathbf{h}^{(2)}\right\|_{\mathcal{C}_{\beta}^{2, \mu}(\Gamma)} .
$$

Defining

$$
\mathcal{T}_{j}(\mathbf{f}+\mathbf{h})-\mathcal{T}_{j}(\mathbf{f})-\mathcal{T}_{j}^{\prime}(\mathbf{f}) \mathbf{h}=\mathcal{N}_{j}(\mathbf{h})
$$

we also have

$$
\left\|\mathcal{N}_{j}(\mathbf{h})\right\|_{\mathcal{C}_{1+\beta}^{0, \mu}(\Gamma)} \leq C \varepsilon^{2+\tau}\|\mathbf{h}\|_{\mathcal{C}_{\beta}^{2, \mu}(\Gamma)}
$$

Similarly, $\mathcal{N}_{j}(\mathbf{h})$ is a Lipschitz function of $\mathbf{h}$. Since we have chosen $\mathbf{f}$ to be a solution of the homogeneous version of 6.17) we are left with

$$
\begin{equation*}
\alpha_{0} \varepsilon^{2} J_{\Gamma}\left(h_{j}\right)+\mathcal{T}_{j}^{\prime}(\mathbf{f})(\mathbf{h})=\tilde{q}_{j}, \quad \tilde{q}_{j}=\widehat{q}_{j}-\mathcal{N}_{j} \tag{6.18}
\end{equation*}
$$

The left-hand side of this equation is the linearized Jacobi-Toda system, and now Lemma 3.6 can be employed directly to solve (6.17) using the Banach fixed point theorem. As similar arguments can be found for instance in [7] and [8], we omit the details here. With this last step we complete our proof.

## 7 An Example of a Traveling Wave with a Nonconvex Front

In this section we will prove Theorem 1.2. We will begin with some preliminary facts about the asymptotic form of the nonconvex traveling front.

### 7.1 Traveling, Catenoidlike Surface

We will summarize here an existence result proven in [5].
Proposition 7.1. For each $R>0$ there exist rotationally symmetric, graphical solutions to the mean curvature flow, given by $F_{R}^{ \pm}: \mathbb{R}^{N} \backslash B_{R}(0) \times \mathbb{R} \rightarrow \mathbb{R}$, and translating with speed $c=1$, where

$$
F_{R}^{ \pm}(r, t)=t+W_{R}^{ \pm}(r)
$$

The functions $W_{R}^{ \pm}$satisfy

$$
\begin{equation*}
W_{R}^{ \pm}(r)=\frac{r^{2}}{2(N-1)}-\log r+C^{ \pm}+\mathcal{O}\left(r^{-1}\right), \quad r \rightarrow \infty \tag{7.1}
\end{equation*}
$$

Moreover, the union of these graphs forms a complete nonconvex translating solution to the mean curvature flow.

In what follows by $\Sigma$, we denote the surface obtained by taking the union of the graphs of $W_{R}^{ \pm}$, and by $\Sigma_{\varepsilon}$ we denote its scaled version. The individual graphs of each function $W_{\mathbb{R}}^{ \pm}$will be referred to as the ends of $\Sigma$ and will be denoted by $\Sigma^{ \pm}$, respectively, with a similar notation for the scaled versions. We assume that the constants $C^{ \pm}$appearing in $(7.1)$ are such that $C^{-}<C^{+}$, and we will call $\Sigma^{-}$ $\left(\Sigma^{+}\right)$the lower (the upper) end of $\Sigma$. Also, in order to not to complicate notation, we will not indicate explicitly the dependence of the surface $\Sigma$ on $R$. Nevertheless, the reader should keep in mind that our results are valid for the whole family of traveling catenoids parametrized by $R$.

The surface $\Sigma$ is an embedded, rotationally symmetric, and genus 0 surface in $\mathbb{R}^{N}$, and in some sense it is a counterpart of the usual catenoid, now in the context of the eternal solutions of the mean curvature flow. Another important, obvious property is its nonconvexity.

Comparing the asymptotic formula (7.1) with the asymptotic formula for $F$, we notice that as $r \rightarrow \infty$ the ends of $\Sigma$ remain at a constant distance from $\Gamma$. Indeed, we have

$$
\begin{equation*}
\left|F(r)-1-W_{R}^{ \pm}(r)+C^{ \pm}\right|=\mathcal{O}\left(r^{-1}\right), \quad r \rightarrow \infty \tag{7.2}
\end{equation*}
$$

This is important in the calculation of various geometric characteristics of $\Sigma$. In fact, formula 7.2 says that the mean curvature $H_{\Sigma}$, the second fundamental form $A_{\Sigma}, \nabla_{\Sigma}$, and $\Delta_{\Sigma}$ are, for $r$ sufficiently large, very close to their counterparts on $\Gamma$. Thus in what follows, we may omit many of the explicit calculations and appeal to the calculation we have already done for $\Gamma$.

### 7.2 An Improvement of the Initial Profile

The fact that the ends of $\Sigma$ are asymptotically parallel means that if we want to use its scaled version $\Sigma_{\varepsilon}$ as a model for a traveling wave with the speed $c=\varepsilon$, we must perturb the ends of the surface. To see this, let us denote the signed distance to $\Sigma_{\varepsilon}$ by $z=z(x)$ for $x \in \mathbb{R}^{N+1}$ close to $\Sigma_{\varepsilon}$. Then it is natural to take $u_{\varepsilon}=H(z)$ as the first approximation to the solution. A short calculation will convince the reader that, since the ends of $\Sigma_{\varepsilon}$ are parallel, the error $S\left(u_{\varepsilon}\right)$ of this approximation contains a term of order $\mathcal{O}\left(e^{-1 / \varepsilon}\right)$. This means that $S\left(u_{\varepsilon}\right)$ is globally a very small function of $\varepsilon$ but it is not a decaying function of $r=\left|x^{\prime}\right|$ along $\Sigma_{\varepsilon}$.

To remedy this situation we will consider an improvement of the initial profile $\Sigma_{\varepsilon}$. In general we want a new surface $\widetilde{\Sigma}_{\varepsilon}$ to be a normal graph over $\Sigma_{\varepsilon}$, to be
identical with $\Sigma_{\varepsilon}$ on a compact set, and to have ends that are diverging from one another as $r \rightarrow \infty$.

To give a formal definition, we need to introduce some notation. We let $\chi$ be a smooth cutoff function such that $\chi(t)=0, t \leq 1$, and $\chi(t)=1, t \geq 2$. By $r_{\varepsilon}$ we denote a number to be determined later on and about which we assume initially that, with some $c<C$,

$$
\begin{equation*}
r_{\varepsilon} \gg e^{\frac{c}{\varepsilon}} \quad \text { and } \quad r_{\varepsilon} \ll e^{\frac{C}{\varepsilon}} \tag{7.3}
\end{equation*}
$$

Next, we will fix an orientation on $\Sigma$ in such a way that a unit normal $n$ is interior to this component of $\mathbb{R}^{N+1} \backslash \Sigma$, which contains the origin. By $n_{\varepsilon}(y)=n(\varepsilon y)$ we denote the corresponding normal on $\Sigma_{\varepsilon}$, and by $n^{ \pm}$and $n_{\varepsilon}^{ \pm}$we denote the restrictions of $n$ and $n_{\varepsilon}$ to the ends of $\Sigma$. Finally, by $\Theta \in S^{N-1}$ we denote points on $S^{N-1}$.

The new surface $\widetilde{\Sigma}_{\varepsilon}$ will be a union of its lower and upper ends $\widetilde{\Sigma}_{\varepsilon}^{ \pm}$given by

$$
\begin{equation*}
\widetilde{\Sigma}_{\varepsilon}^{ \pm}=\left\{\left.\left(r \Theta, \frac{1}{\varepsilon} W_{R}^{ \pm}(\varepsilon r)\right)+\chi\left(\frac{r}{r_{\varepsilon}}\right) n^{ \pm}(\varepsilon r, \Theta) f^{ \pm}(\varepsilon r) \right\rvert\, r \geq R, \Theta \in S^{N-1}\right\}, \tag{7.4}
\end{equation*}
$$

where the radial functions $f^{ \pm}: \Sigma \rightarrow \mathbb{R}$ are still to be determined.

## Construction of $f^{ \pm}$

Choosing the functions $f^{ \pm}$is a subtle point of our problem. To give some motivation, let us recall how in the preceding considerations we have determined the functions $f_{1}, f_{2}: \Gamma \rightarrow \mathbb{R}$, which model the traveling fronts near $\Gamma_{\varepsilon}$. Restricting our attention to $r \gg 1$, we observe that, to main order we needed to solve an algebraic equation

$$
\begin{equation*}
\left|A_{\Gamma}\right|^{2} u=\frac{e^{-u}}{\delta^{2}}, \quad \delta=\frac{\varepsilon \sqrt{\alpha_{0}}}{2^{3 / 4}} \tag{7.5}
\end{equation*}
$$

and then we obtained, to main order,

$$
f_{1} \approx-\frac{1}{2 \sqrt{2}} u, \quad f_{2} \approx \frac{1}{2 \sqrt{2}} u
$$

Equation (7.5) describes a balance between the interactions of the ends due to the exponential decay of the heteroclinic to the stable phases $\pm 1$ and the geometry of the moving front $\Gamma$. Now we need to discover the analogue of (7.5) with $\Gamma$ replaced by $\Sigma$. The natural guess would be to take $\left|A_{\Sigma}\right|^{2}$ on the right-hand side and leave the exponential function on the left. However, the story is not so simple because, altering the ends of $\Sigma$ by adding normal perturbations as described above, we have changed the character of the surface - the new surface is not a translating solution of the mean curvature flow anymore. To take this into account, we solve (instead of (7.5) the following problem:

$$
\begin{equation*}
\frac{F_{r r}}{1+\left|F_{r}\right|^{2}} u=\frac{e^{-u}}{\widetilde{\delta}^{2}}, \quad \tilde{\delta}=\tilde{\alpha} \varepsilon, \tag{7.6}
\end{equation*}
$$

where $\tilde{\alpha}>0$ is a constant to be specified later. In what follows, we will show that we can choose $\tilde{\alpha}$ in such a way that defining

$$
\begin{equation*}
f^{ \pm}=\frac{1}{2 \sqrt{2}} u \tag{7.7}
\end{equation*}
$$

and using the modified surface as a model for the traveling wave, we can achieve the following:
(1) If the approximate solution is defined by $u_{\varepsilon}=H(z)$, where $z$ is the signed distance from $\widetilde{\Sigma}_{\varepsilon}$, then, at least near this surface, the error of the approximation $S\left(u_{\varepsilon}\right)$ is a small function of $\varepsilon$, and it also decays as $r \rightarrow \infty$ at an algebraic rate in $r$.
(2) The projection of the error onto $H^{\prime}(z)$, namely $\int_{\mathbb{R}} S\left(u_{\varepsilon}\right) H^{\prime}(z)$, is a function that behaves like

$$
\frac{\varepsilon^{2+\kappa}}{\left(1+\varepsilon^{2} r^{2}\right)^{1+\beta}} \quad \text { as } r \rightarrow \infty
$$

These two claims, which we will make more precise later, are sufficient to implement a Lyapunov-Schmidt construction quite similar to the one presented in the previous sections and, as a result, prove Theorem 1.2.

Let us go back to equation (7.6). Based on the known asymptotic behavior of the function $F(r)$ and its derivatives, one can prove the following:
Lemma 7.2. Let $u=u(r)$ be the solution of (7.6), and let $f^{ \pm}=f^{ \pm}(r)$ be the functions defined in (7.7). There exist $r_{0}>R$ and $C>0$ such that for all $r>r_{0}$ it holds that

$$
\begin{equation*}
\left(f^{+}(r)+f^{-}(r)\right) \geq \frac{2}{\sqrt{2}} \log \left(\frac{1+r^{2}}{\varepsilon^{2}}\right)-C \log \log \left(\frac{1+r^{2}}{\varepsilon^{2}}\right) \tag{7.8}
\end{equation*}
$$

From now on $\widetilde{\Sigma}_{\varepsilon}$ will be the surface we defined in (7.4) with $f^{ \pm}$as in Lemma 7.2. By $\tilde{n}_{\varepsilon}$ we will denote its unit normal, and by $\tilde{n}$ the unit normal of its scaled version $\widetilde{\Sigma}$. These vectors are chosen in such a way that $\widetilde{n}_{\varepsilon}$ is interior with respect to the connected component of $\mathbb{R}^{N+1} \backslash \widetilde{\Sigma}_{\varepsilon}$, which contains the origin.

### 7.3 Construction of the Initial Approximation

We will consider the Fermi coordinates associated with the surface $\widetilde{\Sigma}_{\varepsilon}$ :

$$
x \longmapsto(y, z), \quad y \in \tilde{\Sigma}_{\varepsilon}, \quad z=\operatorname{dist}\left(x, \tilde{\Sigma}_{\varepsilon}\right)
$$

in a neighborhood of $\mathcal{U}_{\varepsilon}$ of this surface. We let $\mathcal{U}_{\varepsilon}$ be such that this map is a diffeomorphism; namely we define

$$
\mathcal{U}_{\varepsilon}:=\left\{\begin{array}{c|c}
x \in \mathbb{R}^{N+1} & |z| \leq \frac{C(\Sigma)}{\varepsilon}\left[1-\chi\left(\frac{r}{r_{\varepsilon}}\right)\right]+\frac{1}{2} \chi\left(\frac{r}{r_{\varepsilon}}\right)\left(f^{+}(\varepsilon r)+f^{-}(\varepsilon r)\right),  \tag{7.9}\\
x=y+z \widetilde{n}_{\varepsilon}(y), r=r(y)
\end{array}\right\} .
$$

The constant $C(\Sigma)>0$ depends on $\Sigma$ only. As before, for $u: \mathcal{U}_{\varepsilon} \rightarrow \mathbb{R}^{k}$, by ( $\left.X_{\varepsilon}^{*} u\right)(y, z)$ we denote the pullback of $u$ by this diffeomorphism. At this point we
will choose conveniently the value of $r_{\varepsilon}$ by letting it be a solution of the following equation:

$$
\begin{equation*}
\frac{C(\Sigma)}{\varepsilon}=\frac{1}{2}\left(f^{+}\left(2 \varepsilon r_{\varepsilon}\right)+f^{-}\left(2 \varepsilon r_{\varepsilon}\right)\right) \Longrightarrow r_{\varepsilon} \sim e^{\frac{\widetilde{\tau}}{\varepsilon}} . \tag{7.10}
\end{equation*}
$$

As a next step we define a smooth cutoff function $\rho_{\varepsilon}$ that is supported in $\mathcal{U}_{\varepsilon}$ and such that

$$
\left(X_{\varepsilon}^{*} \rho_{\varepsilon}\right)(y, z)=1, \quad \operatorname{dist}\left(x, \partial \mathcal{U}_{\varepsilon}\right) \leq 1, \quad x=y+z \widetilde{n}_{\varepsilon}(y) .
$$

To be more precise, we take for instance a smooth cutoff function $\rho(t)$ such that $\rho(t)=1, t \leq-1$, and $\rho(t)=0, t \geq 0$, and set

$$
\begin{equation*}
\left(X_{\varepsilon}^{*} \rho_{\varepsilon}\right)(y, z)=\rho\left(|z|-\frac{C(\Sigma)}{\varepsilon}\left[1-\chi\left(\frac{r}{r_{\varepsilon}}\right)\right]-\frac{1}{2} \chi\left(\frac{r}{r_{\varepsilon}}\right)\left|f^{+}(\varepsilon r)+f^{-}(\varepsilon r)\right|\right) . \tag{7.11}
\end{equation*}
$$

In order to use a Lyapunov-Schmidt reduction procedure, we have to allow possible further perturbations of the surface $\widetilde{\Sigma}_{\varepsilon}$. They will be given as normal graphs over $\widetilde{\Sigma}_{\varepsilon}$ of $\mathcal{C}_{\beta}^{2, \mu}\left(\widetilde{\Sigma}_{\varepsilon}\right)$ functions. More precisely, we start with radial functions $h: \widetilde{\Sigma} \rightarrow \mathbb{R}$ such that

$$
\begin{equation*}
\|h\|_{\mathcal{C}_{\beta}^{2, \mu}(\tilde{\Sigma})} \leq \varepsilon^{\tau}, \quad \text { some } \tau>0, \beta>0 \tag{7.12}
\end{equation*}
$$

We will also make the usual identification $h_{\varepsilon}(r)=h(\varepsilon r)$ and consider normal graphs of these functions over $\widetilde{\Sigma}_{\varepsilon}$ as admissible perturbations. Numbers $\tau, \beta>0$ will be specified later on.

We denote the two components of $\mathbb{R}^{N+1} \backslash \Sigma_{\varepsilon}$ by $D_{\varepsilon}^{ \pm}$, respectively. We agree that $D_{\varepsilon}^{+}$is the component containing the set $\mathcal{U}_{\varepsilon} \cap\{z>0\}$, and $D_{\varepsilon}^{-}$is "interior" to $\widetilde{\Sigma}_{\varepsilon}$. Finally, by $\chi_{D_{\varepsilon}^{ \pm}}$we denote the characteristic functions of these sets.

With this notation we set

$$
\left(X_{\varepsilon}^{*} u_{\varepsilon}\right)(y, z)=H\left(z-h_{\varepsilon}(r)\right), \quad r=\left|y^{\prime}\right|,
$$

and define the approximate solution

$$
\begin{equation*}
w_{\varepsilon}(x)=\rho_{\varepsilon}(x) u_{\varepsilon}(x)+\left(1-\rho_{\varepsilon}(x)\right)\left(\chi_{D_{\varepsilon}^{+}}(x)-\chi_{D_{\varepsilon}^{-}}(x)\right) . \tag{7.13}
\end{equation*}
$$

### 7.4 Error of the Approximation

In this section we will compute the error of the approximation, namely,

$$
S\left(w_{\varepsilon}\right)=\Delta w_{\varepsilon}+\varepsilon \partial_{x_{N+1}} w_{\varepsilon}+w_{\varepsilon}\left(1-w_{\varepsilon}^{2}\right)
$$

Using (7.13) we can write

$$
\begin{align*}
S\left(w_{\varepsilon}\right)= & \rho_{\varepsilon} S\left(u_{\varepsilon}\right)+\underbrace{w_{\varepsilon}\left(1-w_{\varepsilon}^{2}\right)-\rho_{\varepsilon} u_{\varepsilon}\left(1-u_{\varepsilon}^{2}\right)}_{\mathcal{I}}  \tag{7.14}\\
& +\underbrace{\left[\Delta, \rho_{\varepsilon}\right] u_{\varepsilon}-\left(\Delta \rho_{\varepsilon}\right)\left(\chi_{D_{\varepsilon}^{+}}-\chi_{D_{\varepsilon}^{-}}\right)+\varepsilon \partial_{x_{N+1}} \rho_{\varepsilon}\left(u_{\varepsilon}-\chi_{D_{\varepsilon}^{+}}+\chi_{D_{\varepsilon}^{-}}\right)}_{\mathcal{J}} .
\end{align*}
$$

As done in (6.5), we write $S\left(u_{\varepsilon}\right)$ in Fermi coordinates and denote $\left(X_{\varepsilon}^{*} u\right)(y, z)=$ $\widetilde{u}_{\varepsilon}(y, z)$ for brevity. Thus we get

$$
\begin{align*}
\left(X_{\varepsilon}^{*} S\left(u_{\varepsilon}\right)\right)= & \Delta_{\tilde{\Sigma}_{\varepsilon}} \tilde{u}_{\varepsilon}+\left[\partial_{z}^{2} \widetilde{u}_{\varepsilon}+f\left(\widetilde{u}_{\varepsilon}\right)\right]+\left[\varepsilon \partial_{z}\left(\pi_{\varepsilon, N+1}\right)-H_{\tilde{\Sigma}_{\varepsilon}}\right] \partial_{z} \widetilde{u}_{\varepsilon} \\
& -z \mid A_{\tilde{\Sigma}_{\varepsilon}}{ }^{2} \partial_{z} \tilde{u}_{\varepsilon}+\varepsilon \nabla_{\widetilde{\Sigma}_{\varepsilon}}\left(\pi_{\varepsilon, N+1}\right) \cdot \nabla_{\tilde{\Sigma}_{\varepsilon}} \widetilde{u}_{\varepsilon}  \tag{7.15}\\
& +\mathbb{A}_{\varepsilon}\left[\tilde{u}_{\varepsilon}\right]+\mathbb{B}_{\varepsilon}\left[\tilde{u}_{\varepsilon}\right]-z^{2} R_{\tilde{\Sigma}_{\varepsilon}} \partial_{z} \tilde{u}_{\varepsilon} .
\end{align*}
$$

To proceed we need to calculate various geometric quantities appearing in (7.15) in terms of the parametrization of $\widetilde{\Sigma}_{\varepsilon}$ given in (7.4). These are standard computations, and we will only summarize the most important points in the form of a lemma.

Lemma 7.3. Let $n_{\varepsilon}^{ \pm}$be the unit normal, $g_{\varepsilon, i j}^{ \pm}$be the coefficients of the metric, and $\mathbb{k}_{\varepsilon, j}^{ \pm}$be the principal curvatures of the ends $\Sigma_{\varepsilon}^{ \pm}$of the surface $\Sigma_{\varepsilon}$, and let $\tilde{n}_{\varepsilon}^{ \pm}$, $\widetilde{g}_{\varepsilon, i j}^{ \pm}$, and $\widetilde{\mathbb{k}}_{\varepsilon, j}^{ \pm}$be the corresponding quantities on $\widetilde{\Sigma}_{\varepsilon}^{ \pm}$, expressed in terms of the local coordinates $(r, \Theta) \in \mathbb{R}_{+} \times S^{N-1}$. Then, the following holds:

$$
\begin{align*}
\tilde{n}_{\varepsilon}(r, \Theta)= & n_{\varepsilon}(r, \Theta) \mp\left(0, \varepsilon \chi\left(\frac{r}{r_{\varepsilon}}\right) \frac{\partial_{r}^{2} W_{R}^{ \pm}(\varepsilon r) f^{ \pm}(\varepsilon r)}{1+\left|\partial_{r} W_{R}^{ \pm}(\varepsilon r)\right|^{2}}\right) \\
& +\varepsilon \chi\left(\frac{r}{r_{\varepsilon}}\right) \mathcal{O}\left(\frac{\left|f^{ \pm}(\varepsilon r)\right|}{\left(1+\varepsilon^{2} r^{2}\right)^{3 / 2}}\right) . \tag{7.16}
\end{align*}
$$

Furthermore, the matrices $g_{\varepsilon, i j}^{ \pm}$and $\widetilde{g}_{\varepsilon, i j}^{ \pm}$are diagonal, and we have the following formulas:

$$
\tilde{g}_{\varepsilon, i j}=g_{\varepsilon, i j}\left(1+\varepsilon \chi\left(\frac{r}{r_{\varepsilon}}\right) \mathcal{O}\left(\frac{\left|f^{ \pm}(\varepsilon r)\right|}{\left(1+\varepsilon^{2} r^{2}\right)^{1 / 2}}\right)\right)
$$

The principal curvatures satisfy

$$
\begin{aligned}
& \widetilde{\mathbb{k}}_{\varepsilon, 1}^{ \pm}=\mathbb{k}_{\varepsilon, 1}^{ \pm}\left(1+\varepsilon \chi\left(\frac{r}{r_{\varepsilon}}\right) \mathcal{O}\left(\frac{\left|f^{ \pm}(\varepsilon r)\right|}{\left(1+\varepsilon^{2} r^{2}\right)^{1 / 2}}\right)\right), \\
& \widetilde{\mathbb{k}}_{\varepsilon, j}^{ \pm}=\mathbb{k}_{\varepsilon, j}^{ \pm}\left(1+\varepsilon \chi\left(\frac{r}{r_{\varepsilon}}\right) \mathcal{O}\left(\frac{\left|f^{ \pm}(\varepsilon r)\right|}{\left(1+\varepsilon^{2} r^{2}\right)}\right)\right), \quad j=2, \ldots, N .
\end{aligned}
$$

Let us recall that asymptotically, as $r \rightarrow \infty$, the ends of $\Sigma_{\varepsilon}$ are parallel to $\Gamma_{\varepsilon}$. As a result, in the above formulas we can replace $g_{\varepsilon, i j}^{ \pm}$and $\mathbb{k}_{\varepsilon, j}^{ \pm}$in the right-hand side by the coefficients of the metric and principal curvatures computed on $\Gamma_{\varepsilon}$. The error created this way will be very small. Another observation we make is that if we take $r_{\varepsilon}$ as in 7.10, then we have

$$
\chi\left(\frac{r}{r_{\varepsilon}}\right) \frac{\left|f^{ \pm}(\varepsilon r)\right|}{\left(1+\varepsilon^{2} r^{2}\right)^{\beta}} \leq \frac{C \varepsilon^{2}}{\left(1+\varepsilon^{2} r^{2}\right)^{\beta^{\prime}}}
$$

for all $\beta^{\prime}<\beta$ provided that $\varepsilon$ is taken sufficiently small.

Then straightforward calculations show that the error of the initial approximation is essentially of the same size in the $\mathcal{C}_{\beta, \sigma}^{0, \mu}\left(\widetilde{\Sigma}_{\varepsilon}\right)$ sense. Namely, we can show the exact analog of Lemma 6.1 for this part of the error:

$$
\begin{equation*}
\left\|\rho_{\varepsilon} S\left(u_{\varepsilon}\right)\right\|_{\mathcal{C}_{\beta \sigma, \sigma}^{0, \mu}\left(\widetilde{\Sigma}_{\varepsilon}\right)} \leq C \varepsilon^{2-\sigma \sqrt{2}}, \quad \beta_{\sigma}=1-\sigma \sqrt{2} \tag{7.17}
\end{equation*}
$$

Now we will estimate the second term in 7.14 denoted by $\mathcal{I}$. For future purposes it is convenient to have an explicit formula:

$$
\mathcal{I}= \begin{cases}3\left(u_{\varepsilon}+1\right)^{2} \rho_{\varepsilon}\left(\rho_{\varepsilon}-1\right)+\left(u_{\varepsilon}+1\right)^{3} \rho_{\varepsilon}\left(1-\rho_{\varepsilon}^{2}\right) & \text { in } D_{\varepsilon}^{-},  \tag{7.18}\\ 3\left(u_{\varepsilon}-1\right)^{2} \rho_{\varepsilon}\left(1-\rho_{\varepsilon}\right)+\left(u_{\varepsilon}-1\right)^{3} \rho_{\varepsilon}\left(1-\rho_{\varepsilon}^{2}\right) & \text { in } D_{\varepsilon}^{-} .\end{cases}
$$

From this, using $H(t)= \pm 1+O\left(e^{-\sqrt{2}|t|}\right)$ and also the asymptotic formula (7.8), we find, with some $\sigma>0$,

$$
\begin{equation*}
\|\mathcal{I}\|_{\mathcal{C}_{\beta \sigma, \sigma}^{0, \mu}\left(\tilde{\Sigma}_{\varepsilon}\right)} \leq C \varepsilon^{2-\sigma \sqrt{2}} \tag{7.19}
\end{equation*}
$$

Our final calculation involves the third term in (7.14) denoted by $\mathcal{J}$. This term is quite important since it represents the interactions between the ends of $\tilde{\Sigma}_{\varepsilon}$. We write

$$
\mathcal{J}=\left(\Delta \rho_{\varepsilon}+\varepsilon \partial_{x_{N+1}} \rho_{\varepsilon}\right)\left(u_{\varepsilon}-\chi_{D_{\varepsilon}^{+}}+\chi_{D_{\varepsilon}^{-}}\right)+2 \nabla \rho_{\varepsilon} \cdot \nabla u_{\varepsilon}
$$

Since $H^{\prime}(t)=O\left(e^{-\sqrt{2}|t|}\right)$ we can estimate

$$
\begin{aligned}
&|\mathcal{J}| \leq C e^{-\sqrt{2}|t|} \chi_{\left\{0<\rho_{\varepsilon}<1\right\}} \\
& \leq C e^{-\sigma|z|} \exp ( -(\sqrt{2}-\sigma)\left\{\frac{C(\Sigma)}{\varepsilon}\left[1-\chi\left(\frac{r}{r_{\varepsilon}}\right)\right]\right. \\
&\left.\left.+\frac{1}{2} \chi\left(\frac{r}{r_{\varepsilon}}\right)\left|f^{+}(\varepsilon r)-f^{-}(\varepsilon r)\right|\right\}\right)
\end{aligned}
$$

By (7.8) we have

$$
\begin{equation*}
\|\mathcal{J}\|_{\mathcal{C}_{\beta \sigma, \sigma}^{0, \mu}\left(\widetilde{\Sigma}_{\varepsilon}\right)} \leq C \varepsilon^{2-\sigma \sqrt{2}} \tag{7.20}
\end{equation*}
$$

We will summarize (7.17)-7.20).
LEMMA 7.4. Let $w_{\varepsilon}$ be the approximate solution defined in 7.13). For any $\sigma \in$ $(0,1)$ the error of this approximation $S\left(w_{\varepsilon}\right)$ satisfies the following estimate:

$$
\begin{equation*}
\left\|S\left(w_{\varepsilon}\right)\right\|_{\mathcal{C}_{\beta \sigma, \sigma}^{0, \mu}\left(\widetilde{\Sigma}_{\varepsilon}\right)} \leq C \varepsilon^{2-\sigma \sqrt{2}}, \quad \beta_{\sigma}=1-\sigma \sqrt{2} \tag{7.21}
\end{equation*}
$$

Assuming that the admissible perturbation of $\widetilde{\Sigma}_{\varepsilon}$ satisfies (7.12), the constant $C$ appearing above depends on $\sigma$ but not on this perturbation.

In addition, as a function of the admissible perturbations, $S\left(w_{\varepsilon}\right)$ is a Lipschitz function from $\mathcal{C}_{\beta}^{2, \mu}\left(\widetilde{\Sigma}_{\varepsilon}\right)$ into $\mathcal{C}_{\beta \sigma, \sigma}^{0, \mu}\left(\widetilde{\Sigma}_{\varepsilon}\right)$ with a Lispchitz constant proportional to $\varepsilon^{2-\sigma \sqrt{2}}$.

### 7.5 An Outline of the Lyapunov-Schmidt Reduction

Given the results of Lemma 7.4, it is rather straightforward to implement a Lyapunov-Schmidt reduction procedure similar to the one used in the proof of Theorem 1.1. In fact, large parts are simply repetitions with some natural changes. Thus we will only give a brief outline of the general scheme. As before we look for a solution of the problem

$$
S(u)=\Delta u+\varepsilon \partial_{x_{N+1}} u+u\left(1-u^{2}\right)=0 \quad \text { in } \mathbb{R}^{N+1}
$$

in the form $u=w_{\varepsilon}+\varphi_{\varepsilon}$. Now we write

$$
\varphi_{\varepsilon}=\rho_{\varepsilon} \phi_{\varepsilon}+\psi_{\varepsilon}
$$

and decompose the original problem into a system as described in Section 4.2. As a result, we get the following analogue of (4.13)-(4.15):

$$
\begin{align*}
\Delta_{\tilde{\Sigma}_{\varepsilon}} \phi_{\varepsilon}+\partial_{z}^{2} \phi_{\varepsilon}+f^{\prime}\left(u_{\varepsilon}\right) \phi_{\varepsilon} & =\mathfrak{g}_{\varepsilon}+\mathrm{c}_{\varepsilon} H^{\prime}\left(z-h_{\varepsilon}\right) & & \text { in } \widetilde{\Sigma}_{\varepsilon} \times \mathbb{R}  \tag{7.22}\\
\left(\Delta+\varepsilon \partial_{x_{N+1}}-2\right) \psi_{\varepsilon} & =\mathfrak{h}_{\varepsilon} & & \text { in } \mathbb{R}^{N+1} \tag{7.23}
\end{align*}
$$

The functions $\mathfrak{g}_{\varepsilon}$ and $\mathfrak{h}_{\varepsilon}$ are similar to their counterparts in 4.2 , and it can be proven that they have all the properties described in 6.1. Also, all the linear theory needed is a verbatim repetition of the content of Section 5. This leads us to the existence result for the nonlinear projected problem as in Section6.2. Namely, we have a solution of the system 7.22 - 7.23 , with

$$
c_{\varepsilon}=\frac{\int_{\mathbb{R}} \mathfrak{g}_{\varepsilon} H^{\prime}\left(z-h_{\varepsilon}\right) d z}{\int_{\mathbb{R}}\left[H^{\prime}\left(z-h_{\varepsilon}\right)\right]^{2} d z}
$$

At this point all that remains to be done is to find $h_{\varepsilon}$ such that ${ }_{C_{\varepsilon}}=0$. Next we will address this problem.

### 7.6 Solution of the Reduced Problem

We note that the leading terms in the projection of $\mathfrak{g}_{\varepsilon}$ onto $H^{\prime}\left(z-h_{\varepsilon}\right)$ come from the projection of the error of the approximation $S\left(w_{\varepsilon}\right)$. To prove this requires somewhat tedious calculations that we omit. Thus we concentrate on

$$
\begin{align*}
& \int_{\mathbb{R}} S\left(w_{\varepsilon}\right) H^{\prime}\left(z-h_{\varepsilon}\right) d z=  \tag{7.24}\\
& \qquad \int_{\mathbb{R}} \rho_{\varepsilon} S\left(u_{\varepsilon}\right) H^{\prime}\left(z-h_{\varepsilon}\right) d z+\int_{\mathbb{R}}(\mathcal{I}+\mathcal{J}) H^{\prime}\left(z-h_{\varepsilon}\right) d z
\end{align*}
$$

Using (7.14) and analyzing the terms involved, we observe that

$$
\begin{align*}
\int_{\mathbb{R}} \rho_{\varepsilon} & S\left(u_{\varepsilon}\right) H^{\prime}\left(z-h_{\varepsilon}\right) d z \\
= & -c_{0} \mathcal{J}_{\tilde{\Sigma}_{\varepsilon}}\left(h_{\varepsilon}\right) \\
& +\int_{\mathbb{R}} \rho_{\varepsilon}\left[\varepsilon \partial_{z}\left(\pi_{\varepsilon, N+1}\right)-H_{\tilde{\Sigma}_{\varepsilon}}\right]\left(H^{\prime}\left(z-h_{\varepsilon}\right)\right]^{2} d z  \tag{7.25}\\
& +\Xi_{\varepsilon}\left(h_{\varepsilon}\right)
\end{align*}
$$

where $\mathcal{J}_{\widetilde{\Sigma}_{\varepsilon}}$ is essentially the Jacobi operator on $\widetilde{\Sigma}_{\varepsilon}$ :

$$
\mathcal{J}_{\widetilde{\Sigma}_{\varepsilon}}\left(h_{\varepsilon}\right)=\Delta_{\widetilde{\Sigma}_{\varepsilon}} h_{\varepsilon}+\varepsilon \nabla_{\widetilde{\Sigma}_{\varepsilon}}\left(\pi_{\varepsilon, N+1}\right) \cdot \nabla_{\widetilde{\Sigma}_{\varepsilon}} h_{\varepsilon}+\left|A_{\widetilde{\Sigma}_{\varepsilon}}\right|^{2} h_{\varepsilon}
$$

and $\Xi_{\varepsilon}\left(h_{\varepsilon}\right)$ is a small term for all admissible functions $h_{\varepsilon}$ in the sense that we have

$$
\left\|\Xi_{\varepsilon}\right\|_{C_{\beta}^{0, \mu}\left(\widetilde{\Sigma}_{\varepsilon}\right)} \leq C \varepsilon^{2+\tau} \quad \text { some } \beta>0, \tau>0
$$

It remains to calculate the second term on the right-hand side of 7.25). We observe that since $\Sigma_{\varepsilon}$ is a translating solution to the mean curvature flow, this term would have been 0 if we had not modified $\Sigma_{\varepsilon}$ to $\widetilde{\Sigma}_{\varepsilon}$. Using the fact that $\partial_{z}\left(\pi_{\varepsilon, N+1}\right)=$ $\tilde{n}_{\varepsilon, N+1}$, i.e., it is simply the $(N+1)^{\text {th }}$ component of the normal on $\widetilde{\Sigma}_{\varepsilon}$ we get, by (7.16) in Lemma 7.3 .

$$
\begin{align*}
& \int_{\mathbb{R}} \rho_{\varepsilon}\left[\varepsilon \partial_{z}\left(\pi_{\varepsilon, N+1}\right)-H_{\tilde{\Sigma}_{\varepsilon}}\right]\left[H^{\prime}\left(z-h_{\varepsilon}\right)\right]^{2} \\
&=-\varepsilon^{2} \int_{\mathbb{R}} \rho_{\varepsilon}^{ \pm} \chi\left(\frac{r}{r_{\varepsilon}}\right) \frac{\partial_{r}^{2} W_{R}^{ \pm}(\varepsilon r) f^{ \pm}(\varepsilon r)}{1+\left|\partial_{r} W_{R}^{ \pm}(\varepsilon r)\right|^{2}}\left[H^{\prime}\left(z-h_{\varepsilon}\right)\right]^{2} d z \\
&+\mathcal{O}_{\mathcal{C}_{\beta}^{0, \mu}\left(\tilde{\Sigma}_{\varepsilon}\right)}\left(\varepsilon^{2+\tau}\right)  \tag{7.26}\\
&=-\varepsilon^{2} \int_{\mathbb{R}} \rho_{\varepsilon}^{ \pm} \chi\left(\frac{r}{r_{\varepsilon}}\right) \frac{\partial_{r}^{2} F(\varepsilon r) f^{ \pm}(\varepsilon r)}{1+\left|\partial_{r} F(\varepsilon r)\right|^{2}}\left[H^{\prime}\left(z-h_{\varepsilon}\right)\right]^{2} d z+\mathcal{O}_{\mathcal{C}_{\beta}^{0, \mu}\left(\widetilde{\Sigma}_{\varepsilon}\right)}\left(\varepsilon^{2+\tau}\right) \\
&=-a_{0} \varepsilon^{2} \chi\left(\frac{r}{r_{\varepsilon}}\right) \frac{\partial_{r}^{2} F(\varepsilon r)\left[f^{+}(\varepsilon r)+f^{-}(\varepsilon r)\right]}{1+\left|\partial_{r} F(\varepsilon r)\right|^{2}}+\mathcal{O}_{\mathcal{C}_{\beta}^{0, \mu}\left(\tilde{\Sigma}_{\varepsilon}\right)}\left(\varepsilon^{2+\tau}\right)
\end{align*}
$$

where $a_{0}>0$ is a constant and

$$
\rho_{\varepsilon}^{ \pm}(y, z)= \begin{cases}\rho_{\varepsilon}(y, z), & y \in \widetilde{\Sigma}_{\varepsilon}^{ \pm} \\ 0 & \text { otherwise }\end{cases}
$$

In (7.26) we have omitted terms that are at most of a size comparable to $\varepsilon^{2+\tau}$ in the sense of $\mathcal{C}_{\beta}^{0, \mu}\left(\widetilde{\Sigma}_{\varepsilon}\right)$, as indicated by the notation. We observe as well that the error in replacing $\partial_{r}^{2} W_{R}^{ \pm}(\varepsilon r)$ and $\partial_{r} W_{R}^{ \pm}(\varepsilon r)$ by $\partial_{r}^{2} F(\varepsilon r)$ and $\partial_{r} F(\varepsilon r)$, respectively, again results in a higher-order term. This justifies the third line in 7.26.

Going back to 7.24 we observe that the projection on $\mathcal{I}$ is again negligible since, by (7.18), we see that $\mathcal{I} \approx e^{-2 \sqrt{2}|z|} \chi_{\left\{0<\rho_{\varepsilon}<1\right\}}$. Thus it remains to calculate

$$
\begin{align*}
& \int_{\mathbb{R}} \mathcal{J} H^{\prime}\left(z-h_{\varepsilon}\right) d z=  \tag{7.27}\\
& \quad \int_{\mathbb{R}}\left[\left(\Delta \rho_{\varepsilon}+\varepsilon \partial_{x_{N+1}} \rho_{\varepsilon}\right)\left(u_{\varepsilon}-\chi_{D_{\varepsilon}^{+}}+\chi_{D_{\varepsilon}^{-}}\right)+2 \nabla \rho_{\varepsilon} \cdot \nabla u_{\varepsilon}\right] H^{\prime}\left(z-h_{\varepsilon}\right) d z
\end{align*}
$$

Using the definition of $\rho_{\varepsilon}$ in (7.11) and the identity $1-H^{2}=\sqrt{2} H^{\prime}$, after some integrations by parts we get

$$
\begin{array}{rl}
\int_{\mathbb{R}} & \mathcal{J} H^{\prime}\left(z-h_{\varepsilon}\right) d z \\
= & \int_{\mathbb{R}}\left[\rho_{\varepsilon}^{\prime \prime}\left(H^{\prime}\left(z-h_{\varepsilon}\right)-\chi_{D_{\varepsilon}^{+}}+\chi_{D_{\varepsilon}^{-}}\right)+2 \rho_{\varepsilon}^{\prime} H^{\prime}\left(z-h_{\varepsilon}\right)\right] H^{\prime}\left(z-h_{\varepsilon}\right) d z \\
& +\mathcal{O}_{\mathcal{C}_{\beta}^{0, \mu}\left(\widetilde{\Sigma}_{\varepsilon}\right)}\left(\varepsilon^{2+\tau}\right) \\
= & 2 \int_{\mathbb{R}} \rho_{\varepsilon}^{\prime}\left[H^{\prime}\left(z-h_{\varepsilon}\right)\right]^{2} d z+\mathcal{O}_{\mathcal{C}_{\beta}^{0, \mu}\left(\widetilde{\Sigma}_{\varepsilon}\right)}\left(\varepsilon^{2+\tau}\right)  \tag{7.28}\\
= & a_{1} \exp \left\{-2 \sqrt{2} \frac{C(\Sigma)\left[1-\chi\left(\frac{r}{r_{\varepsilon}}\right)\right]}{\varepsilon}\right\} \\
& \quad \times \exp \left\{-\sqrt{2} \chi\left(\frac{r}{r_{\varepsilon}}\right)\left(f^{+}(\varepsilon r)+f^{-}(\varepsilon r)\right)\right\} e^{-2 \sqrt{2} h_{\varepsilon}}+\mathcal{O}_{\mathcal{C}_{\beta}^{0, \mu}\left(\tilde{\Sigma}_{\varepsilon}\right)}\left(\varepsilon^{2+\tau}\right)
\end{array}
$$

Summarizing (7.25, (7.26), and (7.28) we get that the reduced problem amounts to solving for $h_{\varepsilon}$ the following equation:

$$
\begin{align*}
& c_{0} \mathcal{J}_{\widetilde{\Sigma}_{\varepsilon}}\left(h_{\varepsilon}\right)+\widetilde{c}_{1} \exp \left\{-2 \sqrt{2} \frac{C(\Sigma)\left[1-\chi\left(\frac{r}{r_{\varepsilon}}\right)\right]}{\varepsilon}\right\} \\
& \times \exp \left\{-\sqrt{2} \chi\left(\frac{r}{r_{\varepsilon}}\right)\left(f^{+}(\varepsilon r)+f^{-}(\varepsilon r)\right)\right\} h_{\varepsilon}= \\
&= a_{0} \varepsilon^{2} \chi\left(\frac{r}{r_{\varepsilon}}\right) \frac{\partial_{r}^{2} F(\varepsilon r)\left[f^{+}(\varepsilon r)+f^{-}(\varepsilon r)\right]}{1+\left|\partial_{r} F(\varepsilon r)\right|^{2}}  \tag{7.29}\\
&-a_{1} \exp \left\{-2 \sqrt{2} \frac{C(\Sigma)\left[1-\chi\left(\frac{r}{r_{\varepsilon}}\right)\right]}{\varepsilon}\right\} \\
& \times \exp \left\{-\sqrt{2} \chi\left(\frac{r}{r_{\varepsilon}}\right)\left(f^{+}(\varepsilon r)+f^{-}(\varepsilon r)\right)\right\} \\
&+\mathcal{O}_{\mathcal{C}_{\beta}^{0, \mu}\left(\widetilde{\Sigma}_{\varepsilon}\right)}\left(\varepsilon^{2+\tau}\right) .
\end{align*}
$$

This is of course a fixed point problem for $h_{\varepsilon}$ and the term that we have denoted by $\mathcal{O}_{\mathcal{C}_{\beta}^{0, \mu}\left(\widetilde{\Sigma}_{\varepsilon}\right)}\left(\varepsilon^{2+\tau}\right)$ depends in a nonlinear and nonlocal way on $h_{\varepsilon}$. It can be shown
that this term in fact is a Lipschitz contraction of $h_{\varepsilon}$ (and consequently of the admissible functions $h_{\varepsilon}$ ). This is quite similar to the previous part. We concentrate on analyzing the invertibility of the linear operator on the right-hand side of (7.29).

We make first an observation: by the choice of $f^{ \pm}$we have for $r>2 r_{\varepsilon}$

$$
\begin{equation*}
a_{0} \varepsilon^{2} \frac{\partial_{r}^{2} F(\varepsilon r)\left[f^{+}(\varepsilon r)+f^{-}(\varepsilon r)\right]}{1+\left|\partial_{r} F(\varepsilon r)\right|^{2}}-a_{1} \exp \left\{-\sqrt{2}\left(f^{+}(\varepsilon r)+f^{-}(\varepsilon r)\right)\right\}=0 . \tag{7.30}
\end{equation*}
$$

Second, when $r \leq 2 r_{\varepsilon}$, then by the choice of $r_{\varepsilon}$ we have that the whole right-hand side is an $\mathcal{O}_{\mathcal{C}_{\beta}^{0, \mu}\left(\tilde{\Sigma}_{\varepsilon}\right)}\left(\varepsilon^{2+\tau}\right)$ term. As a consequence, arranging some terms suitably, we are left with solving the following problem:

$$
-c_{0} \mathcal{J}_{\widetilde{\Sigma}_{\varepsilon}}\left(h_{\varepsilon}\right)+\chi\left(\frac{r}{r_{\varepsilon}}\right) \exp \left\{-\sqrt{2}\left(f^{+}(\varepsilon r)+f^{-}(\varepsilon r)\right)\right\} h_{\varepsilon}=\mathcal{O}_{\mathcal{C}_{\beta}^{0, \mu}\left(\widetilde{\Sigma}_{\varepsilon}\right)}\left(\varepsilon^{2+\tau}\right) .
$$

Scaling back to the surface $\tilde{\Sigma}$, we are left with the problem of the form

$$
\begin{align*}
& \Delta_{\tilde{\Sigma}} h+\nabla_{\tilde{\Sigma}}\left(\pi_{N+1}\right) \cdot \nabla_{\tilde{\Sigma}} h+\left|A_{\tilde{\Sigma}}\right|^{2} h \\
& \quad+\frac{1}{\varepsilon^{2}} \chi\left(\frac{r}{\varepsilon r_{\varepsilon}}\right) \exp \left\{-\sqrt{2}\left(f^{+}(r)+f^{-}(r)\right)\right\} h  \tag{7.31}\\
& \quad=\mathcal{O}_{\mathcal{C}_{\beta}^{0, \mu}(\tilde{\Sigma})}\left(\varepsilon^{\tau}\right) .
\end{align*}
$$

Since we consider only the radial perturbations of the original surface $\Sigma$ as admissible, then $\tilde{\Sigma}$ is also rotationally symmetric, and the above problem reduces to an ODE. Thus we may use a similar technique as in the previous part, namely solve it by the variation-of-parameters formula, gluing various parts. When $r<\varepsilon r_{\varepsilon}$ our operator is essentially identical with the linearization of the translating graph solution to the mean curvature flow (cf. Lemma 7.3). Inverting this operator is the only significantly different part of the theory, and thus we will present it in some detail. Note that when $r>\varepsilon r_{\varepsilon}$ the operator above resembles the linearized operator $\mathfrak{L}_{\delta}$, treated extensively in Section 3.6. An argument similar to the one in Section 3.6 can be used to control a fundamental set and to write the variation-of-parameters formula.

### 7.7 Jacobi Operator of the Traveling Catenoid $\boldsymbol{\Sigma}$

Our goal is to prove the following:
Lemma 7.5. Let $g \in \mathcal{C}_{\beta}^{0, \mu}(\widetilde{\Sigma}), \beta>1$, be a function depending on the radial variable only. There exists a solution $v=v(r)$ of the problem

$$
\begin{aligned}
& {\left[\Delta_{\tilde{\Sigma}}+\nabla_{\tilde{\Sigma}}\left(\pi_{N+1}\right) \cdot \nabla_{\tilde{\Sigma}}+\left|A_{\tilde{\Sigma}}\right|^{2}\right] v} \\
& \quad+\frac{1}{\varepsilon^{2}} \chi\left(\frac{r}{\varepsilon r_{\varepsilon}}\right) \exp \left\{-\sqrt{2}\left(f^{+}(r)+f^{-}(r)\right)\right\} v=g,
\end{aligned}
$$

with

$$
\|v\|_{\mathcal{C}_{\beta-1}^{2}, \mu}(\tilde{\Sigma}) \leq C\|g\|_{\mathcal{C}_{\beta}^{0, \mu}(\tilde{\Sigma})} .
$$

In the region where $r<\varepsilon r_{\varepsilon}$ the surface $\tilde{\Sigma}$ coincides with the original traveling catenoid $\Sigma$. This is where our problem is different and we need the following result:

LEMMA 7.6. Let us consider the following problem:

$$
\begin{equation*}
\mathcal{J}_{\Sigma}(v)=\left[\Delta_{\Sigma}+\nabla_{\Sigma}\left(\pi_{N+1}\right) \cdot \nabla_{\Sigma}+\left|A_{\Sigma}\right|^{2}\right] v=g \tag{7.32}
\end{equation*}
$$

where $g \in \mathcal{C}_{1+\beta}^{0, \mu}(\Sigma)$, with $\beta>0$, is a function that depends on the radial variable $r$ only. There exists a solution $v=v(r)$ of this problem such that

$$
\begin{equation*}
\|v\|_{\mathcal{C}_{\beta}^{0, \mu}(\Sigma)} \leq C\|g\|_{\mathcal{C}_{1+\beta}^{0, \mu}(\Sigma)} \tag{7.33}
\end{equation*}
$$

Proof. We observe that the Jacobi operator for the surface $\Sigma$ can no longer be expressed in terms of the radial variable globally. In fact, we need to use three charts on $\Sigma$ to write conveniently equation (7.31) in local variables. This in fact is the only new element.

Near the point of the traveling catenoid where $r=R$ we will express the surface as a graph over the $x_{N+1}$-axis. Thus we have, following the results in [5]:

$$
\Sigma \cap B_{R_{1}}=\left\{(q(z) \Theta, z) \mid z \in\left(-z_{0}, z_{0}\right)\right\}
$$

where $R_{1}>R$ and $q$ satisfies

$$
\left(\frac{N-1}{q}-q^{\prime}\right)\left(1+\left(q^{\prime}\right)^{2}\right)=q^{\prime \prime}
$$

With this in mind we express the radial function $g$ on the right-hand side of 7.32 ) in terms of $z=q^{-1}(r)$. We will abuse notation and denote this, and other functions involved, by the same symbols $g$, $v$, etc.

We write the Jacobi operator $\mathcal{J}_{\Sigma}$ restricted to functions of $v=v(z)$ in this chart and get the following ODE:

$$
\begin{align*}
& \frac{v^{\prime \prime}}{1+\left|q^{\prime}\right|^{2}}+\left(\frac{N-1}{q}+\frac{1}{1+\left|q^{\prime}\right|^{2}}\right) v^{\prime}  \tag{7.34}\\
&+\left[\frac{\left|q^{\prime}\right|^{2}}{\left(1+\left|q^{\prime}\right|^{2}\right)^{3}}+\frac{N-1}{q^{2}\left(1+\left|q^{\prime}\right|^{2}\right)}\right] v=g
\end{align*}
$$

We multiply this equation through by $1+\left|q^{\prime}\right|^{2}$ and arrive at the equation in the following form:

$$
v^{\prime \prime}+p_{1}(z) v^{\prime}+p_{2}(z) v=\left(1+\left|q^{\prime}\right|^{2}\right) g=\tilde{g}
$$

Let $\phi_{0}$ and $\phi_{1}$ be two linearly independent elements of a fundamental set of the operator chosen so that

$$
\begin{array}{ll}
\phi_{0}(0)=0, & \phi_{1}(0)=1 \\
\phi_{0}^{\prime}(0)=1, & \phi_{1}^{\prime}(0)=0
\end{array}
$$

Finally, let $P_{1}(z)$ be a primitive of $p_{1}$. Then we can write explicitly

$$
\begin{equation*}
v(z)=\int_{-z_{0}}^{z} \frac{e^{-P_{1}(\zeta)}}{\phi_{0}^{2}(\zeta)} \int_{-z_{0}}^{\zeta} \widetilde{g}\left(\zeta^{\prime}\right) \phi_{0}\left(\zeta^{\prime}\right) e^{P_{1}\left(\zeta^{\prime}\right)} d \zeta^{\prime}+a_{0} \phi_{0}(z)+a_{1} \phi(z) \tag{7.35}
\end{equation*}
$$

Next we write $\mathcal{J}_{\Sigma}$ on the ends $\Sigma^{ \pm} \backslash B_{r_{0}}$, where $r_{0}$ is chosen so that $R<r_{0}<$ $R_{1}$ and the various local charts overlap. The natural parametrization is of course

$$
\Sigma^{ \pm} \backslash B_{r_{0}}=\left\{\left(r \Theta, W_{R}^{ \pm}(r)\right) \mid(\Theta, r) \in S^{N-1} \times \mathbb{R}_{+}\right\}
$$

In this chart $\mathcal{J}_{\Sigma}$ can be written as an ODE in $r$ for each of the two ends. This is very similar to what we did in Lemma 3.4. Denoting by $\phi_{0}^{ \pm}$and $\phi_{1}^{ \pm}$the elements of a fundamental set corresponding to $\phi_{0}$ and $\phi_{1}$ in Lemma 3.4, and letting $\widetilde{g}^{ \pm}=$ $\left(1+\left|\partial_{r} W_{R}^{ \pm}\right|^{2}\right)$, we get the following formula:

$$
\begin{align*}
v^{ \pm}(r)= & -\phi_{0}^{ \pm} \int_{r_{0}}^{r} \frac{\left.\phi_{1}^{ \pm}(\rho)\right) \tilde{g}^{ \pm}(\rho)}{W^{ \pm}(\rho)} d \rho \\
& +\phi_{1}^{ \pm} \int_{r_{0}}^{r} \frac{\left.\phi_{0}^{ \pm}(\rho)\right) \widetilde{g}^{ \pm}(\rho)}{W^{ \pm}(\rho)} d \rho+a_{1}^{ \pm} \phi_{1}^{ \pm}(r) \tag{7.36}
\end{align*}
$$

for a general solution $v$ in $\mathcal{C}_{\beta}^{2, \mu}(\Sigma)$. Note that we have

$$
\phi_{0}^{ \pm}(r) \sim \frac{1}{1+r}, \quad \phi_{1}^{ \pm}(r) \sim r e^{-r^{2}}, \quad r \gg 1,
$$

which is why in (7.36 we have included only constant multiplicities of $\phi_{1}^{ \pm}$.
Next we need to choose the four constants $a_{0}, a_{1}$, and $a_{1}^{ \pm}$in such a way that

$$
\begin{aligned}
v^{ \pm}\left(r_{0}\right) & =v \circ\left(W_{R}^{ \pm}\right)^{-1}\left(r_{0}\right), \\
\partial_{r} v^{ \pm}\left(r_{0}\right) & =\partial_{r} v \circ\left(W_{R}^{ \pm}\right)^{-1}\left(r_{0}\right) .
\end{aligned}
$$

This is a matter of solving a simple system of four linear equations.
After this is done we have a solution defined now on the whole surface $\Sigma$. Estimate (7.33) follows directly from the explicit formulas we have derived. This ends the proof.

Next, we describe how to solve the linearized problem (7.31). Note that as long as $r<\varepsilon r_{\varepsilon}$ we are dealing with the Jacobi operator discussed in the lemma above. Thus, at least up to $r=\varepsilon r_{\varepsilon}$, we will have no problem in defining a solution $v$ in $\mathcal{C}_{\beta-1}^{2, \mu}\left(\Sigma \cap\left\{r<r_{\varepsilon}\right\}\right)$ (here we take $\left.\beta>1\right)$. What is left is to solve a problem of the form

$$
\begin{equation*}
\mathcal{J}_{\tilde{\Sigma}^{ \pm}}\left(v^{ \pm}\right)+\frac{1}{\varepsilon^{2}} \chi\left(\frac{r}{\varepsilon r_{\varepsilon}}\right) \exp \left\{-\sqrt{2}\left(f^{+}(r)+f^{-}(r)\right)\right\} v^{ \pm}=g^{ \pm} \tag{7.37}
\end{equation*}
$$

on each end $\Sigma^{ \pm}$, with $r>r_{\varepsilon}$ for radial functions $v^{ \pm}$, with initial data given by the solution $v$, already found, at $r=\varepsilon r_{\varepsilon}$.

Now we notice that because of the definition on $f^{ \pm}$in (7.6)-(7.7), the operator appearing in (7.37) is very similar to the operator $\mathfrak{L}_{\delta}$ considered in Section 3.5. In fact, we can write (7.37) in the form

$$
\frac{(1+o(1)) v_{r r}^{ \pm}}{1+\left|\partial_{r} W_{R}^{ \pm}\right|^{2}}+\frac{(N-1)(1+o(1)) v_{r}^{ \pm}}{r}+p_{\varepsilon}^{ \pm}(r) v^{ \pm}=g^{ \pm}, \quad r>\varepsilon r_{\varepsilon}
$$

where

$$
p_{\varepsilon}(r) \sim \frac{1}{1+r^{2}} \log \left(\frac{1+r^{2}}{\varepsilon^{2}}\right), \quad r>\varepsilon r_{\varepsilon}
$$

which is in agreement with the behavior of the function $p_{\delta}$ in 3.37), and the $o(1)$ term above means terms that are small both in $\varepsilon$ and $r$. Since we are interested in this problem only for large values of $r \geq \varepsilon r_{\varepsilon} \sim e^{\widetilde{c} / \varepsilon}$, we see that the argument in Section 3.6 can be repeated verbatim to solve our problem. Having the inverse of the operator in 7.31 at hand, we proceed in the same way as in the previous case to finally solve a fixed point problem for $h$. We omit the details.

Acknowledgment. The research of Manuel del Pino was supported in part by Fondecyt Grant 1110181, Proyecto Anillo ACT 125 CAPDE, and Fondo BasalCMM, Chile. Michał Kowalczyk was partly supported by Fondecyt Grant 1090103, Proyecto Anillo ACT 125 CAPDE, and Fondo Basal-CMM, Chile. Juncheng Wei was partly supported by General Research Fund from RGC of Hong Kong.

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